

Aiding Basic Space Science in Developing Nations

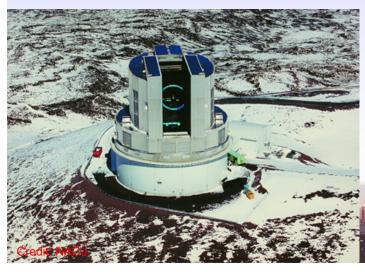
The Official Development Assistance (ODA) programme of Japan

Kaz Sekiguchi
National Astronomical Observatory of Japan

Scientific and Technical Subcommittee 47th session, COPUOS Vienna, 2010/02/8-19

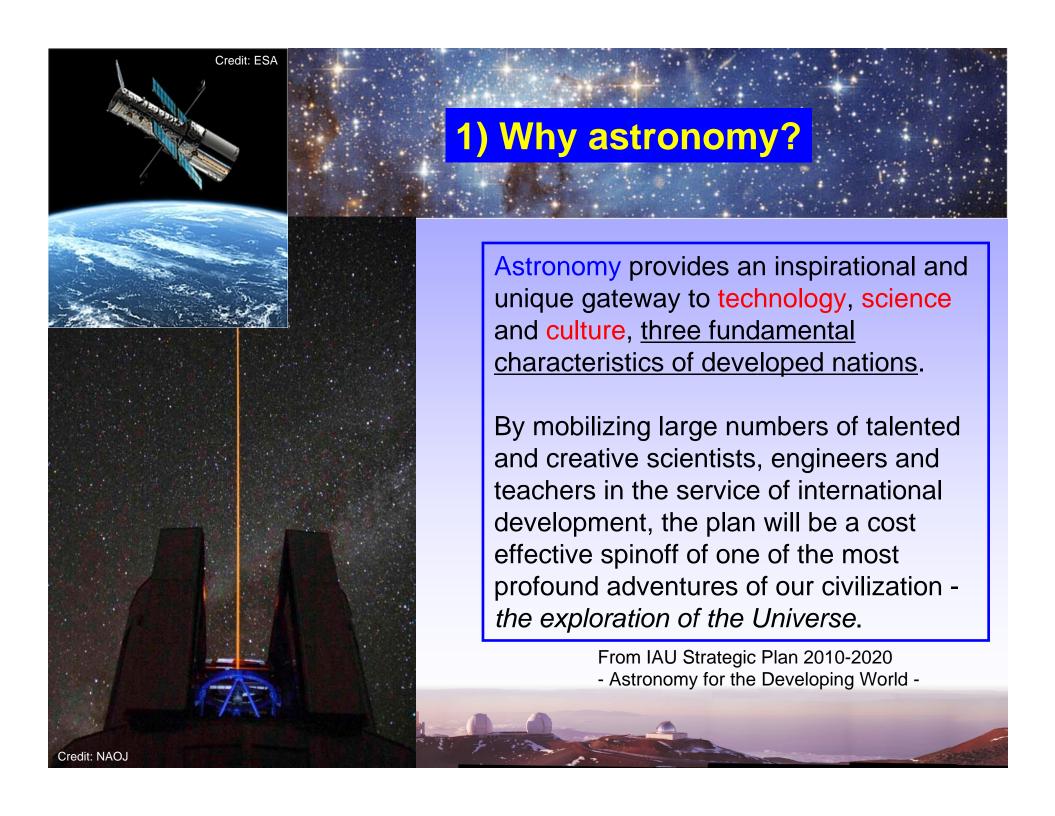


- 1) Why astronomy?
- 2) TRIPOD concept
- 3) Japan's Official Development Assistance (ODA)
- 4) Over 25 years of commitments









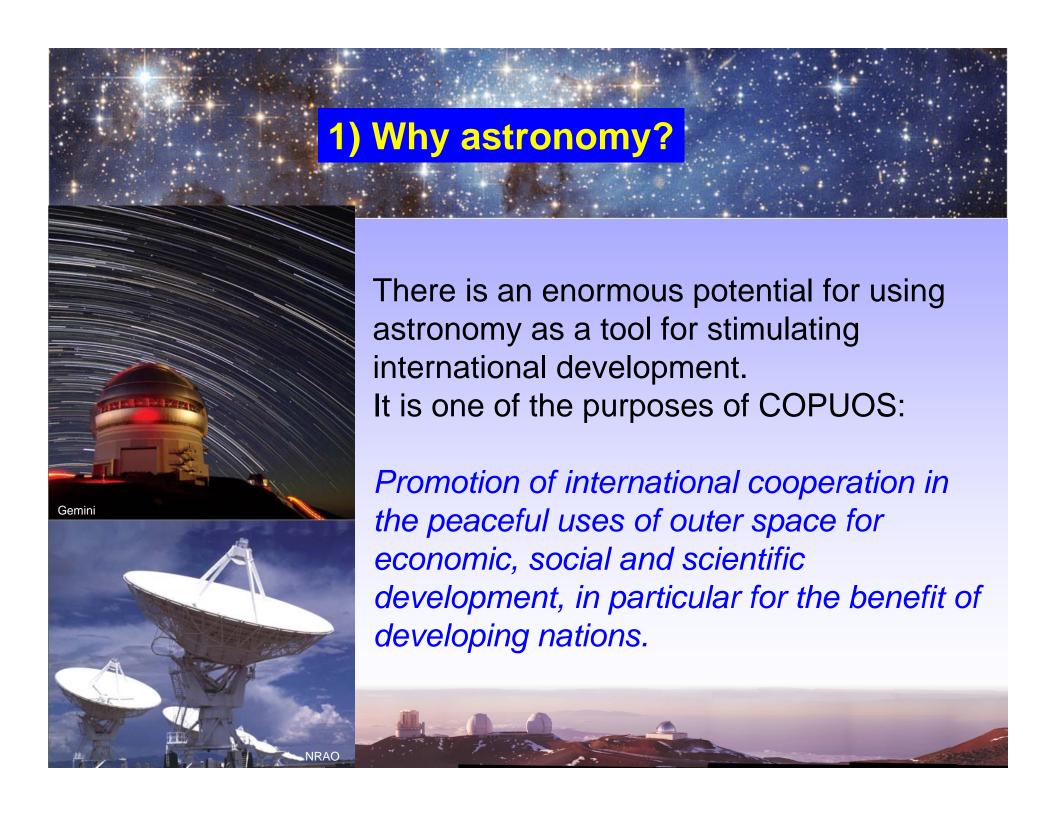


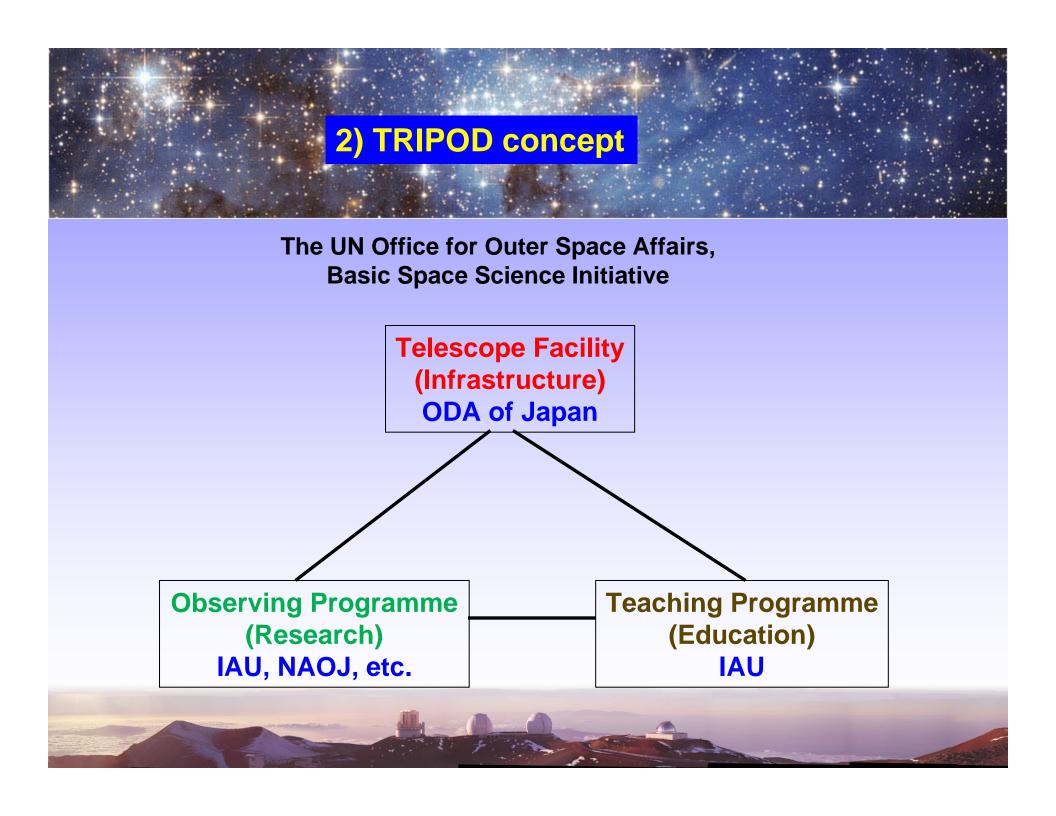
Astronomy combines science and technology with inspiration and excitement, it can play a unique role in facilitating <u>education</u> and <u>capacity building</u> and in furthering <u>sustainable development</u> throughout the world.

• It's an exciting gateway into physics, chemistry, biology and mathematics.

 The need to study the faintest celestial objects has driven advanced developments in electronics, optics and information technology.

 It inspires teenagers to choose a career in science and technology





2) TRIPOD concept

Astronomical telescope facilities were supplied by the Official Development Assistance (ODA) of Japan.



NATIONAL UNIVERSITY OF ASUNCION REPUBLIC OF PARAGUAY



45-cm telescope

2) TRIPOD concept

Research programmes were coordinated and supported by the National Astronomical Observatory of Japan and the IAU.



Sri Lanka astronomer adjusting CCD spectrograph for the reflector.

Spectroscopic Observations of Novae V1065 CENTAURI and V1280 SCORPII using 45cm Cassegrain Telescope at Arthur C Clarke Institute, Sri Lanka

> J. Adassuriya, S. Gunasekera, and I. Medagangoda Arthur C Clarke Inst. for Modern Tech., Sri Lanka

The spectroscopic observations of two novae namely nova Centauri 2007 (V1065 CEN) and nova Scorpii 2007 (V1280 SCO) were made by 45 cm Cassegrain telescope with f/12 at Arthur C Clarke Institute, Sri Lanka during the period at 31st January to 20th Feb 2007. High resolution ($\alpha\Delta\lambda$ =22000) profiles in H α (6563 oA) region were obtain for V1065 CEN, 6, 15 and 20 days after maximum and H α profiles of the same resolution were obtained for V1280 SCO, 4 days after maximum, covering the early decline stages of novae.

V1065 CEN is He/N-type spectra which characterize a broad (Gaussian FWHM 49 oA), saddle shaped and asymmetric Hα emission line with out prominent P-Cyg absorption component. Completely different Hα profile of V1280 SCO shows prominent P-Cyg absorption and narrow emission line (Gaussian FWHM 26 oA) which can be classified as Fe II type nova. The absence of prominent P-Cyg structure in V1065 CEN suggests that the emission causes by discrete shell while the prominent P-Cyg structure in V1280 SCO evidences a wind-like structure. The expansion velocities of these two systems measured from the minima of the P-Cyg profiles are close to 2300 km/s for V1065 CEN, 6 days after the maximum and 716 km/s for V1280 SCO, 4 days after the maximum

The light curves V-t, B-t and visual-t have been used to estimate the distances of both novae. Based on the photometric analysis, the Nova V1065 CEN can be classified as fast (11<32<25) nova with the parameters t2V 21 days, t3V=328 days and t2B 23 days, t3B=31 days. The derived absolute magnitudes at maximum for nova V1065 CEN to be Mo,V = -7.58-0.18 and Mo,B= -7.75±0.25. The mean distance module 16.57 and the color excess EB-V = \pm 0.6 correspond to a distance 3.51±0.33 kpc. The parameters t2V=12 days and t3V=14 days were calculated from visual-t light curve for nova V1280 SCO and It can be determine that the nova is in between very fast and fast nova. The rate of decline at C, 0.48 mag/d (very fast=0.2 mag/d) clearly determine that V1280 SCO is classified as very fast nova. The mean absolute magnitude at maximum is calculated to be Mo,V=-8.7±0.1. Neglecting the interstellar reddening the estimated distance to the nova V1280 SCO is 3.240.2 kpc.

A research paper presented at the 2009 UN/NASA/ESA/JAXA Workshop on BSS & IHY 2007 in Daejeon, Korea, (Sept. 21-25, 2009)



Teaching material and hands-on astrophysics material has been developed for the operation of such astronomical telescope facilities in an university environment.





2) TRIPOD concept

A similar TRIPOD concept is being developed for the planetarium facilities.

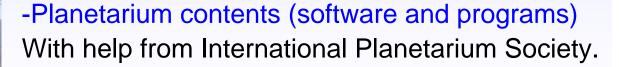


-Planetarium facility

Supplied by the ODA of Japan

-Operation support

Follow up assistance programmes by Japan International Cooperation Agency (JICA) and



3) Official Development Assistance (ODA) programme

Official Development Assistance (ODA) programme

(http://www.mofa.go.jp/policy/oda/)

Started in 1954

Objectives:

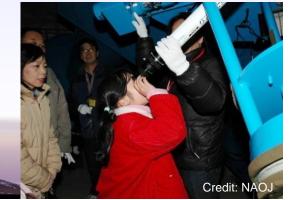
- Contribute to the peace and development of the international community
- Help ensure Japan's own security and prosperity.

Major parts of the ODA support goes to:

- economic and social infrastructure development,

- human resource development,

- institution building.



3) Official Development Assistance (ODA) programme

In 1975, Cultural Grant Aassistance (CGA) programme (http://www.mofa.go.jp/policy/oda/category/cultural/index.html) was introduced.



The Cultural Grant Assistance (CGA) provides to contribute to the promotion of cultural and higher educational activities and preservation of cultural heritage in developing countries.





4) Over 25 years of commitments by Japanese Government

Since 1982, the Japanese Government has donated 27 units of astronomical equipment to 22 nations.



National Astronomical Observatory of Tarija, Bolivia

7 of the items donated were professional-grade reflecting telescopes with scientific instruments, such as CCD cameras and/or the spectrographs, which can be used for photometric and spectroscopic observations.

In addition to these, 20 planetarium systems have been installed at universities and space-education museums.



Table 1: Astronomical equipment (reflective telescopes and accessories) donated by Japan

Receiving Institutions	Country	Year	Tel. size(φ)
- Science Centre	Singapore	1987	40cm
- Bosscha Observatory,			
Bandung Institute of Technology	Indonesia	1988	45cm
- Chulalongkorn University, Bangkok	Thailand	1989	45cm
-Arthur C. Clarke Institute for Modern			
Technologies	Sri Lanka	1995	45cm
- Asuncion National University	Paraguay	1999	45cm
-Philippine Atmospheric, Geophysical and			
Astronomical Services Administration	The Philippines	2000	45cm
-Cerro Calan Astronomical Observatory,			
Univ. of Chile	* Chile	2001	45cm

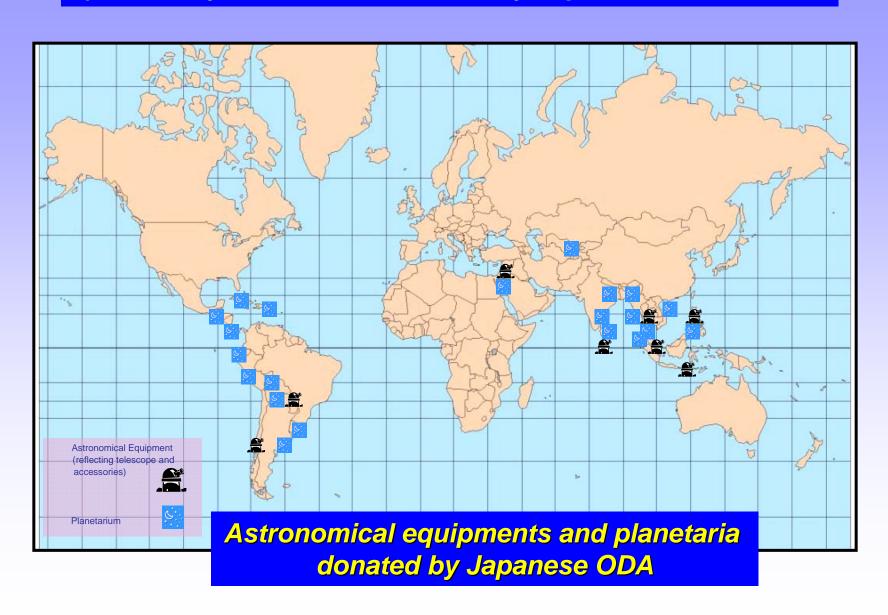


4) Over 25 years of commitments by Japanese Government

20 Planetarium equipment

Table 2: Planetarium equipment donated by Japan			
Planetarium/location		Country	Year
Pagoda Cultural Centre, Ynagon	•	Myanmar	1986
Haya Cultural Centre, Amman		Jordan	1989
Space Science Education Centre, Kuala Lumpur	(• <u> </u>	Malaysia	1989
Auxiliary projectors for the already existing planetarium, Manila		The Philippines	1990
Meghnand Saha Planetarium, University of Burdwan Golapbag	•	India	1993
Planetario de la Ciudad, Buenos Aires	•	Argentina	1993
Planetario de la Ciudad, Montevideo	*	Uruguay	1994
Ho-Chi Minh Memorial Culture Hall, Vinh City Planetarium	*	Viet Nam	1998
Auxiliary projectors for the already-exsisting planetarium, Bangkok		Thailand	1998
Auxiliary projectors for the already-exsisting planetarium, Colombo		Sri Lanka	1998
Anna Science Centre Planetarium, Tamilnadu Science and Technology Centre		India	1999
City Park, Tashkent	C::::		2000
Asuncion National University	6	Paraguay	2001
Planetario Municipal, Cuenca	¥		2002
Children Museum, San Pedro Sula	1+1		2002
National Costa Rica University	9	Costa Rica	2003
Laboratorio Central del Instituto Geofisico, Lima	&	Peru	2004
National Astronomical Observatory of Tarija		Bolivia	2007
Tin Marin Children's Museum, San Salvador	ĕ	El Salvador	2007
National History Museum, Habana		Cuba	2009

4) Over 25 years of commitments by Japanese Government







Photometric and Spectroscopic Studies of BW Eri

Desima Kristyowati, Hakim L. Malasan, and Hanindyo Kuncarayakti

Abstract New CCD photometric and spectroscopic studies of eclipsing binary BW Eridani are presented. BVRI photometric observations were carried out using Bosscha's 20-cm (f/10) GAO-ITB Remote Telescope System in 2006 and 28-cm (f/10) Schmidt–Cassegrain telescope in 2007. Low-resolution spectra (R = 400 ~ 500) were obtained using Bosscha's 45-cm (f/12) GOTO telescope equipped with Bosscha Compact Spectrograph (Malasan et al. 2001) in optical window. The investigation of $B,\,V,\,R,\,I$ light curves by fitting method yields in temperature 7,480 \pm 2,950 K and 5,200 \pm 875 K, fractional radii 0.491 \pm 0.126 and 0.280 \pm 0.135, for the primary and secondary components, respectively. An inclination 89° \pm 2.2° is also deduced. We obtained the time for primary eclipse at HJD = 2453769.1760 \pm 0.0118 by Kwee–van Woerden method, which indicate period change. At the orbital phase of 0.955 and 0.511 the star's spectrum is consistent with spectral type G8V for the secondary and A7V for the primary, respectively.

Keywords Eclipsing binary · Photometry · Spectroscopy

1 Introduction

The eclipsing binary system BW Eri (SAO 169130; period 0.6384777 days) has been known as a candidate for a system caught in an out-of-contact state. The previous observations by Baade in 1976 and Duerbeck in 1977 were carried out with *UBV* photometers in European Southern Observatory, gave the evidence of period variation. The complete photometric study was first announced by Baade et al. (1982), who leads some conclusion i.e. BW Eri is a semi-detached system, with the primary filling its critical volume, component with very different surface temperatures, and a period variation that seems to indicate a mass transfer.

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A research paper published in Proceedings of the Third UN/NASA/ESA/NASA Workshop on BSS & IHY 2007, ASSP, p149 (2010)

D. Kristyowati (🗷), H.L. Malasan, and H. Kuncarayakti Bosscha Observatory, Institut Teknologi Bandung, Lembang, Bandung 40139, Indonesia e-mail: mydesima@gmail.com

H.J. Haubold and A.M. Mathai (eds.), Proceedings of the Third UN/ESA/NASA Workshop on the International Heliophysical Year 2007 and Basic Space Science, Astrophysics and Space Science Proceedings, DOI 10.1007/978-3-642-03325-4_13, © Springer-Verlag Berlin Heidelberg 2010





Example of CCD photometry using the 45-cm telescope in Paraguay

Fredy Doncel, Arexis Toroche and Takeshi Noguchi Universidad Nacional de Asuncion Facultad Politecnica Observatorio Astronomico

Abstract

A SX Phe-type pulsating variable KZ Hya (HD94033) was observed with CCD set attached to the 45-cm reflector at Asuncion Astronomical Observatory in Paraguay. In the present work, 7 maximum phases were covered. A new ephemeris has been obtained, and the result suggests a probable change of the pulsation period of KZ Hya.

1. Introduction

CCD photometric observations of KZ Hya (α =10^h51^m54.1^s, δ =-25^d21^m11^s, 2000) were made during 4 nights from April 18 to May 9, 2002, with the 45-cm reflector (made by Goto) at Asuncion Astronomical Observatory (Longitude=57. d012W, Latitude=-25. d08, h=25m) in Paraguay. A photograph of the 45-cm telescope is shown in figure 1, and the observatory building with sliding roof is shown in figure 2.



Figure 1. The 45-cm telescope at Asuncion Observatory.



Figure 2. The Observatory building with a sliding roof.

2. Observation and reduction

Observations of KZ Hya were carried out using CCD camera with BVRI color filters, attached to the Cassegrain focus the 45-cm telescope. The present ST-8 type CCD camera has 1530*1020 pixels with field of view of about 8.7*5. system was made by SBIG (Santa Barbara Instrument Group).

Figure 3 shows the result of CCD photometry during the night of May 8, 2002 as an example. Exposure times of the used color-bands *BGRI* were 30 second (*B*-band), 10 second (*G*-band), 10 second (*G*-band), 10 second (*R*-band),

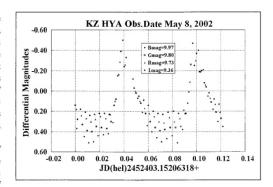


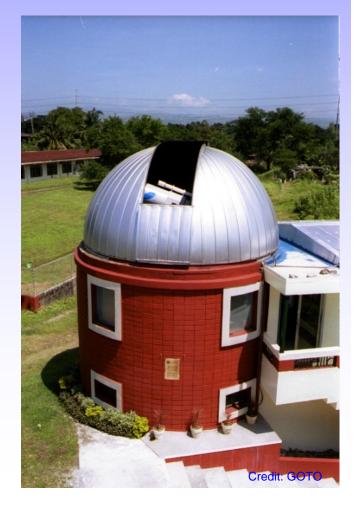
Figure 3. The observation of CCD photometry of the KZ Hya with BGRI colors which were made during the night of May 8, 2002

A research paper presented at the 2009 UN/ESA/NASA Workshop on BSS & IHY 2007 in Tokyo, Japan, (June 18-22, 2007)





PHILIPPINE ATMOSPHERIC GEOPHYSICAL AND ASTRONOMICAL SERVICE ADMINISTRATION





UNIVERSIDAD DE CHILE



Científico Mutsumi Ishitsuka y dirigente Luis Sakoda fueron condecorados por gobierno japonés

da, a quienes distinguió con las con-decoraciones de Primavera. El Gobierno japonés otorgó al Dr. Ishitsuka la condecoración en la or-den Tanguióku Shiojushu y en tan-to que a Sakoda le correspondió la

En la ceremonia estuvieron pre-sentes, la esposa, hijos, familiares y amigos de los condecorados, ade-Japonesa, que preside Carlos Saito y el presidente de la Asociación Es-tadio La Unión, Julio Gushiken.

MUTSUMI ISHTISUKA

vida a la investigación en el Observatorio Solar Cosmos, en Huanca-yo, que se convirtió en referencia mundial, porque allí estaba instala-do uno de los más modernos coro-

nógrafos.

I 1992 fue nombrado Director del Observatorio del Centro de Investigación Geofísica de Ancón. Mutsumi Ishitsula ha recibido múltiples distinciones, entre ellas, el título Doctor Honoris Causa de la Universidad Nacional de Ingenieria y recientemente, fue inaugurado en Ate el planetario solar, que en homenaje al cientifico apporés lleva su nombre. En la ceremonia, Mutsumi Ishitsuka estuvo acompañado por su essuka estuvo acompañado por su essuka estuvo acompañado por su es-

LUIS SAKODA

Luis Sakoda Shinyashiki es descen-diente de la prefectura de Kagoshi-ma. Ha desempeñado los máximos



nikkei, como la presidencia de la Asociación Estadio La Unión, en 1970 y de la Asociación Panameri-cana Nikkei filial Perú, en 1980.

misión de Festejos del 90 aniver-sario de la Inmigración Japonesa al

ción de la Clínica Centenario, la Aso-ciación Emmanuel y Perú Kagoshi-ma Kenjinkai.

Luis Sakoda acudió al acto de condecoración acompañado de su esposa, Juana Sakoda.

Sakoda.

Sakoda.

Sakoda.

Desde Japón llegarán dos telescopios para Región Ica

cospios que serán destinados a la Región Ica. "Uno de ellos está conformado por un conjunto de seis telesco-pios solares que se encuentran operande ne el Observatorio Hida, Universidad de Kyoto, los cuales serán instalados en el mes de se-tiembre en la Estación Solar de la Universidad Nacional San Luis Gon-zaga de Ica. Per ello se ha plani-ficado el entrenamiento de dos jó-venes estudiantes en el manejo de este telescopio en el Observa-

El Dr. Musumi Ishitsuka anunció la liegada desde Japón de dos telecopios que serán destinados a la Región Ica Región Ica Región Ica de los está conformado por un conjunto de seis telescopios solares que se encuentran operando en el Observatorio Hida, Universidad de Kyoto, los cuales serán instalados en el mes de se-tiembre en la Estación Solar de la Universidad Nacional San Luis Gontaga de Ica, Por ello se ha planido de la Universidad nacional San Luis Gontaga de Ica, Por ello se ha planido de la Universidad Nacional San Luis Gontaga de Ica, Por ello se ha planido de la Universidad Nacional San Luis Gontaga de Ica, Por ello se ha planido de la Universidad Nacional San Luis Gontaga de Ica, Por ello se ha planido de la Universidad Nacional San Luis Gontaga de Ica, Por ello se ha planido de la Confección que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la la condecoración que le hizo llegar el Gobierno Japonio de la condecoración que le hizo llegar el Gobierno Japonio de la condecoración que le hizo llegar el Gobierno Japonio de la condecoración que la lacación por lacación po quien realizando una colecta logró recaudar los fondos necesarios para la construcción de este tele-







INSTITUTO GEOFISICO DEL PERU, PERU



NATIONAL ASTRONOMICAL OBSERVATORY **OF TARIJA BOLIVIA**



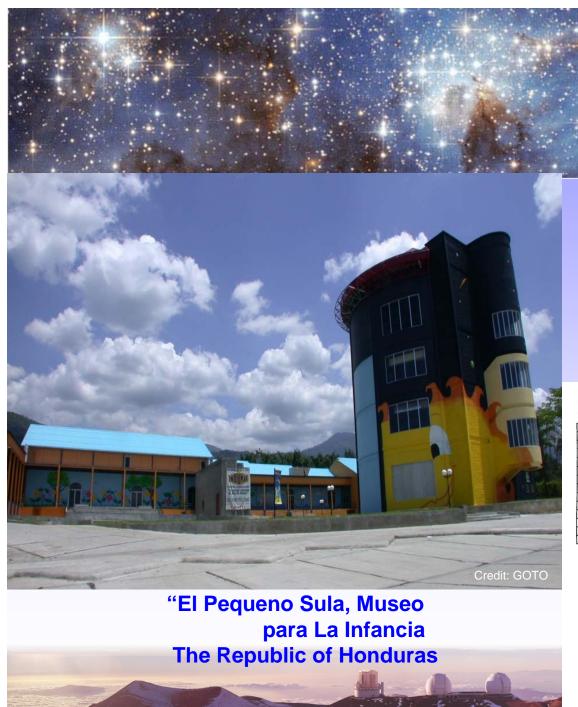
UNIVERSIDAD DE COSTA RICA — THE REPUBLIC OF COSTA RICA



Credit: GOTO



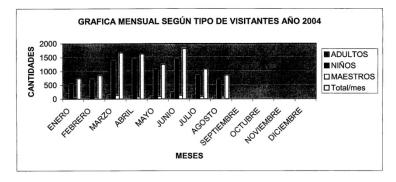






EL PEQUEÑO SULA, MUSEO PARA LA INFANCIA GRAFICA MENSUAL SEGÚN TIPO DE VISITANTES EN EL PLANETARIO AÑO 2004

MES	ADULTOS	NIÑOS	MAESTROS	Total/mes	No. DÍAS	PROM. X DÍA VISIT.
ENERO	172	543	27	742	11	52
FEBRERO	82	718	50	850	12	69
MARZO	147	1391	130	1,668	18	90
ABRIL	94	1451	73	1,618	15	94
MAYO	93	1063	87	1,243	19	59
JUNIO	243	1444	127	1,814	19	87
JULIO	145	850	80	1,075	21	48
AGOSTO	101	695	60	856		#¡DIV/0!
SEPTIEMBRE	0	0	0	-		#¡DIV/0!
OCTUBRE	0	0	0	-		#¡DIV/0!
NOVIEMBRE	0	0	0	-		#¡DIV/0!
DICIEMBRE	0	0	0	-		#¡DIV/0!
TOTAL	1077	8,155	634	9,866		





Office of Havana City's Historian "Rosa Elena Simeon" Science and Technology Cultural Center







Universidad de Costa Rica The Republic of Costa Rica

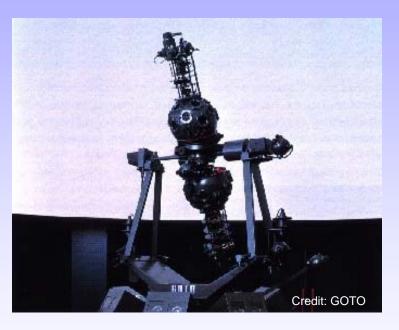


HAYA CULTURAL CENTRE FOR CHILD DEVELOPMENT



TIN MARIN CHILDREN'S MUSEUM









MEGHNANDSAHA PLANETARIUM, UNIVERSITY OF BURDWAN, INDIA



ANNA SCIENCE CENTER PLANETARIUM, INDIA















MUSEO NACIONAL HISTORIA NATURAL

DOMINICAN REPUBLIC



Japan International Cooperation Agency (JICA);

Japanese astronomers and/or engineers are sent to give necessary technical training to the local staff members of the institutions that received the telescopes and/or planetariums.

Public Observatories in Japan;

six-month astronomy research and observation training courses have been provided to staff members of the institutions that received a telescope with a CCD camera.

Assistance by the National Astronomical Observatory of Japan (NAOJ) with the IAU

- 1) Work with the local government to help prepare an application to the CGA.
 - Developing plans for programs.
 - Give IAU endorsement to the application.
- 2) Support follow-up training.
 - Organize seminars and workshops
 - Support collaborative research projects



For more information

The TRIPOD concept have been developed during the last two decades of efforts by Profs. Masatoshi Kitamura, Hans J. Haubold, and many others. Additional information on this subject can be found in the following papers.

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