

# Study of ionospheric perturbations using GNSS ground station and AWESOME VLF data

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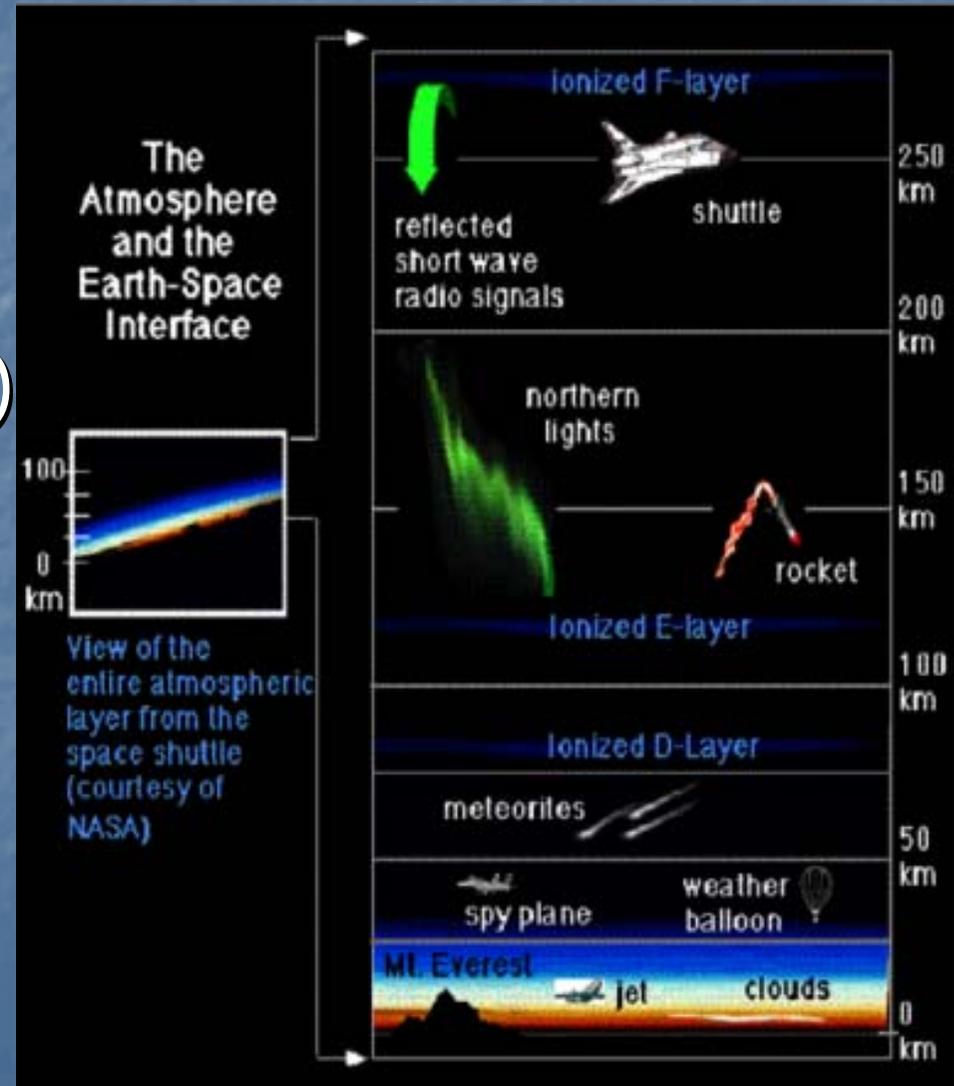
Ulugh Beg Astronomical Institute, Tashkent,  
Uzbekistan

# Outline

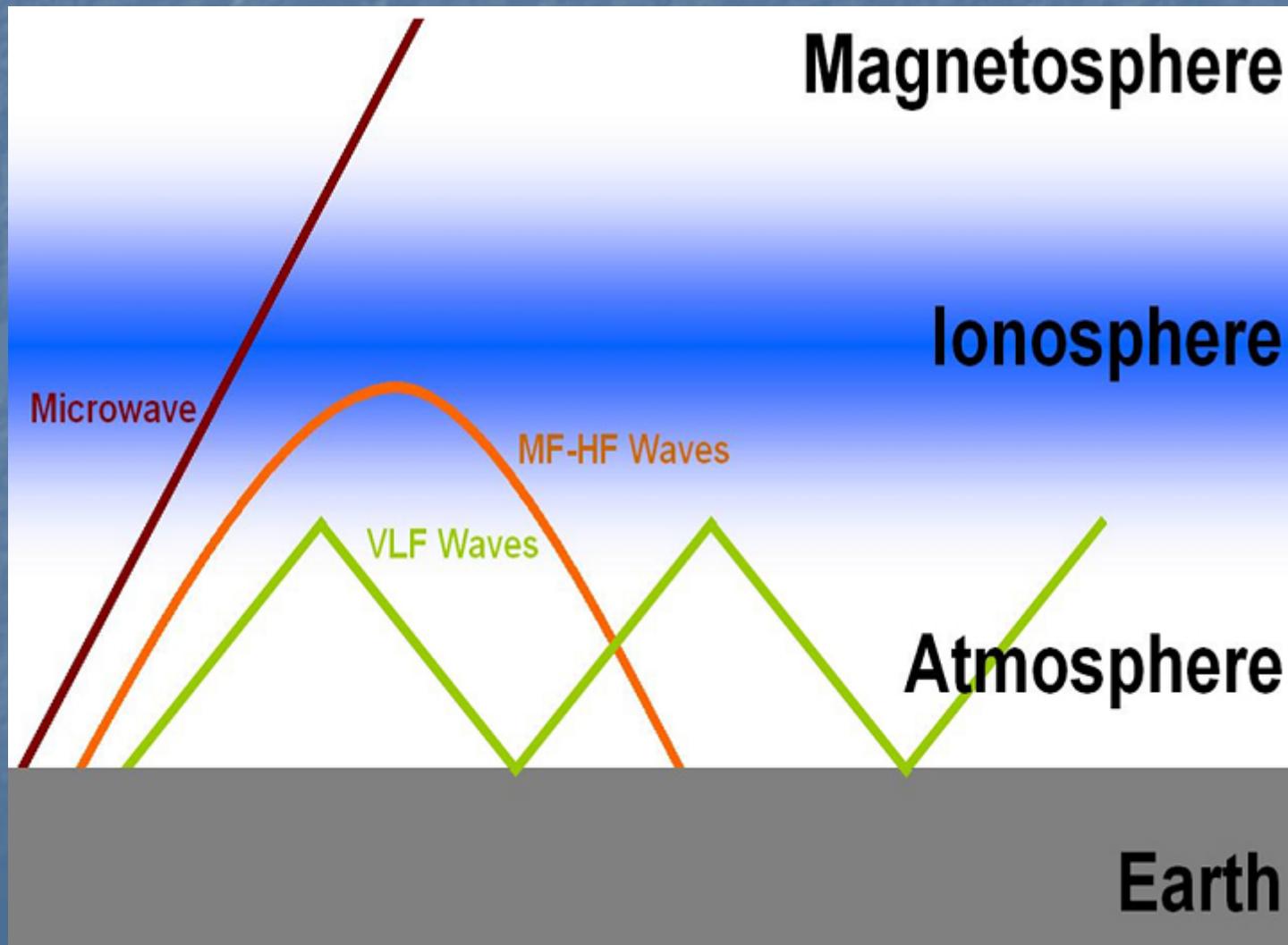
- Introduction
  - Ionospheric layers
  - TEC
- Uzbekistan permanent GNSS ground stations
- TEC variations
- AWESOME project
- Tashkent AWESOME station
- Monitoring of space weather
- Conclusion

# Ionosphere

- D-layer (60-90 km)  
 $n \sim 10^2$  e/cc
- E-layer (90-120 km)  
 $n \sim 10^5$  e/cc
- F-layer (130-1000)  
 $n \sim 10^6$  e/cc



# Earth-ionosphere waveguide



# Ionosphere and total electron content

GPS measurements use time delay between radio signals at two frequencies

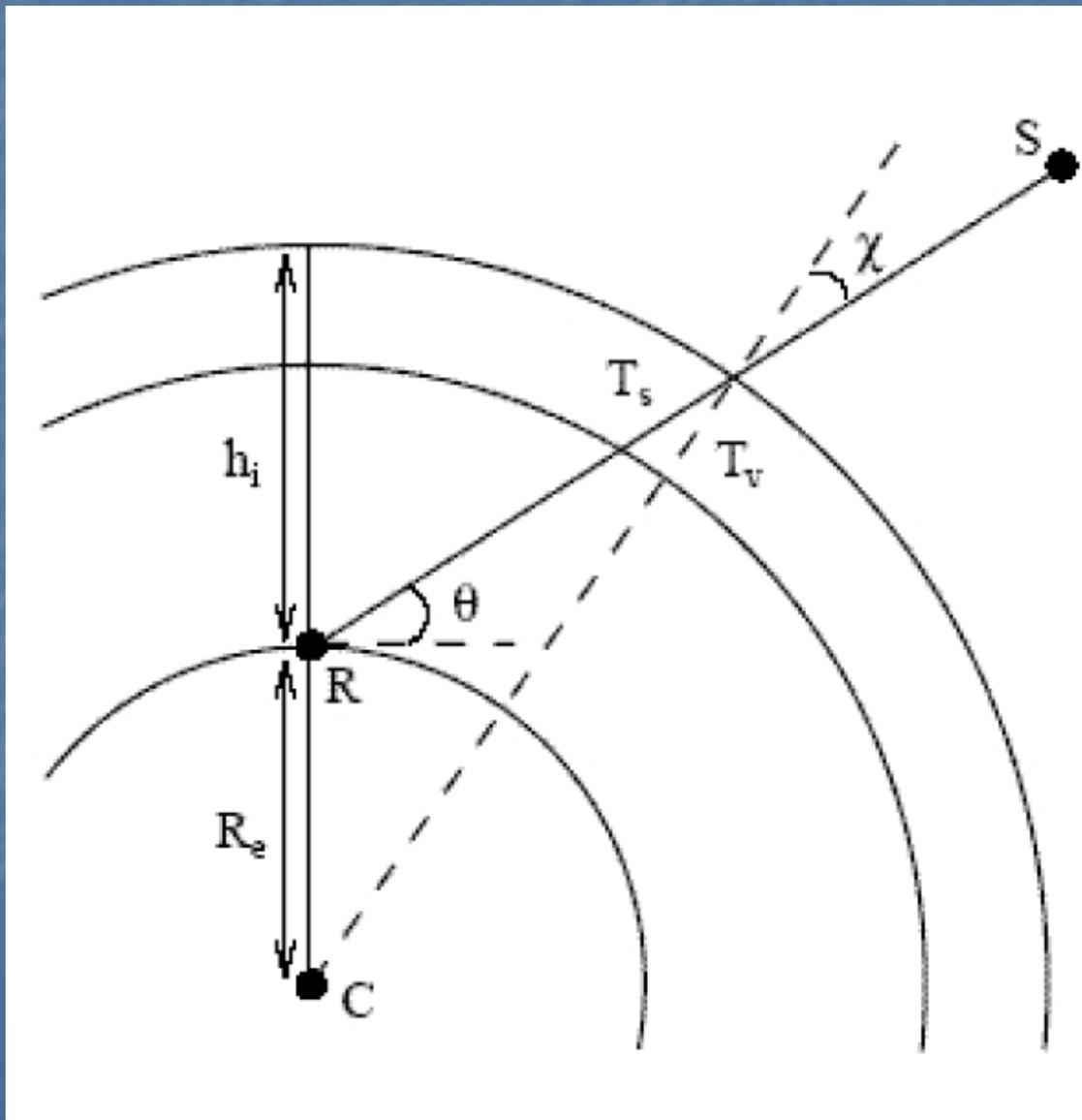
$$L_1 = 1575.42\text{MHz} \quad \text{and} \quad L_2 = 1227.6\text{MHz}$$

Pseudorange

$$P_i = \rho + c(dT - dt) + \Delta_i^{\text{iono}} + \Delta^{\text{trop}}$$

- Layers cause delay of incoming signals
- Signals carry information about the state of the ionosphere
- Thus, signals can be used for remote sensing

# sTEC and vTEC



# GNSS permanent stations in Uzbekistan

1996 Kitab station (KIT3)

- 2001 Tashkent (TASH)

Installed by German Research Center GeoForschungsZentrum (GFZ), Potsdam



# Tashkent station

Receiver: Javad Delta

Antenna: Javad



# Web-interface



## Tashkent / GSS Monitor

GeoForschungsZentrum Potsdam

GPS Receiver

Meteo Station

Options used

Software

Transfer Dir.

Logfiles

sysinfo

webmin

### GPS Receiver Monitor

actual status at 2010/11/30 08:55:16  
GPS TIME: 1612/2 08:55:31 (975142530 sec); diff to computer time: 0 sec  
Sitename: tash, Elevation mask: 0, Sample Int.: 1, Receiver-Version: 3.1.7 May,31,20, DELTA\_00200 OEM\_15658

Satellite: 18

current tash16122iz2700.jps (Epochs: 631, Obs: 631, Ant: 631, Alma: 631, Ephe: 2, Health: 115)  
last tash16122iz1800.jps (Epochs: 900, Obs: 900, Ant: 900, Alma: 900, Ephe: 10, Health: 167)

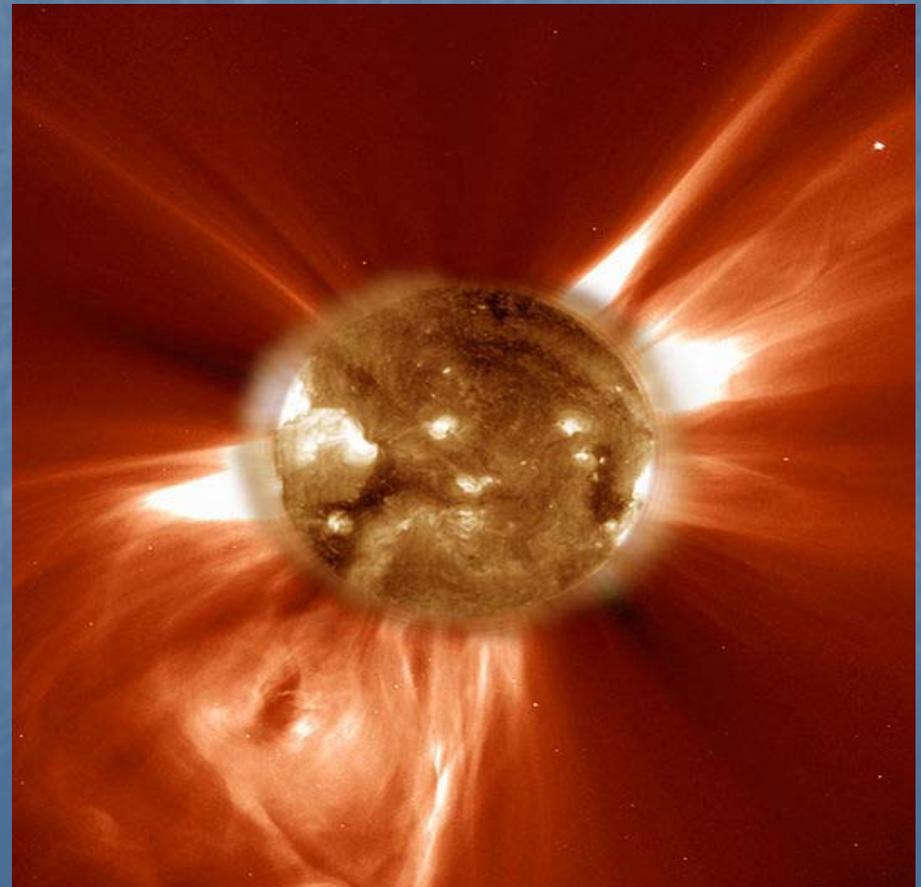
/home/gghmet/src/GPS\_read.pl version is 2.2b.9a

### GPS-Datalogger: Tail of 202 lines of LOG-file: [prt/prt\\_GPS\\_2010\\_334](#)

```
08:42:46: no time synch necessary
08:44:45: tash16122iz1800.jps (Epochs: 900, Obs: 900, Ant: 900, Alma: 900, Ephe: 10, Health: 167)
08:44:45: move /home/gghmet/wrk/gps/tash16122iz1800.jps to /home/gghmet/cmp_mv
08:44:45: create new file: tash16122iz2700.jps; used GPS-time: 975141900
08:44:45: Bytes read / print: 1856348 / 1854566
```

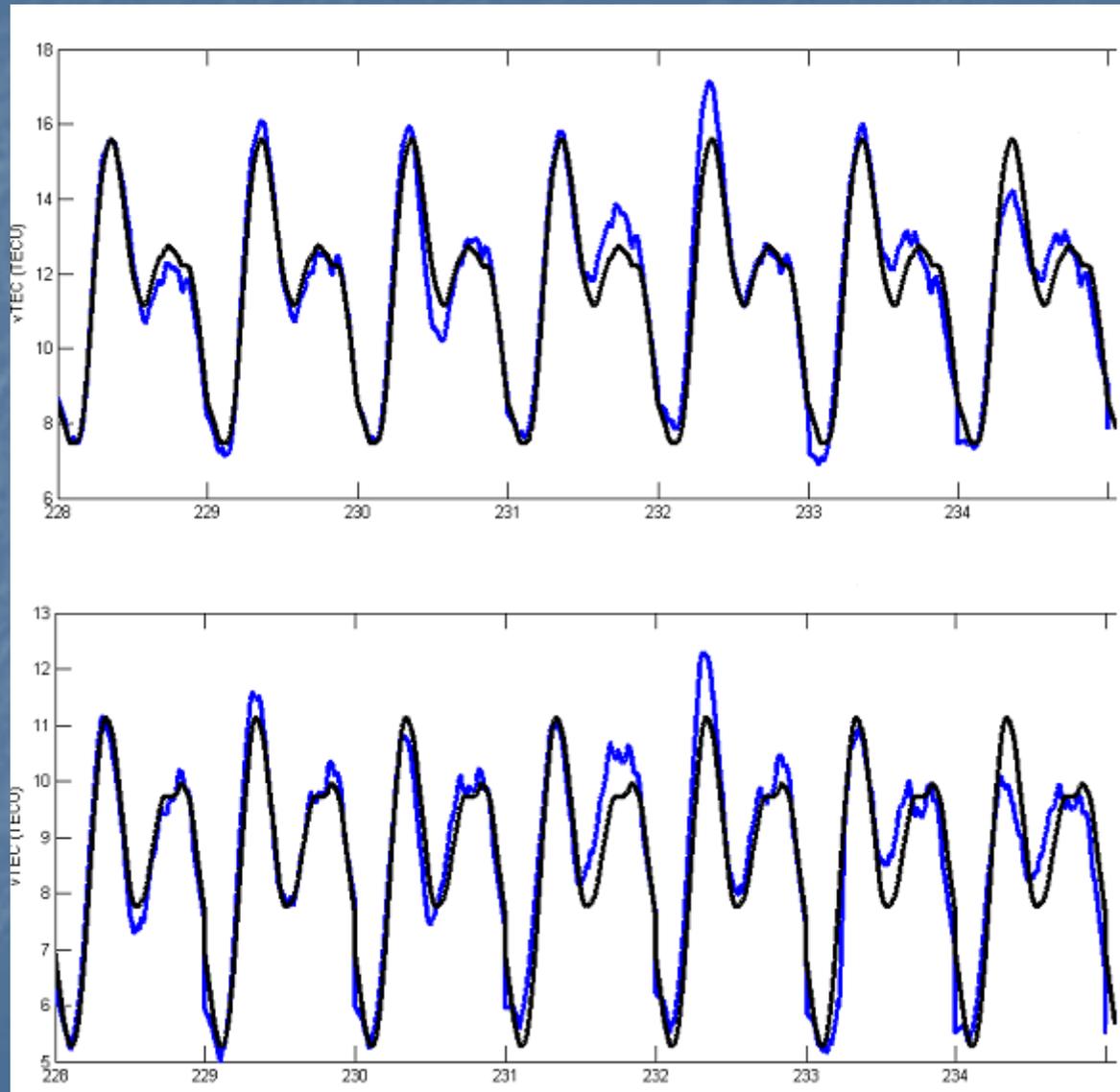
# Solar Flares

Classification  
of X-Ray Solar  
Flare: A, B, C,  
M, X

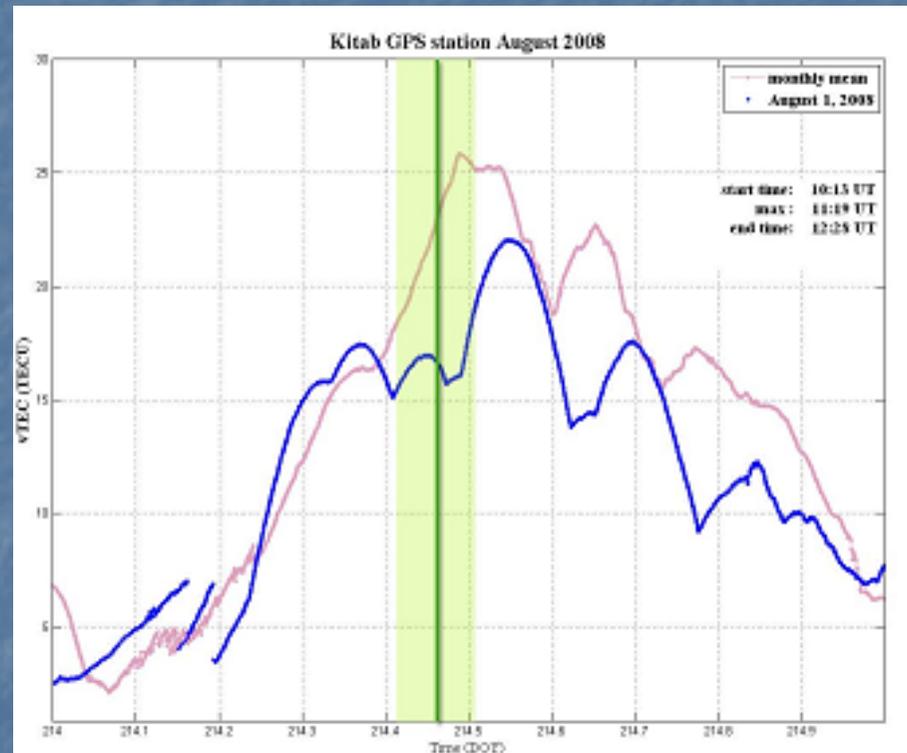


# Analysis method

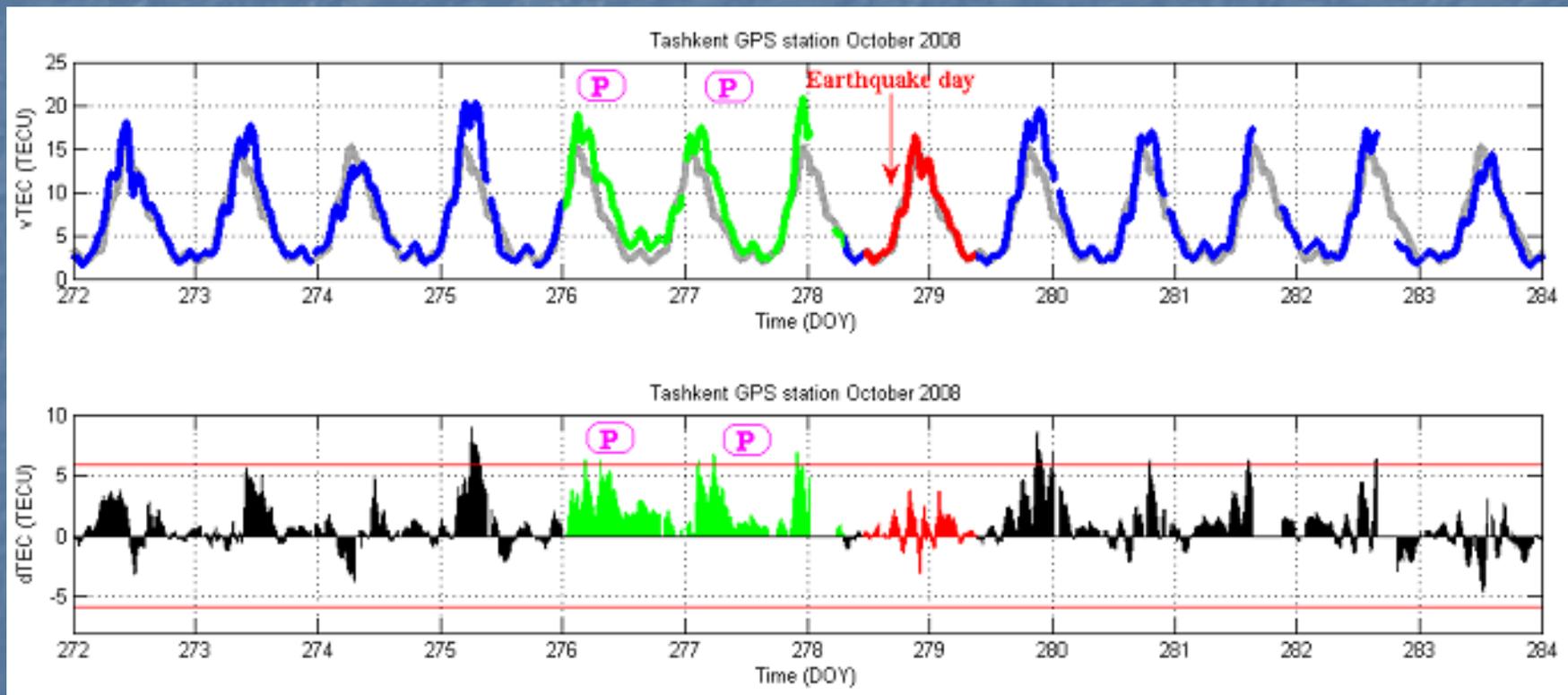
- Monthly mean of TEC around the day of interest is taken as a reference



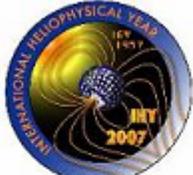
# Solar eclipse



# Anomalies of TEC

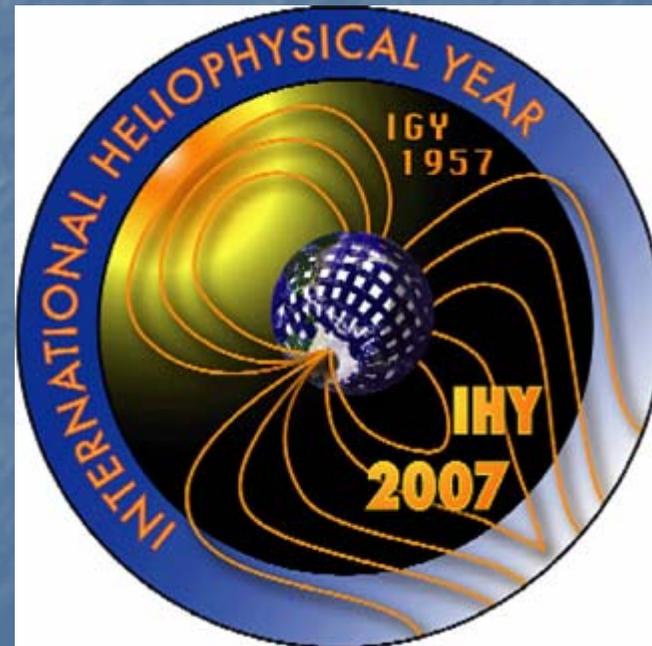


# Space weather monitors

|                                                                                                                                                                                                                |                                                                                                                                                                        |            |                                                                                                                  |
|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|------------------------------------------------------------------------------------------------------------------------------------------------------------------------|------------|------------------------------------------------------------------------------------------------------------------|
| <br><br>UNITED NATIONS                      | <br> | Umran Inan | <b>Prof. Hans J. Haubold</b><br><i>UN Office for Outer<br/>Space Affairs<br/>Vienna International<br/>Centre</i> |
| <br> STANFORD<br>S O L A R<br>C E N T E R | Deborah Scherrer                                                                                                                                                       |            |                                                                                                                  |

# VLF AWESOME project

A tmospheric  
W eather  
E lectromagnetic  
S ystem for  
O bservation  
M odeling and  
E ducation



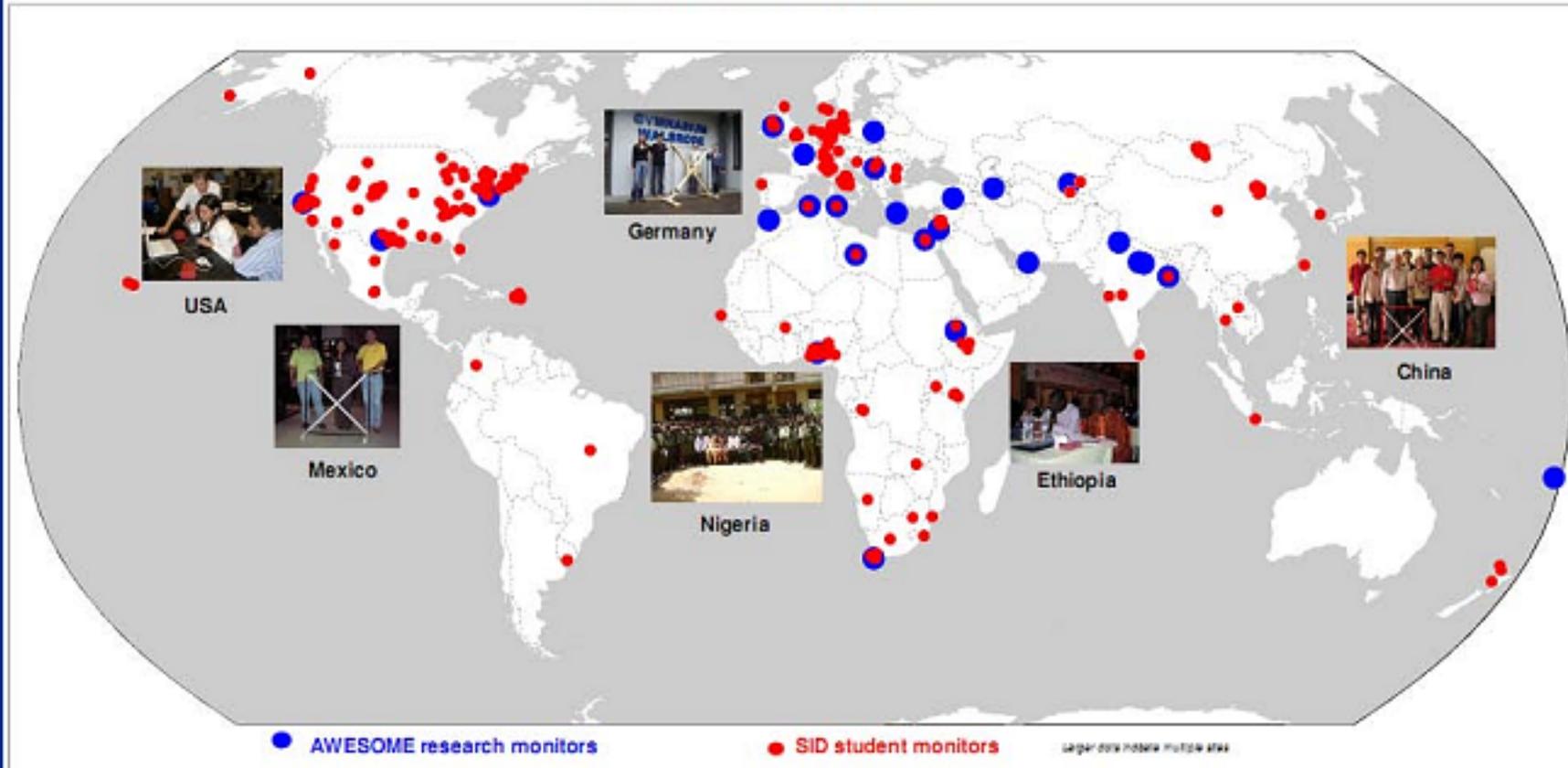
# World SID & AWESOME Sites



## Space Weather Monitor Sites



IHY Distribution 2007-2009



USA



Mexico



Romania



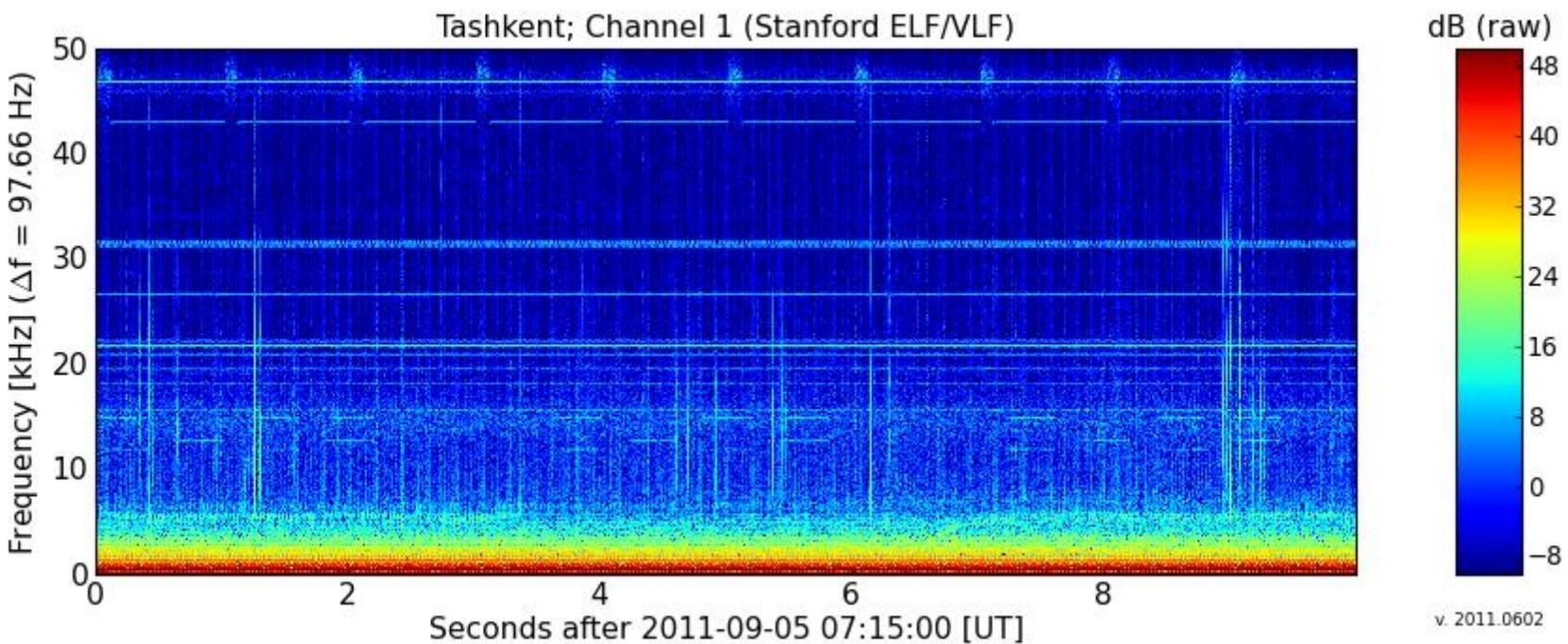
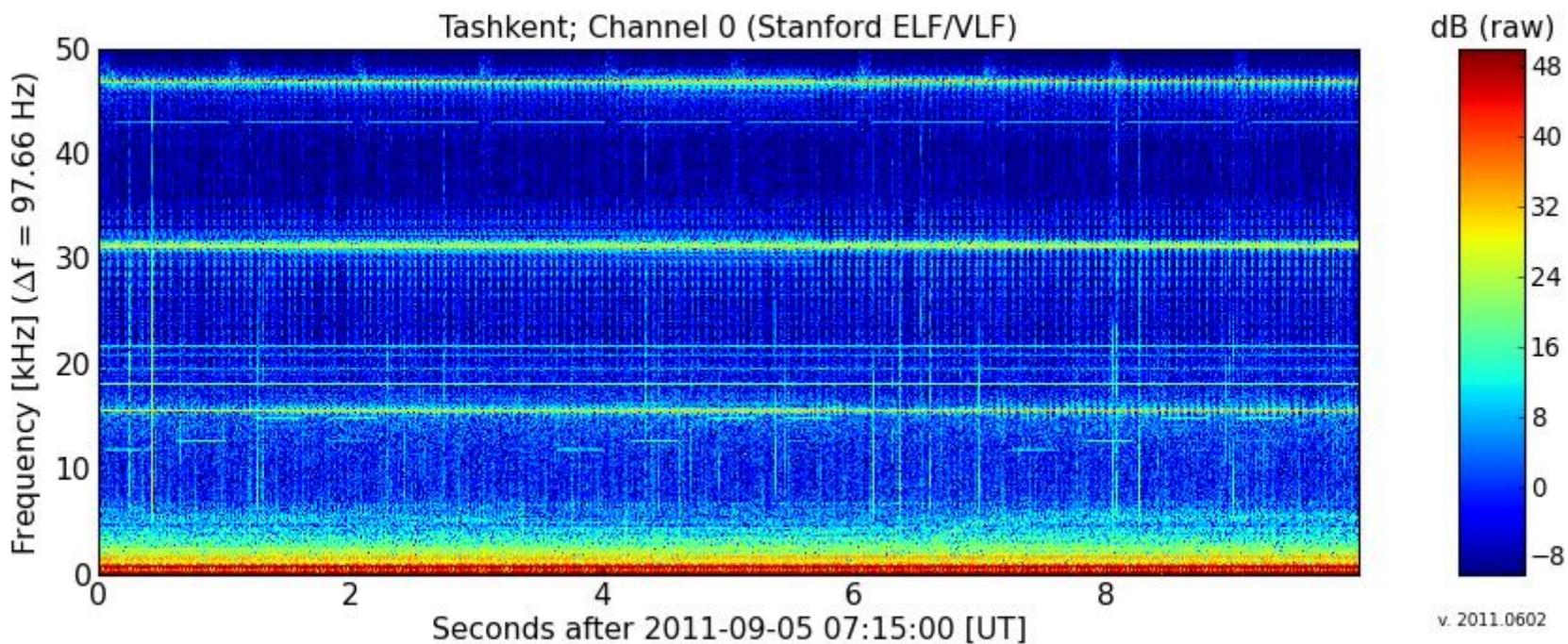
Lebanon

Thailand

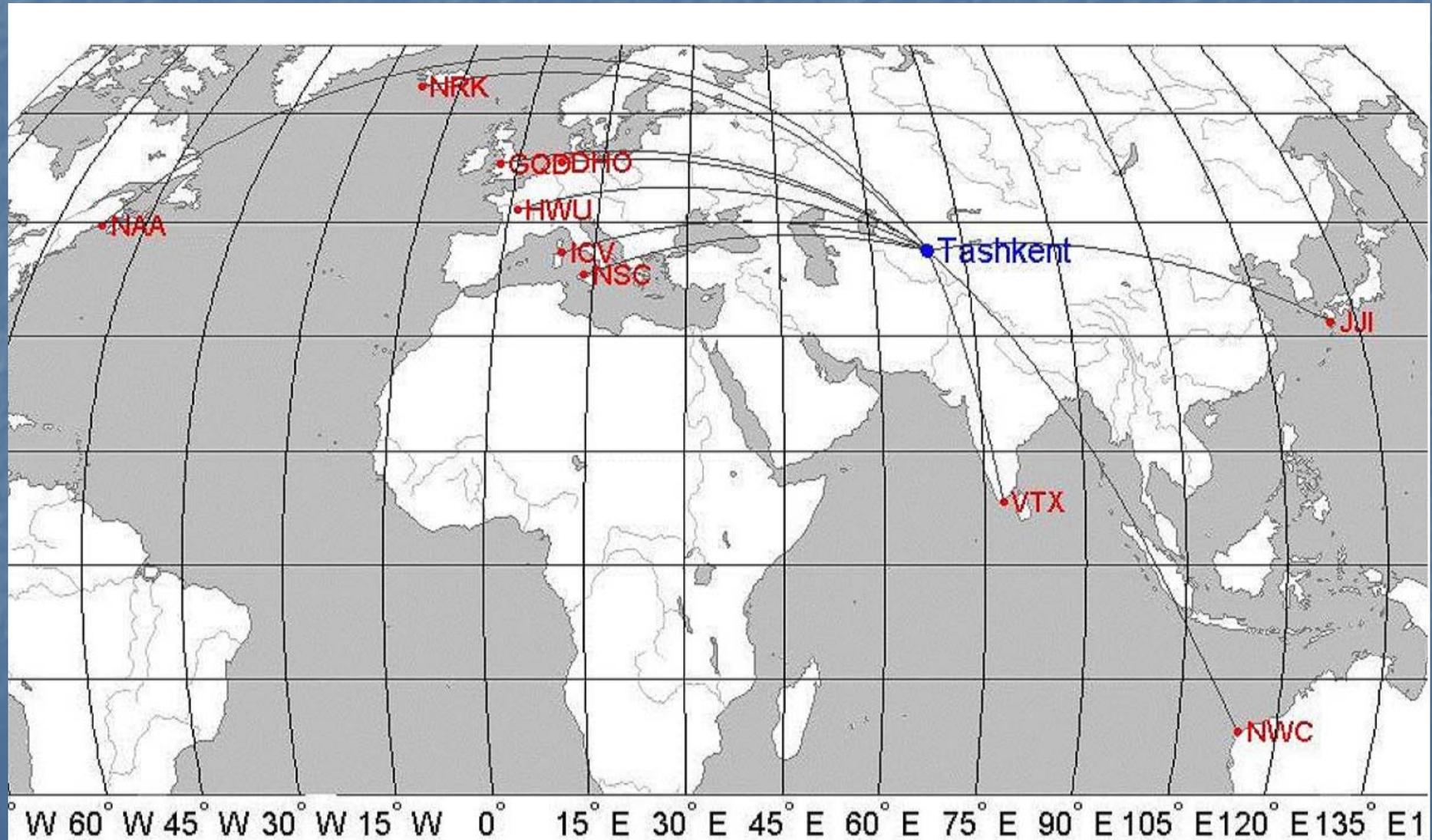


# Tashkent AWESOME station

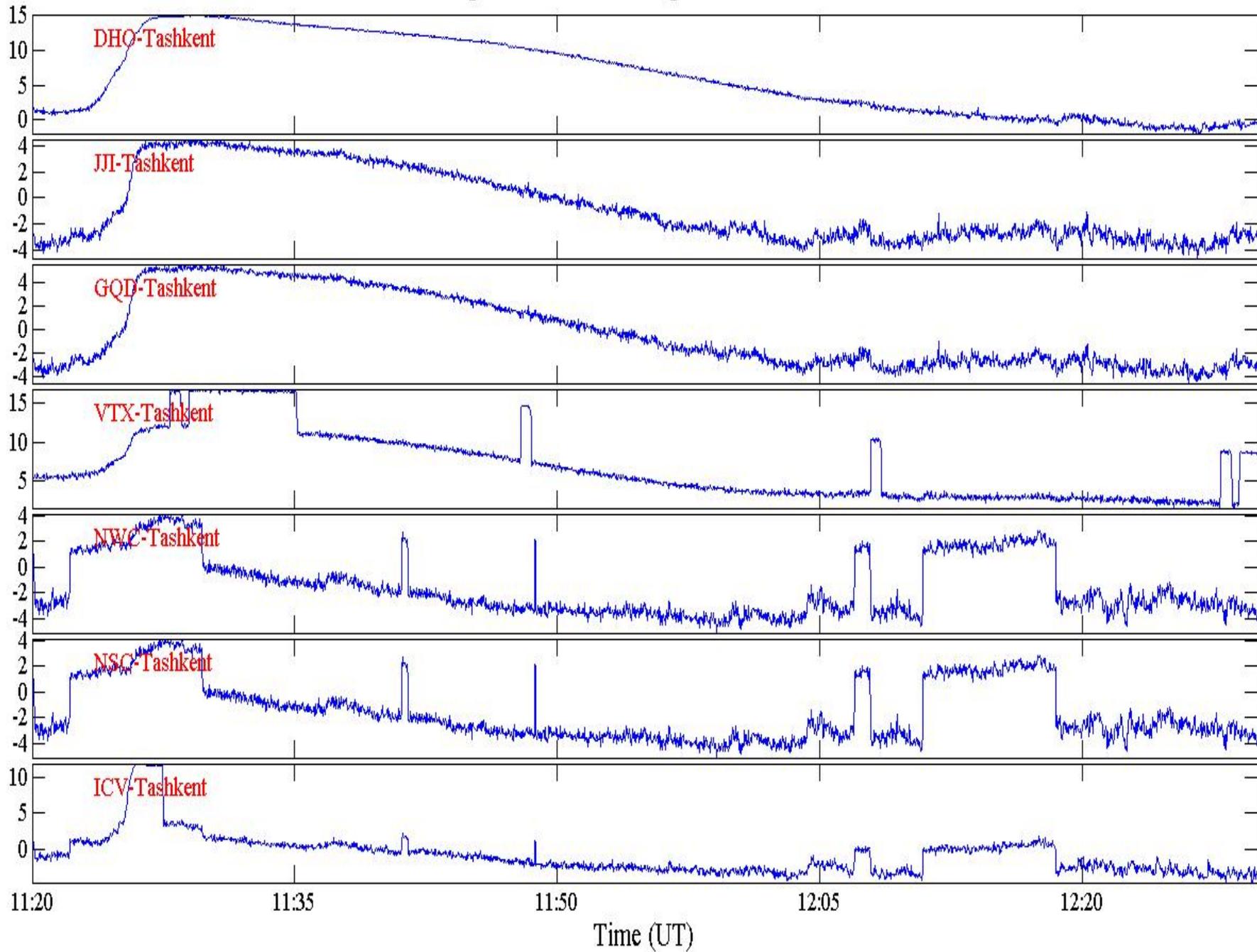




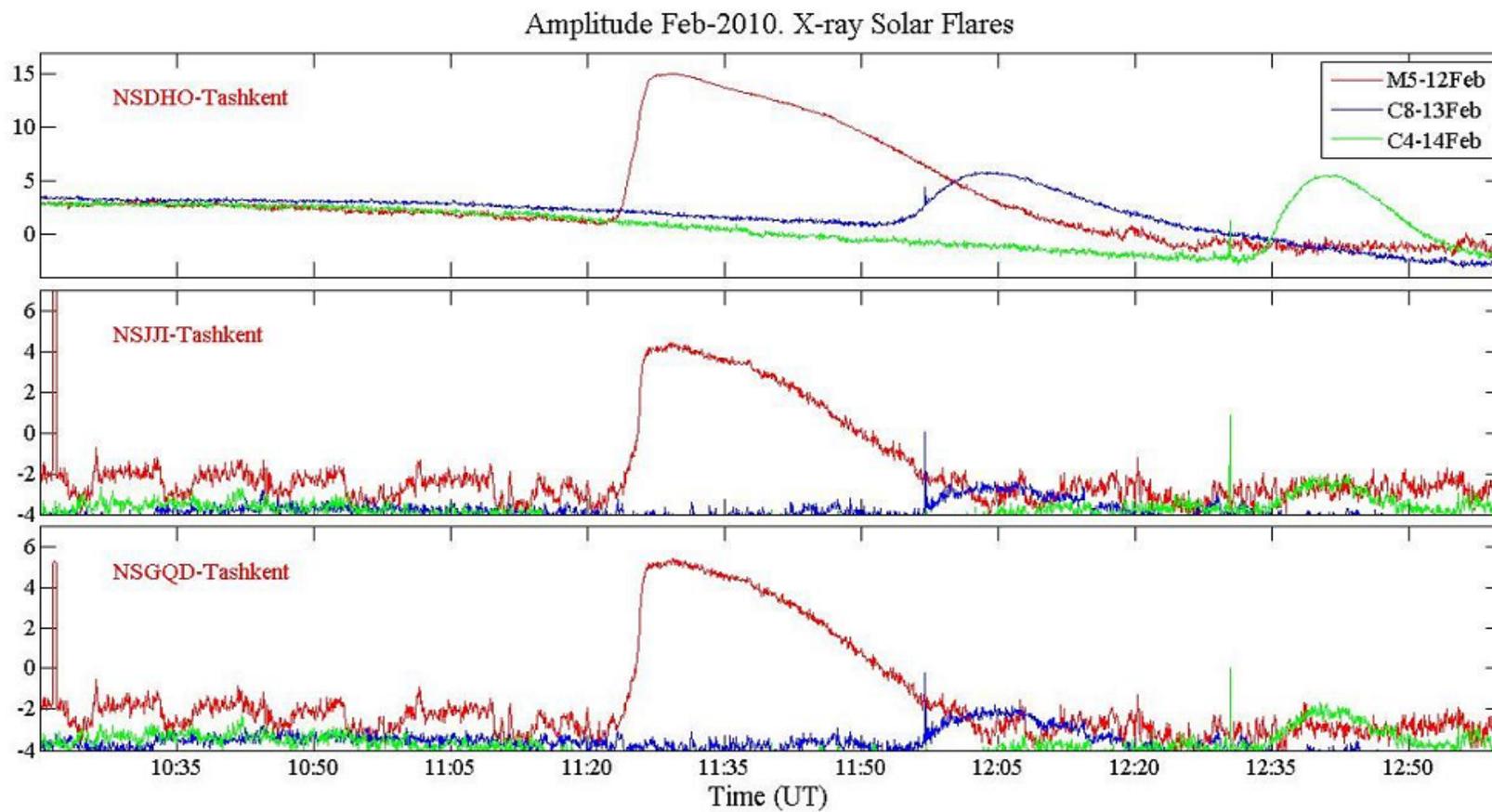
# Transceivers



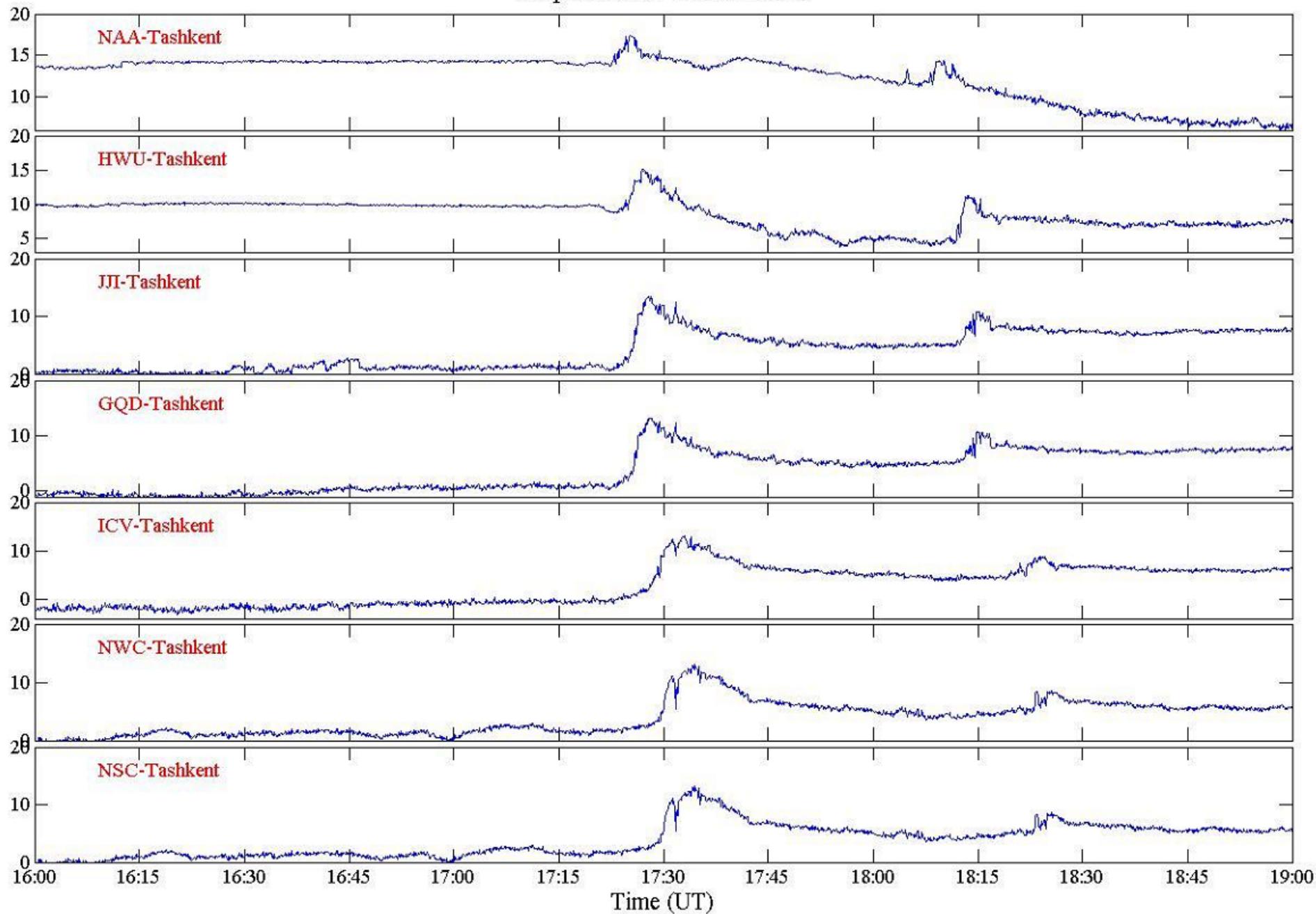
# Amplitude Data Starting On 12-Feb-2010



# X-ray Solar Flare are observed during February 12-14, 2010

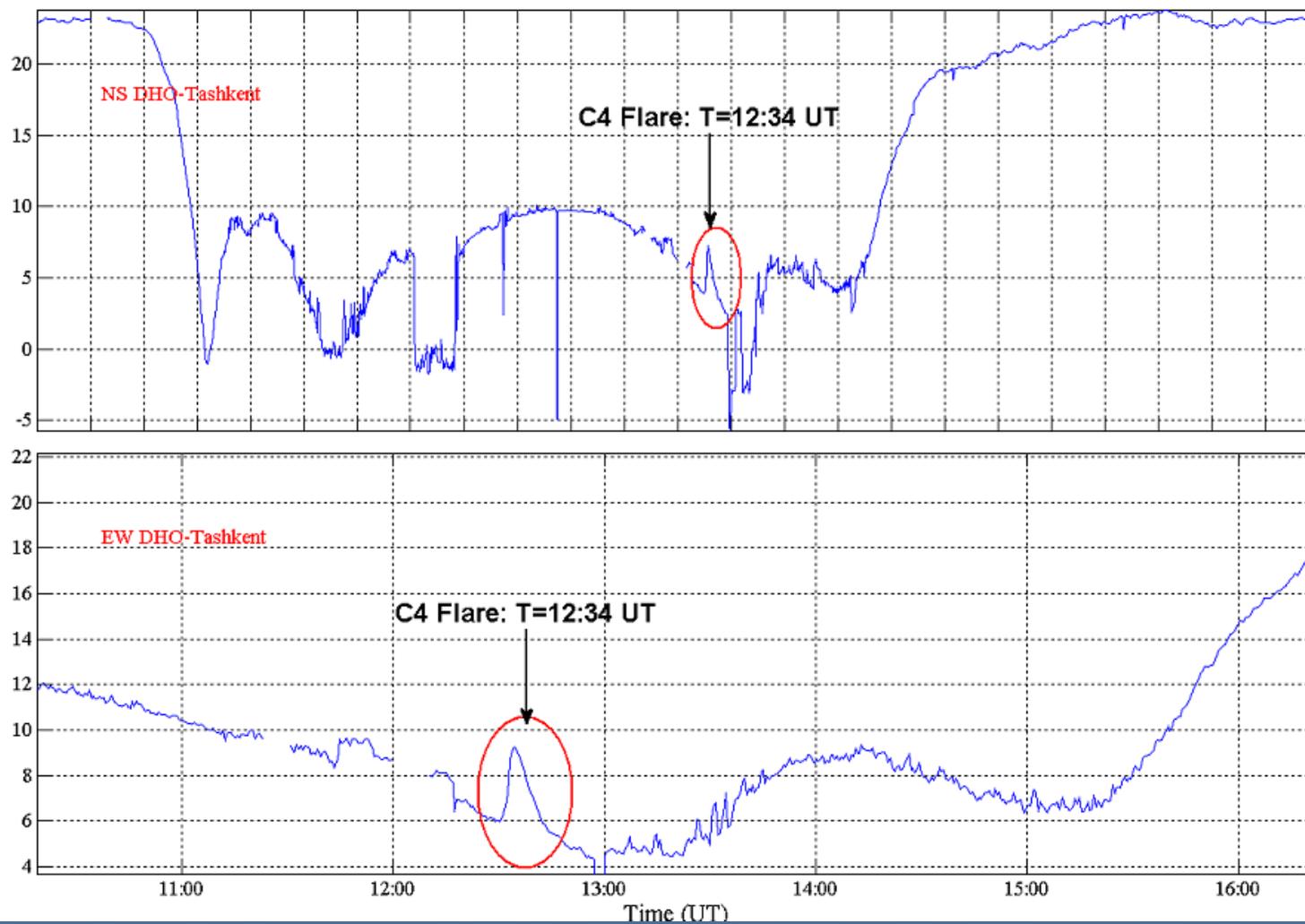


# Amplitude On 09-Mar-2010



## X-ray Solar Flare are observed on February 10, 2011

Amplitude Data Starting On 10-Feb-2011

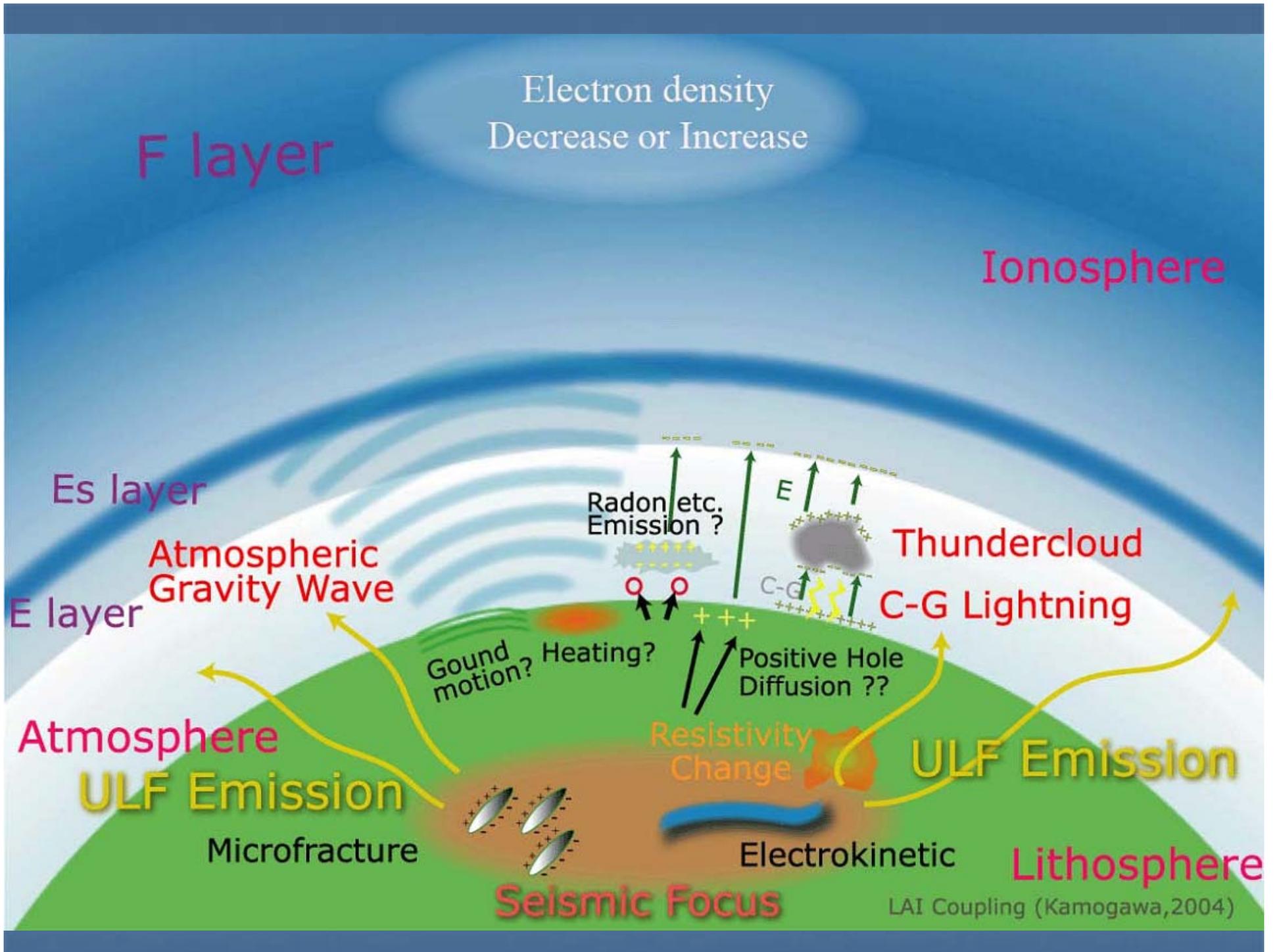


# Conclusion

- Ionospheric TEC variations can be used as a monitor of space weather
- Two independent sensors of the ionosphere can complement each other in better understanding
- Construction of a web-based online near real time space weather monitor is in progress

Thank you





## Lithosphere-Atmosphere-Ionosphere Coupling - Types

- **Electromagnetic Coupling:** Connected with the direct penetration of DC electric field induced due to the appearance of Seismic-related electric charges on the Earth's surface. It can lead to substantial modifications of ionospheric properties.
- **Chemical Coupling:** Determined by the variation of the fair weather electric field in the lower ionosphere due to the enhancement of conductivity of lower atmospheric layer ionized by radon emanating from Seismic faults.
- **Dynamic Coupling:** Implies influence of atmospheric wave processes originating near the Earth surface on the lower ionosphere.

## Ionospheric Precursors: For different ionospheric layers

- **F-Layer:** Critical frequency of F layer ( $f_oF_2$ ) and **TEC**
- **E-Layer:** Critical frequency of Sporadic E-Layer ( $f_oE_s$ )
- **D-Layer:** Phase and Amplitude of **ELF/VLF signals** from navigational transmitters

Out of these three types lot of work have been done by Japanese and Russian group on D-region precursors studies. (Pulinets et al 1991, Lipervosky et al 2000, Hayakawa 1996, Gokhberg et al., 1982; Gufeld et al., 1992 )