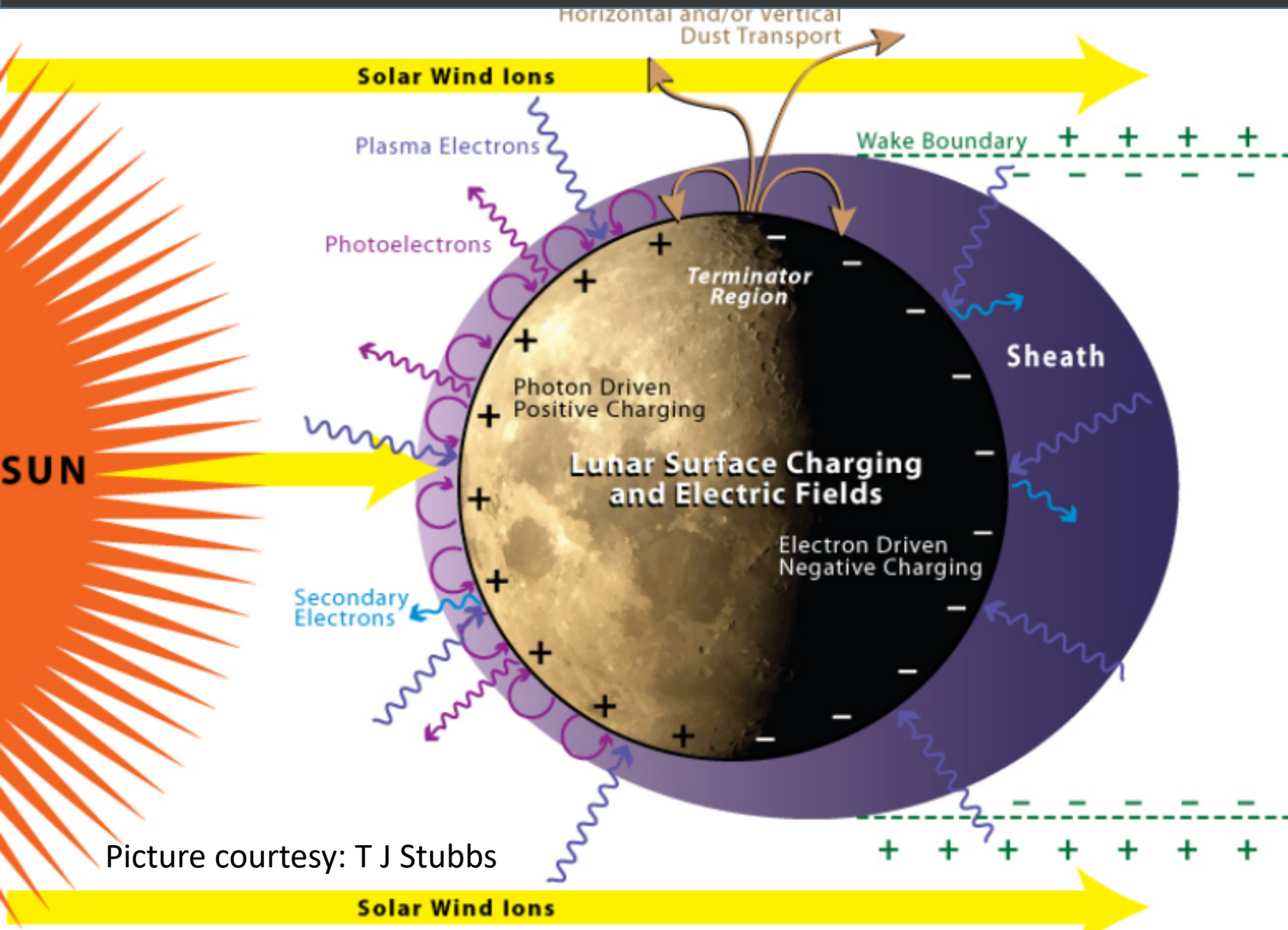




Chandrayaan3 Mission

# Scientific Motivation for Indian Lunar Exploration



Picture courtesy: T J Stubbs

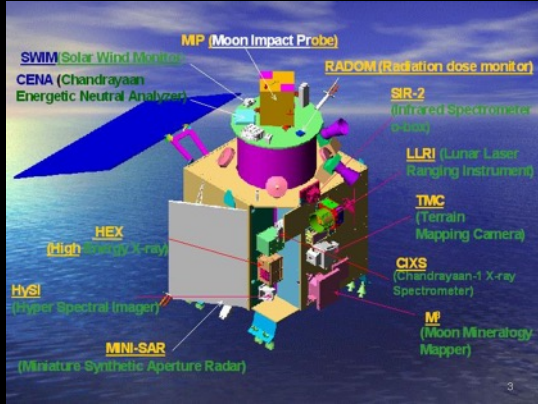
- Airless body → All the wavelengths unhindered (since no thick atmosphere)
- No Global Magnetic Field → Exposed directly to the solar wind ( $H^+$ ,  $He^{++}$ , other ions PLUS electrons)

## Domains for Study:

Elemental composition,  
mineralogy, geology;  
Exosphere studies;  
Particle environment

# India's Lunar Science Programme: Scientific Rationale

## Chandrayaan-1 (2008)



### Indian

- ☐ HYSI
- ☐ TMC
- ☐ LLRI
- ☐ HEX
- ☐ MIP

### European

- ☐ SIR-2
- ☐ RADOM

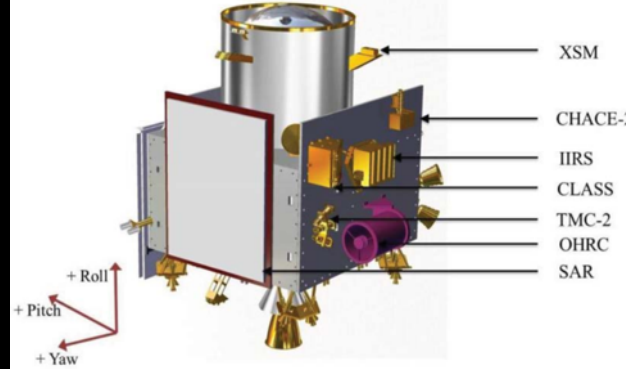
### European-Indian

- ☐ C1xS
- ☐ SARA

### United States

- ☐ M3
- ☐ Mini-SAR

## Chandrayaan-2 (2019)



Orbiter High Resolution Camera (OHRC)

Terrain Mapping Camera – 2 (TMC-2)

Chandrayaan-2 Large Area Soft X-ray Spectrometer (CLASS)

Solar X-ray Monitor (XSM)

Imaging Infra-Red Spectrometer (IIRS)

Dual frequency Synthetic Aperture Radar (DFSAR)

Chandra's Atmospheric Composition Explorer – 2 (CHACE-2)

Dual Frequency Radio Science Experiment (DFRS)

## Chandrayaan-3 (2023)



Instrument for Lunar Seismic Activity (ILSA)

Radio Anatomy of Moon Bound Hypersensitive ionosphere and Atmosphere (RAMBHA) – Langmuir Probe (RAMBHA-LP)

Chandra's Surface Thermo-physical Experiment (ChaSTE)

Alpha Particle X-Ray Spectrometer (APXS)

Laser Induced Breakdown Spectroscope (LIBS)

Spectro-polarimetry of HAbitable Planet Earth (SHAPE)

“Ground truth of the surface and near-surface properties at the landing site”

“Chemical, mineralogical, geologic mapping of the Moon”

# From Chandrayaan-1 to Chandrayaan-2

Ch-1 payloads	Specifications	Ch-2 payloads	Improved Specifications	Enhanced science
<b>HySI</b> <b>SIR-2</b> <b>M3</b>	32 spectral bands 0.93 - 2.4 micron 0.7 – 3.0 micron	<b>IIRS</b>	256 spectral bands, 0.8 to 5.0 micron; Res. 0.020-0.025 micron	Unambiguous detection of OH, water and water-ice signatures.
<b>Mini-SAR</b>	S-band, circular polarimetry	<b>DFSAR</b>	L and S band	Greater depth of penetration (~5-10m i.e twice that of S-band); circular and full-polarimetry
<b>ChACE-1 on MIP</b>	1-100 amu; short duration	<b>CHACE-2</b>	1-300 amu; continuous	Global dynamics of exospheric species
<b>C1XS</b>	0.5 – 10 keV	<b>CLASS</b>	0.8 – 15 keV, ~ 3 times larger area, 12.5 km resolution	High resolution elemental maps, minor elements detection, geotail studies
<b>SXM</b>	1–20 keV; 250 eV at 6 keV	<b>XSM</b>	1 – 15 keV; ~180 eV @5.89 keV	Better resolution and high cadence
<b>TMC-1</b>	Panchromatic	<b>TMC-2 + OHRC</b>	Miniaturised version & Hi-res	Mass reduction & high-res imaging
<b>RO Experiment</b>	Single freq – S band	<b>DFRS</b>	Dual freq.- S and X bands	Vertical profiling of ionosphere, electron density measurements, mitigation of scintillation effects



# From Chandrayaan-2 (global) to Chandrayaan-3 (*in-situ*)

## Chandrayaan-2 Orbiter Payloads

DFRS  
(RO experiment)

CLASS + XSM  
(Surface elemental composition)

CHACE-2  
(Exospheric neutral Composition)

DF-SAR & IIRS  
(Subsurface hydration, surface roughness,  
surface hydration)

## Chandrayaan-3 Lander and Rover Payloads

RAMBHA-LP  
(measurement of the near-surface plasma)

APXS and LIBS  
(Elemental Composition of the surface)

ILSA  
(Moon-quakes, ground acceleration)

CHASTE  
(Thermophysical properties of the regolith)

Complementary obs.

Ground truth on elemental  
composition

Internal release

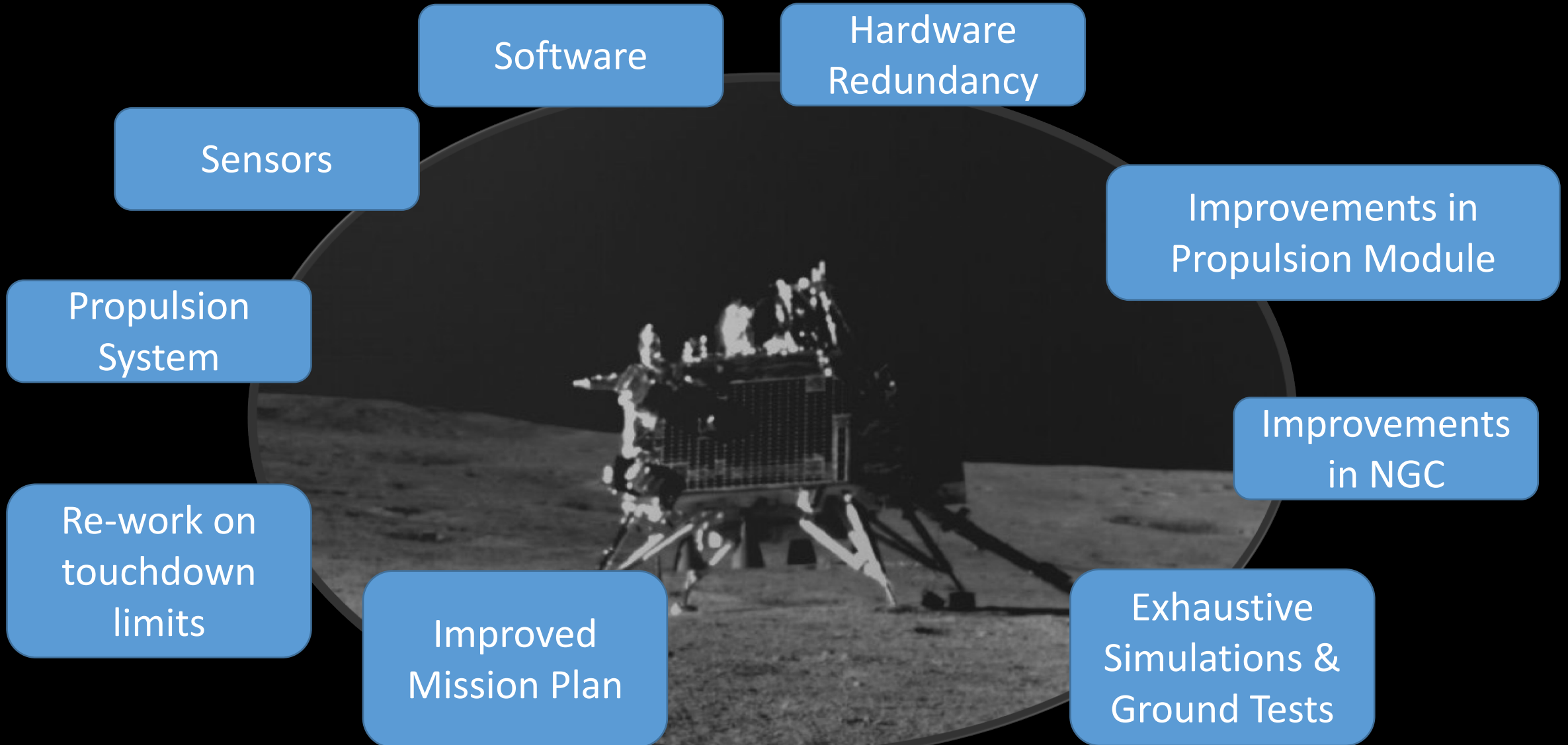
Neutral-to-plasma  
connection

Correlations & Constraining  
the models

Local in-situ observations, plus looking for any variations of properties in the vicinity of the landing site

# Improvements in Chandrayaan-3 Lander

Based on the lessons learnt from Chandrayaan-2 Lander



# Choice of the Landing Site of Chandrayaan-3

## Scientific

1. Higher latitudes are relatively unexplored
2. Unique condition – Sun shines at grazing angle  
→ lesser deposition of energy, solar wind ions and electrons
3. Scientifically interesting Topography

## Technological

1. Conditions on global slope
2. Conditions on the sizes of boulders and crates
3. Shadow analysis; power generation
4. Communication with Earth





# The Journey of Chandrayaan-3



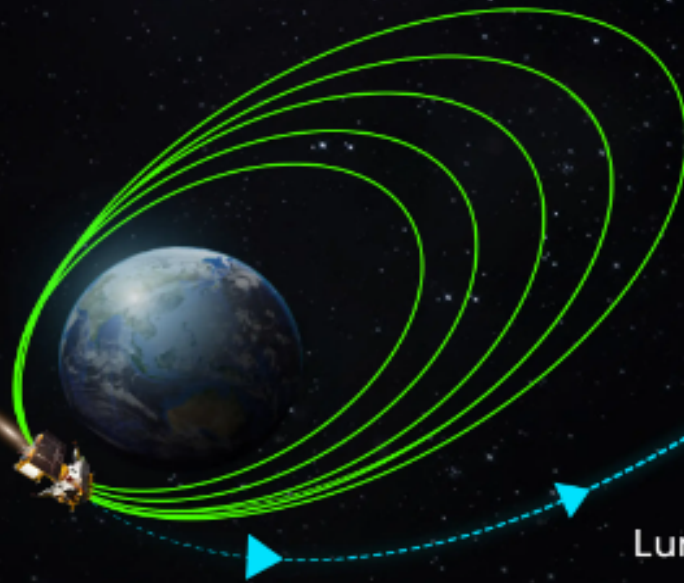
**July 14, 2023, at 14:35 Hrs. IST from the  
Second Launch Pad, SDSC-SHAR, Sriharikota**

# The Journey of Chandrayaan-3 (contd.)

August 01, 2023, 00:15 Hrs. IST



Earth-bound Maneuvers



Lunar Transfer Trajectory

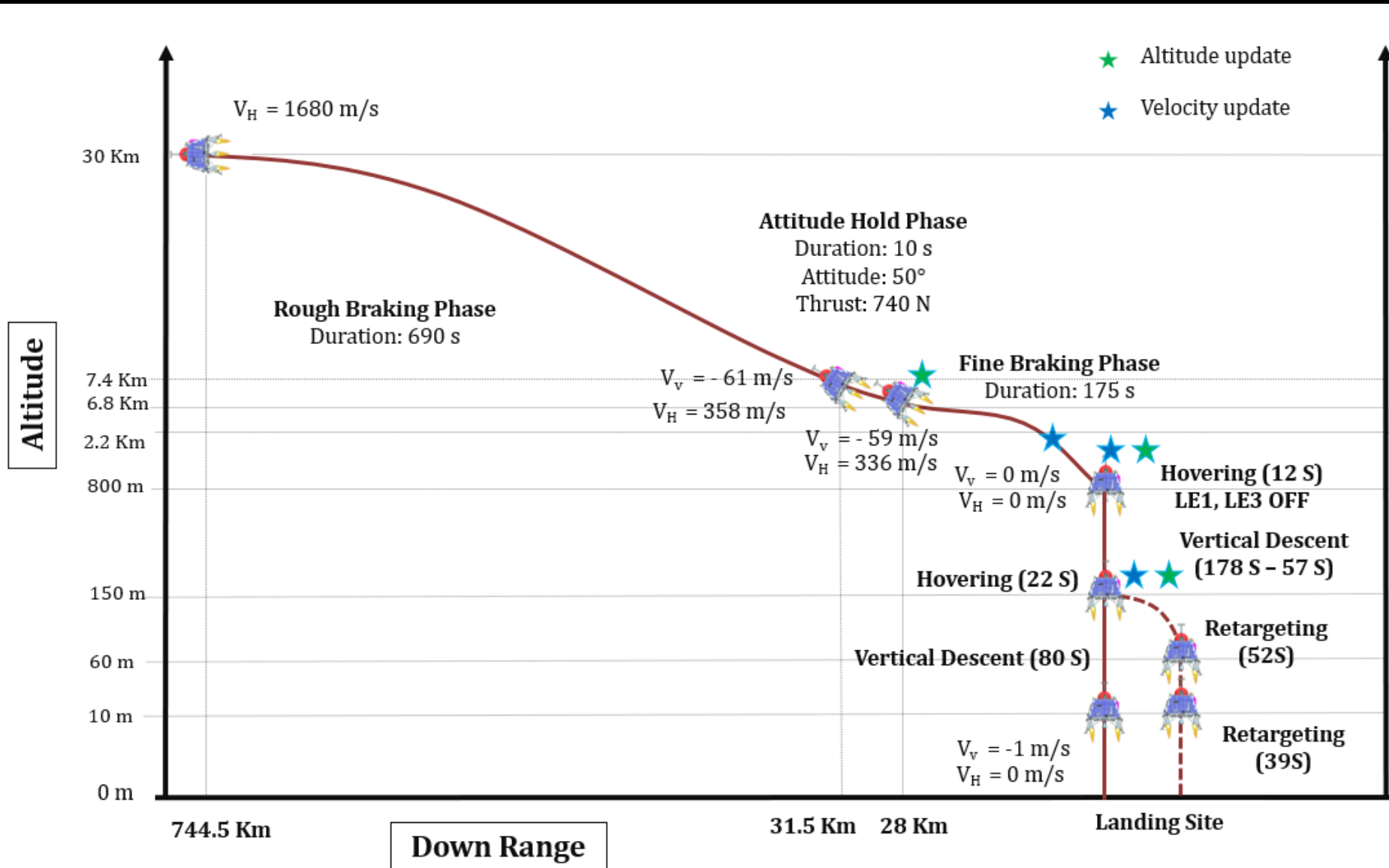


Orbits are not up to the scale

Moon-bound Maneuvers



# Landing Sequence of Chandrayaan-3



23 August, 2023, Wednesday, 18:04 IST



CHANDRAYAAN 3

## Achieved soft landing on the Moon

India:

The 1<sup>st</sup> country to have soft-landed in higher lunar latitude  
The 4<sup>h</sup> country to have soft-landed on the Moon





# Chandrayaan-3 Mission Update

## Original Mission Objectives

- To soft-land the lunar lander-rover module on the pre-determined landing site at the Southern higher latitudes of the Moon
- Demonstration of the rover movement on the lunar surface

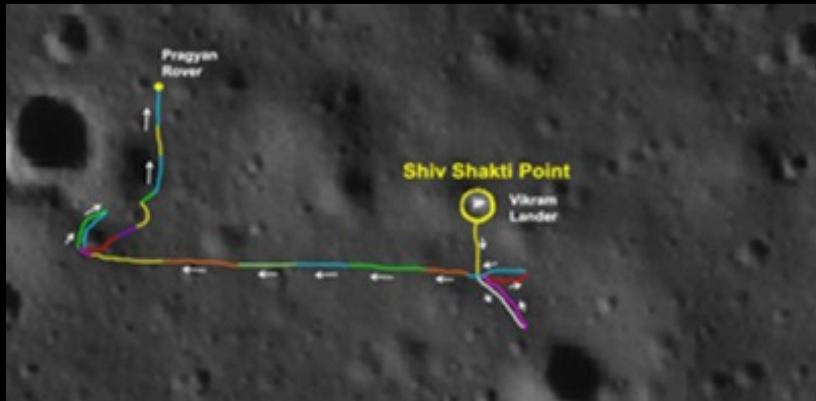
## Extended Mission Accomplishments

- Demonstration of hopping of the lander by re-firing the engines – Sep 3, 2023
- Return of the propulsion module (PM) to the Earth's orbit

Aug 23, 2023: Ch-3  
soft-landed on Moon



Vikram Lander on Lunar Surface:  
Picture taken by the Pragyan Rover



Total distance travelled by Pragyan: ~ 101 m

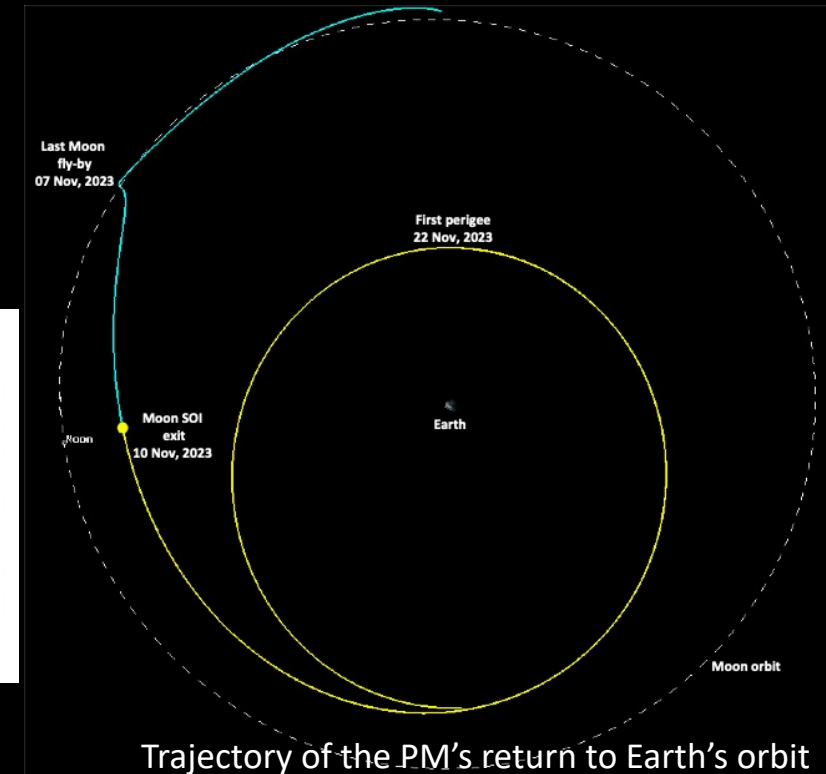
Ramp Deployed Position Captured on  
25-08-2023



Ramp Deployed Position Captured on  
03-09-2023 after post Hopping



Pre and Post hop ramp images captured by LI-1 cam



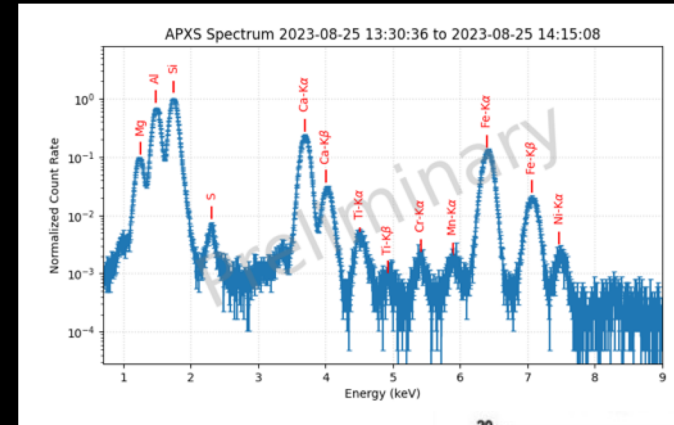
Trajectory of the PM's return to Earth's orbit

Preliminary steps towards Lunar Sample Return Mission

# First-Cut Observations from Chandrayaan-3

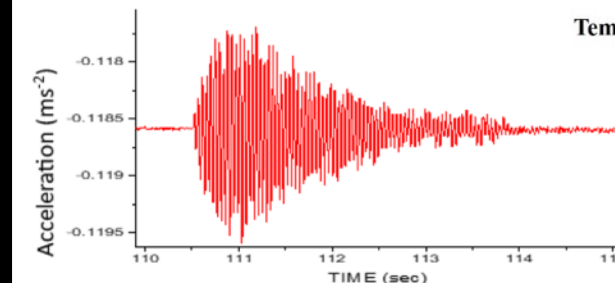
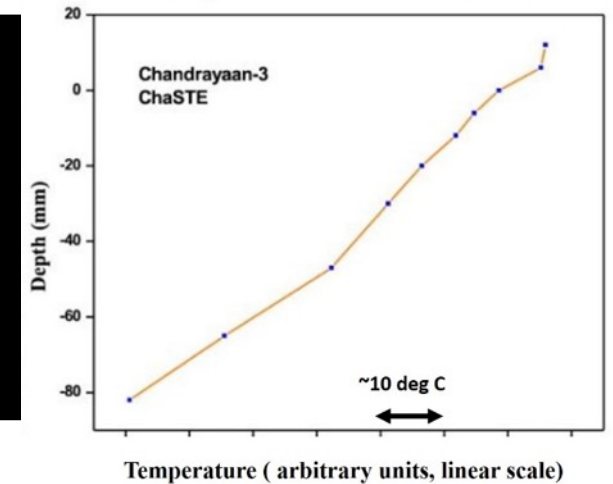
1. First-ever detection of S on the lunar regolith, along with trace constituents like C, N, P, Ti, Mn, Cr, Ni; quantification in progress
2. First-ever temperature profiling of the lunar regolith up to  $\sim 10$  cm depth, results show good thermal insulating properties of the lunar soil
3. A few events of ground vibrations of the lunar surface are recorded
4. First-ever characterization of the near-surface lunar plasma at higher lunar latitude; indicate that only a few tens to hundreds of electrons per cc; dependence on local time

(detailed analyses in progress)



Elemental composition

Temperature profile



Vibration event

**Thank You for Kind Attention**