

17th Meeting of the International Committee on Global Navigation Satellite Systems



Data and Applications of Space Weather Payloads onboard of BDS Satellites

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1. Space environment disruptions near Earth, caused by solar activities





2. Possible consequences of space weather events to GNSS





3. Incident Electrons and Charging Events





The Solar Cycle 25 is HERE









SILSO graphics (http://sidc.be/silso) Royal Observatory of Belgium 2023 October 1



The Solar Cycle 25 is Here

2. Solar activity-Geomagnetic storm-Increasing atmospheric density-Satellites Decay





The Solar Cycle 25 is Here





BDS









The Solar Cycle is Here

4. Increasingly active solar activity









1. Energetic Electron Detection Packages

Payload	Characteristic Parameter	Function
Medium-energy Electron Spectrometer (MES)	Energy: $50 \sim 600 \text{keV}$ FOV: $30^{\circ} \times 180^{\circ}$ Geometric factor: $< \sim 2.0 \times 10^{-3}$	Measure the energy spectra and flux changes of medium electrons in the outer radiation belt.
High-energy Electron Detector (HED)	Energy: 0.5~3.0MeV FOV: 30° cone-angle Geometric factor: <~1.0×10 ⁻²	Measure the energy spectra and flux changes of high electrons in the outer radiation belt.
Deep Dielectric Charging Monitor (DDCM)	Charging Voltage: -2.5 kV to 0 V Charging Current: 0.01-50 pA	Measure the deep dielectric charging current and voltage.





2. Space Plasma and Satellite Surface Charging Monitor

Payload	Characteristic Parameter	Function
Low Energy Electron/Ion Spectrometer	Energy: 0.1~15 keV FOV: 2π Resolution: <15%±2%	Detect parameters of in-situ electrons and ions, such as energy, flux, density and velocity.
Magnetometer	Range: -65000 nT ~ +65000 nT Noise: 10 nT	Measure the environmental magnetic field around the satellites.
Radiation dosimeter	Radiation dosage: $0 \sim 10^7$ rad	Measure total radiation dose to evaluate the lifetime of satellite.
Surface potential detector	Surface potential : 0.1 ~ 10 kV	Monitor the satellite's surface potential.



3. Data of HED (High-energy Electron Detector)

Highly Energy Electron Flux (X: 0.2-0.3MeV)



The red arrows indicate times of four CME (Coronal Mass Ejection) events



4. Data of LEIS (Low Energy Ion Spectrometer)



C: Characteristic for injection of charged particles



4. Analysis of the Surface Charging Events

- Spectrum data of both lons and Electrons are from the same satellite
- The higher electron fluxes appear from the midnight to the dawn side, which may indicate that the electron injection from the magnetotail mainly occurs at the dawn side
- The weakening of the substorm and the weakening of the electron flux observed by the electron detector on the dawn side may indicate a correlation in-betweens



5. A prediction model for high-energy electrons in radiation belt

Model Principle: Delayed correlation between intermediate and high-energy electrons



$$\begin{aligned} \frac{dj}{dt} &= S + L \\ j(t + \Delta t)_{\text{high}} &= j(t)_{\text{high}} + \Delta t(S + L) \\ S &= C_1 C_2^{\left(\frac{j(t-t_0)_{\text{low}}}{j_0}\right)^b} + C_3 j(t - \Delta t)_{\text{high}} \\ L &= -C_4 j(t)_{\text{high}} \\ g &= \log_{10} \left[j(t)_{\text{high}} \right] - \log_{10} \left[j(t - \Delta t)_{\text{high}} \right] \\ \log_{10} \left[j(t + \Delta t)_{\text{high}} \right] &= \log_{10} \left[j(t + \Delta t)_{\text{high}} \right] (1 + fg) \end{aligned}$$

> Model Equations







Model Results : The prediction efficiency and accuracy are both very good









Proposal and Discussions

• Joint response to catastrophic space weather events

- > We should work together to prevent possible damages from severe space weather events to our GNSS satellites
- The workshops for space weather events could be conducted in ICG
- Data products, models and forecast outputs could be shared and discussed in ICG.
- A platform, such as a website, could be built for public to study the impact for space weather events to GNSS



Thank you

http://en.beidou.gov.cn

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