PRESTO - Predictability of the variable solar-terrestrial coupling The new SCOSTEP's scientific program in 2020-2024

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SCOSTEP

Scientific Committee on Solar-Terrestrial Physics

- > one of the Thematic Organizations of the International Science Council (ISC).
- permanent observer at the United Nations Committee on the Peaceful Uses of Outer Space (UNCOPUOS).
- Runs long-term (4-5 years) international interdisciplinary scientific programs of solar terrestrial physics since 1966
- Interacts with national and international programs involving solar terrestrial physics elements
- Engages in Capacity Building activities such as via annual Space Science Schools
- Disseminates new knowledge on the Sun-Earth System and how the Sun affects life and society as outreach activities









PRESTO: Predictability of the variable Solar Predictability of the variable Solar Terrestrial Coupling

The SCOSTEP's new Scientific Program



International interdisciplinary programs in solar-terrestrial physics operated by SCOSTEP

1976-1979: IMS (International Magnetosphere Study) 1979-1981: SMY (Solar Maximum Year) **1982-1985: MAP (Middle Atmosphere Program)** 1990-1997: **STEP** (Solar-Terrestrial Energy Program) 1998-2002: Post-STEP (S-RAMP, PSMOS, EPIC, and ISCS) 2004-2008: CAWSES (Climate and Weather of the Sun-Earth System) 2009-2013: CAWSES-II (Climate and Weather of the Sun-Earth System-II) 2014-2018: VarSITI (Variability of the Sun and Its Terrestrial Impact) **2020-2024:** PRESTO (Predictability of the variable Solar-Terrestrial Coupling)

SCOSTEP's Next Scientific Program (NSP) committee

In October 2017, the SCOSTEP Bureau established the NSP committee with the following members:

Ioannis A. Daglis, Chair (National and Kapodistrian University of Athens, Greece) Loren Chang (National Central University, Taiwan) Sergio Dasso (Universidad de Buenos Aires, Argentina) **Olga Khabarova (IZMIRAN, Russia) Emilia Kilpua** (University of Helsinki, Finland) **Daniel Marsh** (National Center for Atmospheric Research, USA) Katja Matthes (GEOMAR and Christian-Albrechts Universität, Germany) **Dibyendu Nandi (IISER Kolkata, India)** Annika Seppälä (University of Otago, New Zealand) **Rémi Thiéblemont** (Univ. Pierre et Marie Curie, France) **Qiugang Zong (Peking University, China)**

SCOSTEP Next Scientific Program (NSP) Forum at ISSI-Beijing, China, 2018

Detailed documentation is available at: http://www.issibj.ac.cn/Publications/Forum_Reports/201404/W020190620592906717714.pdf

- Amal Chandran, Nanyang Technological University, Singapore
- Ioannis A. Daglis, NSP Committee Chair / National and Kapodistrian University of Athens, Greece
- Sergio Dasso, NSP Committee Member / IAFFE, Argentina
- Katya Georgieva, VarSITI / Bulgarian Academy of Sciences
- Nat Gopalswamy, SCOSTEP Bureau / NASA GSFC, USA
- Mamoru Ishii, NICT, Japan
- Olga Khabarova, NSP Committee Member / IZMIRAN, Russia
- Kanya Kusano, Nagoya University, Japan
- William Liu, National Space Science Center, CAS, China
- Shinobu Machida, Nagoya University, Japan
- Takahiro Obara, Tohoku University, Japan



Figure 6: Group picture of the participants of the Forum in 2018.

SCOSTEP Next Scientific Program (NSP) Forum at ISSI-Bern, Switzerland, 2019

- Seth Claudepierre, The Aerospace Corporation and UCLA, USA
- Ioannis A. Daglis, NSP Committee Chair / National and Kapodistrian University of Athens, Greece
- Patricia Doherty, Boston College, USA
- Katya Georgieva, VarSITI / Bulgarian Academy of Sciences
- Nat Gopalswamy, SCOSTEP Bureau / NASA GSFC, USA
- Olga Khabarova, NSP Committee Member / IZMIRAN, Russia
- Emilia Kilpua, NSP Committee Member / University of Helsinki, Finland
- Petra Koucká Knížová, Czech Academy of Sciences
- Vladimir Kuznetsov, SCOSTEP Bureau / IZMIRAN, Russia
- Franz-Josef Luebken, SCOSTEP Bureau / IAP, Germany
- Daniel Marsh, NSP Committee Member / NCAR, USA
- Dibyendu Nandi, NSP Committee Member / IISER Kolkata, India
- Nick Pedatella, UCAR, USA
- Eugene Rozanov, PMOD and ETH, Switzerland
- Marianna Shepherd, SCOSTEP Bureau / York University, Canada Figure 7: Group picture of the participants of the Forum in 2019.
- Kazuo Shiokawa, VarSITI / Nagoya University, Japan
- Alphonse Sterling, NASA MSFC, USA
- Manuela Temmer, University of Graz, Austria
- Rémi Thiéblemont, NSP Committee Member / IPSL, France
- Qiugang Zong, NSP Committee Member / Peking University, China

Detailed documentation is available at: http://www.issibj.ac.cn/Publicatio ns/Forum_Reports/201404/W0201 90620592906717714.pdf



An integrated view of solar-terrestrial prediction

Overlap of various Solar-Terrestrial phenomena with various spatial & temporal scales



Detailed documentation is available at:

http://www.issibj.ac.cn/Publications/Forum_Reports/201404/W020190620592906717714.pdf





SCOSTEP/PRESTO Predictability of the Solar-Terrestrial Coupling

PRESTO is a science program that seeks to improve the predictability of energy flow in the integrated Sun-Earth system on times scales from a few hours to centuries through promoting international collaborative efforts. PRESTO is sponsored by SCOSTEP, the Scientific Committee on Solar Terrestrial Physics.



Chair: Ramon Lopez



Co-chair: Jie Zhang



Co-chair: Eugene Rozanov

PRESTO is comprised of 3 pillars:

- Pillar 1: Sun, Interplanetary Space and Geospace
- Pillar 2: Space Weather and the Earth's Atmosphere
- Pillar 3: Solar Activity and Its Influence on the Climate of the Earth System.

The PRESTO Science Program is featured in the SCOSTEP/PRESTO Newsletter Volume 22



https://scostep.org/newsletter-archive/





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Pillar 1: Sun, Interplanetary Space and Geospace

Heliospheric transients originating from the Sun, stream interaction regions (SIRs), fast streams, coronal mass ejections, as well as Solar Energetic Particles (SEPs) are key agents driving disturbances in geospace. Their formation and evolution, solar wind – magnetospheric coupling and ensuing magnetospheric dynamics form a complex chain.

Predicting accurately and reliably various geospace disturbances, including changes in near-Earth plasma waves and radiation environment, requires understanding the key aspects of related phenomena, physical processes and their interplay from Sun to Earth operating over timescales ranging from milliseconds to days.

Pillar 1: Co-Leaders



Allyson Jaynes



Emilia Kilpua



Spiros Patsourakas





Space weather meets Earth's atmosphere

Pillar 2: Space Weather and Earth's Atmosphere

The thermosphere and the ionosphere are conventionally known to be directly affected by the varying magnitudes of solar activity.

When incidence of solar flux or energetic charged particles increases suddenly due to the occurrence of solar flares or coronal mass ejections combined with magnetic reconnection in the Earth's magnetosphere, the adverse effects witnessed in the ionosphere-thermosphere (IT) system as a whole are referred to as Space Weather effects.

Consequently, these space weather effects lead to adverse effects in several space and ground-based applications as illustrated in the figure above. These effects and consequences are to be understood at both fundamental and applications levels.



Pillar 2: Co-Leaders

Loren C. Chang



Duggirala Pallamraju



Nick Pedatella



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Pillar 3: Solar Activity and its Influence on Climate

The next 5 years spanning the start of Solar Cycle 25 through its (near) peak provide an excellent opportunity for evaluating our understanding and ability to predict solar activity on decadal time scales.

Predictability requires improved understanding of the physical pathways wherein solar variability impacts the atmosphere, from the magnetosphere through the troposphere. The nonlinear and stochastic mechanisms which modulate the solar cycle and affect the scope of the prediction and predictability of the solar cycle are still open questions.

In some areas, decadal-scale solar forcing remains to be quantified to climate-relevant accuracies, which challenge our ability to determine causal connections in the pathways explaining solar forcing impacts on climate.

Furthermore, due to wave-driven coupling in the atmosphere, improving predictability of solar forcing also requires improved characterization of the atmosphere- ocean response to the forcing.

Pillar 3: Co-Leaders



Odele Coddington



Jie Jiang



Stergios Misios



To get involved in PRESTO, contact the chairs or relevant pillar leaders.

Join the SCOSTEP Mailing list. Send a note of introduction to <u>scostep@bc.org</u> with your request to join the mailing list.

SCOSTEP/PRESTO Opportunities

SCOSTEP/PRESTO provides support for organizing international campaigns and meetings. The activities should be strictly related to one or more PRESTO Pillars and contribute to PRESTO activities.

Funding for campaigns and meetings is up to \$5000. The deadline for 2021 proposals is extended to December 31, 2020.

Go to the following website for guidelines on these proposals: https://scostep.org/grant-proposals/

To get involved in PRESTO, contact the chairs and /or relevant pillar leaders.

Watch for upcoming online seminars.

Join the SCOSTEP Mailing list.

For all information: https://scostep.org/presto/



For subscription on the SCOSTEP-all mailing list: drop e-mail to "scosteprequest@bc.edu".

PRESTO Website at Boston College: www.bc.edu/scostep/

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SCOSTEP: PRESTO

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PRESTO

SCOSTEP is tasked to organize organize long-term scientific programs in solar terrestrial physics. The program for 2020-2024 is **Pre**dictability of Variable **S**olar-**T**errestrial Coupling (PRESTO). PRESTO was defined based on a community effort with numerous white papers submitted to the Committee for Definition of the Next Scientific Program (NSP) and two fora organized by the International Space Science Institute (ISSI) in Beijing in 2018 and in Bern in February 2019. PRESTO's goals are to "address the predictability of 1) space weather on timescales from seconds to days and months, including processes at the Sun, in the heliosphere and in the Earth's magnetosphere, ionosphere and atmosphere, [and] 2) sub-seasonal to decadal and centennial variability of the Sun-Earth system, with a special focus on climate impacts and a link to the World Climate Research Program (WCRP) Grand Challenge Near-Term Climate Predictions as well as the IPCC." (NSP Concept Text).

PRESTO Logo Contest Winner Announced

We are pleased to share the results of the PRESTO logo contest. A total of 37 clever designs were submitted for consideration. After much thought, the PRESTO committee selected the design provided by Mr. Vishal Jagatsing Pawar. Mr. Pawar is a Master's Degree Candidate in Atmospheric and Space Science at Pune University in Maharashtra, India.

Congratulations to Mr. Vishal Pawar and thanks to all who submitted a design.

- SCOSTEP/PRESTO Grants to support meetings and campaigns
- SCOSTEP capacity building grant to support schools and capacity building activities

SCOSTEP-PRESTO ONLINE SEMINAR SERIES

1st SCOSTEP/PRESTO Online Seminar Title: A challenge to Physics-based Prediction of Giant Solar Flares Author: Kanya Kusano (Institute for Space-Earth Environmental Research (ISEE), Nagoya University, Japan) Date/time: May 26 (Tue), 2020, 12:00-13:00 UT 2nd SCOSTEP/PRESTO Online Seminar Title: Extreme solar events: A new paradigm Author: Ilya Usoskin (University of Oulu, Finland) Date/time: July 20 (Mon), 2020, 12:00-13:00 UT **3rd SCOSTEP/PRESTO Online Seminar** Title: Developing a Highliy Predictable Geomagnetic Index to Gauge Magnetospheric Activity and Space Weather Author: Joe Borovsky (Space Science Institute, USA) Date: September 10, 2020, 22:00-23:00 UT 4th SCOSTEP/PRESTO Online Seminar **Title:** The lonospheric Connection Explorer - Results from the first year on orbit Author: Thomas Immel (University of California Berkeley, USA) Date: November 17, 2020, 23:00-24:00 UT **5th SCOSTEP/PRESTO Online Seminar** Title: Magnetospheric Response to Interplanetary Shocks: ULF Wave-Particle Interaction Perspective Author: Q.-G. Zong (Peking University, China) Date and Time: Jan 14 (Thu), 2021, 00:00-01:00 UT 6th SCOSTEP/PRESTO Online Seminar Title: Utilizing galactic cosmic rays as signatures of interplanetary transients Author: Mateja Dumbović (University of Zagreb, Croatia) Date and Time: Jan 19 (Tue), 2021, 12:00-13:00 UT



SCOSTEP/PRESTO Newsletter vol.21-26

Articles, Highlight of young scientists, Meeting reports, and Short news



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Dragan Rosa

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SSISTEP

Inside this issue

Highlight on Young Scientists 1: Naoko Takahashi Japan

Highlight on Young Scientists 2: Xianwen Jing USA / China

Upcoming Meetings

Announcement 1: Climate Implications of the Sun transition to high Activity mode

Announcement 2: PRESTO Town Hall a AGU Fall Meeting

Announcement 3: Memorandum of Un-derstanding between SCOSTEP and ISEE, Nadrowa University

Article 1 How small-scale cur-rent sheets and mag-netic islands in the so-lar wind help under-standing the nature of large-scale processes behind space weather SCOSTEP/PRESTO

IEWSLETTER

Vol. 26, January 2021

Space Weather

Pushkov Institute of Terre ropertation of the Russian

opegation of the Nussee

Olga Khabarova^{1,2}

How Small-Scale Current Sheets and Magnetic

*Space Research Institute of the Russian Academy of Sciences (IKI), Moscow 117997, Russia.

Islands in the Solar Wind Help Understanding the Nature of Large-Scale Processes Behind

Current heters (155) occur in the solar (HCS) and CSs formed within high-speed Ownid ubiquitoxity, owing to both flowvitemum, to several proton grounds local dynamical processes and dpuble in the case of CSs involved in the trans-mitipole nature of the solar magnetic leat catcade. The within of CSs is about field. Their trycical domains that ranges several proton grounds independently of from many AUs, e.g. in the case of the their origin, therefore, in this term, CSs quant-stitle helpolyneic current shear new year indicates growthere.

instanced education wave index html (Khabarova et al. 2015, 2016). d) EVLLI er structions of the HCS (white line) distortion by propagating streams (green and low areas), see http://www.deloweather.net/ for details. In the meridional plane, seen that the HCS is crossed twice - before and after the stream passage. Magg end that the HCS is crossed twice - before and after the stream passage. Magg

are formed between the stream and the HCS. e) Sketch illustrating a com tortion of the HCS/HPS-magnetic island system by an approaching CIR is islands within the HPS are grey and outside are yellow. Since the HCS/HP when headed by the CIR, some of magnetic islands are located inside the rippl

Olga

the solar wind speed from http al 2016 2016 d) ENLI



PRESTO Officers

PRESTO chair/ co-chairs



Pillar 1: Sun, Interplanetary space and geospace



USA Finland

Pillar 2: Space weather and the Earth's atmosphere



co-leader Loren Chang Talwan



Greece

Pillar 3: Solar activity and its influence on the climate of the Earth System



co-leader Odele Coddington Jie Jiang China USA

co-leader Stergios Misios Greece

PRESTO (PREdictability of the variable Solar-Terrestrial cOupling) is a science program that seeks to improve the predictability of energy flow in the integrated Sun-Earth system on times scales from a few hours to centuries by promoting international collaborative efforts. PRESTO is sponsored by SCOSTEP, the Scientific Committee on Solar Terrestrial Physics.

The Sun is a variable star and its variability influences the Earth's space environment. Furthermore, changing solar magnetic fields, radiative and energetic particle fluxes force the Earth's atmosphere and climate. Transient energetic events such as flares, coronal mass ejections (CMEs), interplanetary shocks, stream interaction regions (SIRs), corotating interaction regions (CIRs) and energetic particles adversely impact critical technologies based in space and on Earth that our society is increasingly dependent upon. At the same time, the middle and upper atmosphere/ ionosphere are impacted by processes originating at lower altitudes, e.g., by atmospheric gravity waves, tides and planetary waves and changes in radiatively active gases. Solar influence on climate is gaining increasing attention since variations in solar activity do not only impact middle atmosphere chemistry and physics, but also impact decadal variability at the Earth's surface. This is of particular interest and importance for decadal climate predictions. With the enhanced understanding of causal connections in the Sun- Earth system over the last several decades, fueled by both observations and theoretical modelling, we are in a position to transform this understanding to improved predictions of the Sun-Earth coupled system, which is of relevance to the society and the focus of the current PRESTO program.

International interdisciplinary programs conducted by SCOSTEP

1976-1979: IMS (International Magnetospheric Study) 1979-1981: SMY (Solar Maximum Year) 1982-1985: MAP (Middle Atmosphere Program) 1990-1997: STEP (Solar-Terrestrial Energy Program) 1998-2002: SRAMP (STEP-Results, Applications, & Modeling Phase) : PSMOS (Planetary Scale Mesopause Observing System) : ISCS (International Solar Cycle Study) : EPIC (Equatorial Processes Including Coupling) 2004-2008: CAWSES (Climate and Weather of the Sun-Earth System) 2009-2013: CAWSES II 2014-2018: VarSITI (Variability of the Sun and Its Terrestrial Impact) 2020-2024: PRESTO (Predictability of the Variable Solar-Terrestrial Coupling)

PRESTO Leaflet



PRESTO

Predictability of the Variable Solar-Terrestrial Coupling

2020-2024





https://scostep.org

3 Pillars of PRESTO and their science questions

PRESTO Leaflet



Pillar 1: Sun, interplanetary space and geospace

Question 1.1:

Under what conditions are solar eruptions, CMEs, and SEPs produced, and which indicators of pre-CME and pre-flare activity are reliable?

Question 1.2:

What are the required/critical model input parameters for most successfully forecasting the arrival of SEPs and the geoeffectiveness of CMEs, SIRs/CIRs and the consequences of the interactions between SIRs/ CIRs and CMEs?

Question 1.3:

How are different magnetospheric disturbances and waves (which are critical for the ring current and radiation belt dynamics) driven by variable solar wind structures, and/or internal magnetospheric processes?



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Pillar 2: Space weather and Earth's atmosphere

Question 2.1:

How does the thermosphere and ionosphere respond to various forcings from above and from below?

Question 2.2:

How do atmospheric waves and composition changes impact the middle and upper atmosphere?

Question 2.3:

What is the magnitude and spectral characteristics of solar and magnetospheric forcing, needed for accurate predictions of the atmospheric response?

Question 2.4:

What is the chemical and dynamical response of the middle atmosphere to solar and magnetospheric forcing?

Pillar 3: Solar activity and its influence on climate

Question 3.1:

How will future solar activity vary over different timescales and what are the physical reasons for the variations?

Question 3.2:

How will the solar forcing on the Earth's system evolve in the future?

Question 3.3:

What is the role of the coupling between atmospheric regions in the realization of the long- and short-term solar influence on the Earth system and how are those responses affected by increasing green-house gases?



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