Panel on Observing Infrastructure

South American Observing Infrastructure - Current status and the known or potential gaps in the area. Recommendations for future work. Suggestions for local priority needs for the global space weather enterprise.

J.E.R.Costa - INPE – National Institute for Space Science EMBRACE – Regional Warning Center in Brazil

The United Nations/United States of America Workshop on the International Space Weather Initiative: The Decade after the International Heliophysical Year 2007, 31 July – 4 August, 2017

Introduction

First I will address the Space Weather issues as a matter of an operational center:

- It needs cooperation with real time capability (most of time).
- It will be restricted to products being used for monitoring/ forecasting space weather.
- The networks of sensors today in Latin America is mostly driven by science. They are now being used by SW operations through cooperation with existing projects.
- We have now in Latin America two Regional Warning Center (Brazil and Mexico) and we are talking to Argentina to organize their AWC (associate).
- Gaps in the network are being used for planning the investment of EMBRACE program or through international initiatives.
- I will present some conclusions and offer suggestions to overcome some of the main difficulties.



Clezio Marcos Denardini, Sergio Dasso and J. Americo Gonzalez-Esparza

| Group | Name(Acronym) | Origin | Objective of Investigation | Current Status |
|-----------|------------------------|------------|--|------------------|
| · | | (Year) | (Primary Objective) | (FO, PO, NO*) |
| Atm. | Embrace GlowNet | 1998 | Gravity Waves (85 km) and Plasma Bubbles | PO |
| | Airglow imagers | | (260 km) footprints signature and propagations | |
| Atm. | AAS | 1998 | Airglow in mesopause region | FO |
| Âtm. | SARINET | 1999 | D-Region density variations and solar flares | FO |
| | | | related effects | |
| Atm. | SAVNET | 2006 | D-Region (daytime) and E-Region (nighttime) | FO |
| | | | density variations and solar flares related | |
| | | | effects | |
| Atm. | NDMC | 1998 | Mesopause region and changing climate | FO |
| Atm. | RIOM-SMN | 2012 | Absorption in ionosphere | PO |
| Atm. | Schumann station | 2014 | Schumann resonance frequencies | FO |
| Atm. | Embrace DigiNet | 1990 | Bottom-side ionospheric profile measurements | PO |
| | | | and upper-side ionospheric profile theoretical | |
| | lonosondes | | extrapolations, Maximum Usable Frequency | |
| | | | from oblique HF ionospheric reflections | |
| Atm. | RBMC | 1996 | Total ionized atmosphere content and | FO |
| | | | positioning for several purposes like | |
| | GNSS sensors | | geodynamics studies of the South American | |
| | | | plate | |
| Atm. | Calibra | 2012 | Total ionized atmosphere content and | PO |
| (| | | positioning for several purposes like | |
| | Magnetometers | | topographical and geodetic surveys, land | |
| | | | management and offshore operations | |
| Mag. | Embrace MagNet | 2011 | Geomagnetic field variations and geomagnetic | PO |
| | | | storm detections | |
| Source: D | enardini, et.al (2016) | : Review o | n space weather in Latin America-2, Adv | /SR, 58, p.1940. |

| Magnetometersto 2012geomagnetic storm detections, long term secular changes and contribute to IGRFMag.Intermagnet in Argentina1991Geomagnetic field (absolute, variations), geomagnetic storm detections, long term secular changes and contribute to IGRFMag.SMN-LAS1961Geomagnetic Storm detections, long term secular changes and contribute to IGRFMag.SMN-LAS1961Geomagnetic Storms. South American Magnetic Anomaly.FO1. M.GMDN in Brazil2006Cosmic ray, ICMEs trough Forbush decreaseFOI. M.SciCRT-GMDN in Mexico2011Cosmic ray, ICMEs trough Forbush decreaseFOI. M.LAGO2011Cosmic ray, ICMEs trough Forbush decreaseFOI. M.LAGO2011Cosmic ray, ICMEs trough Forbush decreaseFOI. M.Auger - Low energy modes2005Cosmic ray, ICMEs trough Forbush decreaseFOI. M.Cosmic Ray Observatory1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.MEXART tracking of solar wind disturbances using the interplanetary scintillationFOSunSolar Submillimeter Solar Submillimeter1999Active regionFOSunCallisto-MEXART Solar Neutron Telescope2005Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFONetworkNetworkNetworkSolar energetic particlesFO |
|--|
| Magnettonneters 2012 secular changes and contribute to IGRF Mag. Intermagnet in Argentina 1991 Geomagnetic field (absolute, variations), geomagnetic storm detections, long term secular changes and contribute to IGRF Mag. SMN-LAS 1961 Geomagnetic Storm detections, long term secular changes and contribute to IGRF Mag. SMN-LAS 1961 Geomagnetic Storms. South American Memory FO I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy modes 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO I. M. Solar Submillimeter 1999 Active region FO Sun Solar Submi |
| Mag. Intermagnet in Argentina 1991 Geomagnetic field (absolute, variations), geomagnetic storm detections, long term secular changes and contribute to IGRF FO Mag. SMN-LAS 1961 Geomagnetic Storms. South American FO FO Mag. SMN-LAS 1961 Geomagnetic Storms. South American FO FO I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary Scintillation FO Interplanetary Scintillation tracking of solar wind disturbances using the interplanetary scintillation technique FO Sun Solar Submillimeter 1999 Active region FO |
| Mag. SMN-LAS 1961 Geomagnetic storm detections, long term secular changes and contribute to IGRF Mag. SMN-LAS 1961 Geomagnetic Storms. South American Magnetic Anomaly. FO I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy modes 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO Interplanetary Scintillation interplanetary scintillation FO FO Sun Solar Submillimeter 1999 Active region FO Sun Solar Neutron Telescope 2004 Sola |
| Mag. SMN-LAS 1961 Geomagnetic Storms. South American Magnetic Anomalv. FO I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Maxico Cosmic Rays Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy modes 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO Telescope Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope |
| Mag. SMN-LAS 1961 Geomagnetic Storms. South American Magnetic Anomaly. FO I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in Mexico 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particles FO Network Network FO FO FO |
| I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Image: Cosmic Rays Image: Cosmic Rays FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO Interplanetary Scintillation interplanetary scintillation technique FO Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particl |
| I. M. GMDN in Brazil 2006 Cosmic ray, ICMEs trough Forbush decrease FO I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays Image: Cosmic Rays Image: Cosmic Rays FO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Auger - Low energy 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary scintillation FO Interplanetary Scintillation interplanetary scintillation technique FO Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particles FO Network FO FO FO FO FO |
| I. M. SciCRT-GMDN in 2014 Cosmic ray, ICMEs trough Forbush decrease FO Mexico Cosmic Rays 2011 Cosmic ray, ICMEs trough Forbush decrease PO I. M. LAGO 2011 Cosmic ray, ICMEs trough Forbush decrease PO I. M. Auger - Low energy 2005 Cosmic ray, ICMEs trough Forbush decrease FO modes I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary Scintillation FO Interplanetary Scintillation interplanetary scintillation technique FO Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particles FO Network Extended Extended FO FO |
| MexicoCosmic RaysI. M.LAGO2011Cosmic ray, ICMEs trough Forbush decreasePOI. M.Auger - Low energy modes2005Cosmic ray, ICMEs trough Forbush decreaseFOI. M.Cosmic Ray Observatory modes1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.Cosmic Ray Observatory modes1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.MEXART metary Scintillation2005tracking of solar wind disturbances using the interplanetary scintillation techniqueFOSunSolar Submillimeter Telescope1999Active regionFOSunCallisto-MEXART Solar Neutron Telescope2004Solar energetic particlesFONetworkFOFO |
| I. M.LAGO2011Cosmic ray, ICMEs trough Forbush decreasePOI. M.Auger - Low energy modes2005Cosmic ray, ICMEs trough Forbush decreaseFOI. M.Cosmic Ray Observatory1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.MEXART2005tracking of solar wind disturbances using the interplanetary scintillationFOSunSolar Submillimeter Telescope1999Active regionFOSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFO |
| I. M. Auger - Low energy modes 2005 Cosmic ray, ICMEs trough Forbush decrease FO I. M. Cosmic Ray Observatory 1999 Cosmic ray, ICMEs trough Forbush decrease FO I. M. MEXART 2005 tracking of solar wind disturbances using the interplanetary Scintillation FO Sun Solar Submillimeter 1999 Active region FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particles FO |
| modesI. M.Cosmic Ray Observatory1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.MEXART2005tracking of solar wind disturbances using the interplanetary scintillationFOSunSolar Submillimeter Telescope1999Active regionFOSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFO |
| I. M.Cosmic Ray Observatory1999Cosmic ray, ICMEs trough Forbush decreaseFOI. M.MEXART2005tracking of solar wind disturbances using the interplanetary scintillationFOSunSolar Submillimeter1999Active regionFOSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFO |
| I. M.MEXART2005 Interplanetary Scintillationtracking of solar wind disturbances using the interplanetary scintillation techniqueFOSunSolar Submillimeter1999Active regionFOTelescope5050FOFOSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFONetworkFOFO |
| Sun Solar Submillimeter 1999 Active region FO Telescope FO FO FO Sun Callisto-MEXART 2015 Solar radio burst FO Sun Solar Neutron Telescope 2004 Solar energetic particles FO Network FO FO FO |
| SunSolar Submillimeter1999Active regionFOTelescopeTelescopeTelescopeFOSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFONetworkNetworkFOFO |
| TelescopeSunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFONetworkNetworkFOFO |
| SunCallisto-MEXART2015Solar radio burstFOSunSolar Neutron Telescope2004Solar energetic particlesFONetwork |
| Sun Solar Neutron Telescope 2004 Solar energetic particles FO Network Network FO |
| Network |
| |
| SunItapetinga Radio1974Galactic and extra-galactic studies, solarFO |
| physics specially the solar continuous emission |
| solar observatories and transition region monitoring |
| SunBrazilian Decimetric2004Galactic and extra-galactic studies, highPO |
| Array resolution solar physics |
| Sun Northeast Radio Space 1993 Geodetic studies as well as galactic and extra- FU |
| Observatory galactic studies. |

* FO = Fully operational, PO = Partially Operational, NO = Not Operational

Source: Denardini, et.al (2016): Review on space weather in Latin America-2, AdvSR, 58, p.1940.

GNSS Network for determinations of TEC

Net: RBMC (RBMC -Brazilian Network for Continuous Monitoring of the GNSS Systems.)

- The EMBRACE program does not own all the instruments used to complete the South America TEC map (now being extended to Latin America).
- A successful program of national and international cooperation was set to grant a denser GNSS network.
- In this example the Brazilian RBMC net.



GNSS Network for determinations of TEC Net: LISN – Boston College, USA



GNSS Network for determinations of TEC Net: RAMSAC - Argentina



GNSS Network for determinations of TEC Net: IGS - International



GNSS Network for determinations of TEC Net: ALL NETWORKS

- Even including all sensors it is still a low density array.
- Mainly in the North of Amazon region and some neighbors countries.
- This map does not include Central America.



GNSS Network for determinations of TEC Net: TLALOCNet-Mexico and COCONet-Central

America

- The network is now being extended to whole Latin America including these two Networks.
- The gaps can be seen in the TEC maps.





Magnetometer Network

- A low density network
 Net: EMBRACE
- Although the magnetometers are cross-calibrated some of them are noisy due to electrical devices near.
- As the Magnetic Equator has moved to the West there is no magnetometer under the Equatorial Electro Jet.
- EEJ may amplifies the geomagnetic variations that give origin to the GIC. Need to have more focus on research.



MAGNETOM





Airglow Network

Net: EMBRACE

- Low density network.
- Very important for the plasma bubble monitoring.
- Suffers from adequate sites without light pollution, clear sky and infrastructure.
- Need investment in F-P interferometer to understand better the plasma irregularity called plasma bubble.
- Two F-P were acquired to measure thermospheric winds and temperature (EMBRACE).
- Initial discussion with Dr. Qian Wu -NCAR (2 FPI) to be part of a North_South network.
- Initial discussion with Dra. Hsieh,
 Syau-Yun (John Hopkins) (4 All-Sky) to be deployed in Brazil



etwork

All Sky Network Map



GNSS Network with S4 index Net: EMBRACE



Metwork

- Low density network.
- Very important for monitoring disturbance in radio wave propagation (index S4, others).
- The full network suffers from old GNSS receivers without capability to measure scintillations.

Scintilation Network Map



Ionosonde Network Net: EMBRACE



IONOSONDE:

- Low density network.
- Need lonosondes at the crest of the anomaly.
- Lack of local technical/scientific training to stimulate the partnership in planned strategic sites.

lonosonde Network Map



Other Networks

SCiESMEX-Mexico

A radiotelescope MEXART, an array of 4096 dipoles operating at 140 MHz to track solar wind large-scale disturbances using the interplanetary scintillation technique.



Source: Mejia-Ambriz et.al, Solar Phys (2010) 265: 309–320

EMBRACE-Brazil

A muon detector part of the Global Muon Detector Network –GMPR 13 Shinshu University





1)Nagoya, Japan; 2)Hobart, Australia; 3)City of Kuwait, Kwait; 4)São Martinho da Serra, Brazil The majority of the ISWI instruments in the South America are in the LISN network.

 NSSC -International Meridian Circle Project.

- Ionosonde(1)
- Magnetometer(1)
- GNSS (1)
- All Sky (1)

ISWI Network

ISWI in Brazil (48) AMBER(2), CALLISTO(2), CSSTE(1), GMDN(1), LISN(22), MAGDAS(2), RENOIR(2), SAVNET(6), SCINDA(3), SID(7)



Conclusions

- The cooperation among existent research groups must be within association with an Operational Center for SW (RWC) and follow an operational approach:
 - The Operational Center (e.g. RWC) commit **to monitor the pipeline** of the instrument data to report quickly to the PI of instruments any interruption. Assuring an efficient duty cycle.
 - The Operational Center commit to **organize a data bank and promptly release** the information.
 - The Operational Center may **offer some technical support** for the instrument maintenance when needed (not always possible).
 - Initiatives to deploy instruments such as ISWI should always **use the support of an Operational Center** to be responsible for the instrument operation.
- International initiatives should work hard for the creation of data banks of data from ground and space based instruments in data servers around the world to enable the Operational Centers to collect data directly from the data servers.
- Foment open data police.
- In strategic areas foment local technical/scientific training to stimulate the partnership and assure continuity of the instrument operation. Training through capacity building schools.
- Offer technical support scholarships.