

ESA heliophysics archives: a key asset for ISWI

Arnaud Masson

*ESA heliophysics science lead
Cluster deputy project scientist*

*UN-US workshop on ISWI
Boston College, 4 August 2017*

ESA science archives



<http://archives.esac.esa.int>

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- Publications
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ESAC SCIENCE DATA CENTRE

Astronomy Science Archives

esasky	exosat	gaia
herschel	hubble space telescope	iso
lisa pathfinder	planck	xmm-newton

Heliophysics Science Archives

cluster	double star	ISS-SolACES*
proba-2	soho	ulysses

The Planetary Science Archive

cassini huygens	exomars	giotto
mars express	rosetta	smart-1
venus express		

LATEST NEWS

Tweets by @ESAesdc

ESA ESDC @ESAesdc #ESASky is featured in this month's #INTEGRAL Picture of the Month! tinyurl.com/ybs3yzgz @ESA_Integral sky.esa.int

10h

ESA ESDC @ESAesdc European Commission Horizon 2020 Proposers' Day 2017, 9-10 Nov 2017 Budapest, free admission buff.ly/2h8hyfI

28 Jul 2017

ESA ESDC @ESAesdc Cuando ya no esté: Carlos Barrabés (Parte 1/2) I #0 buff.ly/2vdhyhnE "Si yo tuviese 20 años hoy, me dedicaría al espacio"

European Space Agency

<http://archives.esac.esa.int>

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Future Archives

 bepicolombo	 euclid	 juice
 solar orbiter		

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The ESAC science data centre (ESDC)

1. Consolidation of all ESA Space Science Archives at ESAC, with strong re-use across projects, ensuring easier and cheaper long term data preservation
 - Hardware infrastructure
 - Software architecture and code, including technology migration
 - Human technical and scientific expertise
 - **Multi mission, multi instruments science exploitation**
2. Long term preservation of data processing capabilities (on-going)
 - Preserve software coming from various places
 - Provide data processing capabilities as a “service”
 - Bring the “user software to the data” instead of the “data to the user”
3. Sharing and preservation of knowledge, including international cooperation
 - IVOA, IPDA

Towards multi wavelengths Science data exploitation



- **Goal:** to facilitate data discovery and archival science for ALL users
 - Multi-wavelength
 - Project agnostic
 - Exploration
- Interface “on top of” all ESA astronomy archives



ESASky

Legacy:
IUE, ..



Herschel
XMM-
Newton



HST



Planck



ISO



Integral



EXOSAT

Future:
Gaia,
JWST,
Euclid, ..

All-sky maps for all ESA missions

J2000 15 03 23.406 -42 05 21.11

Sky XMM-Newton EPIC color

Upload target list

Target list file: SNRs_list2.txt

- SN 1006
- SNR G315.0-02.3
- 08 52 00.0 -46 20 00
- 23 01 35.0 +58 53 00
- 05 26 30.0 +42 56 00
- 08 34 00.0 -45 50 00
- 15 53 00.0 -56 10 00
- 16 01 06.0 -51 34 00

Screenshot **Multi-target**

Search...

- Web application
- Simple sky exploration interface

Data Panel INTEGRAL#1 XMM-Newton#1 SUZAKU#1 XMM-OM(UV)#1 HST#1 Gaia DR1 TGAS#1

Imaging Observations

Mission	Filter	# results
INTEGRAL	UV	~100
SUZAKU	UV	~100
XMM-Newton	UV	~100
XMM-OM(UV)	UV	~100
XMM-OM(Optical)	UV	~100
HST	UV	~100
WFC3	UV	~100

Source Catalogue

Mission	Filter	# results
HSC	Visible	~1000000
PCCS2E	Visible	~1000000
PCCS2-HFI	Visible	~1000000
PCCS2-LFI	Visible	~1000000
PSZ	Visible	~1000000
PGCC2	Visible	~1000000

Click on histograms bars to start retrieving metadata.

Quick access to science-ready observations and catalogues

FoV: 1.53°

Close data panel

esa European Space Agency

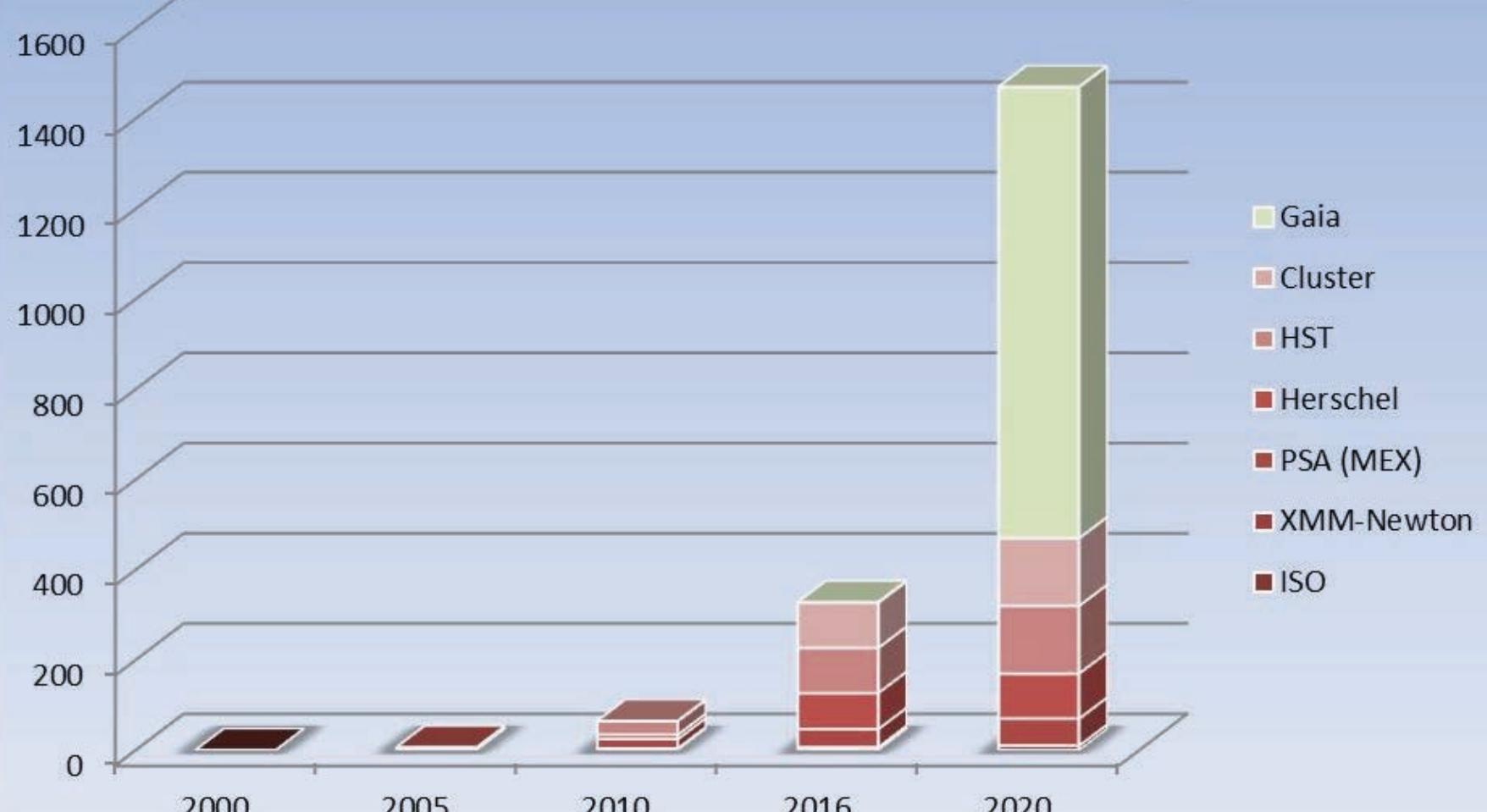
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Archives Volume evolution (2000-2020)

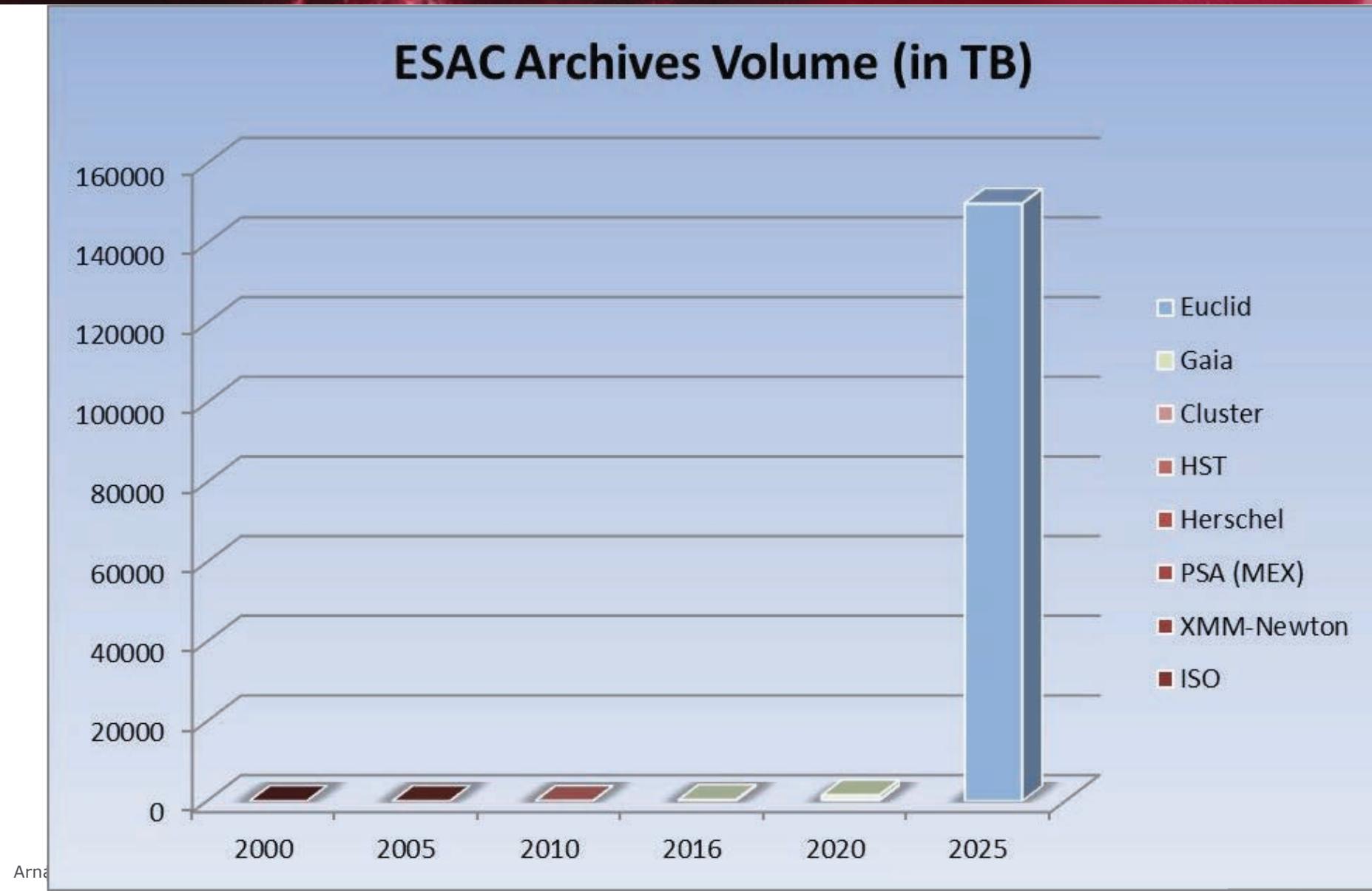


ESAC Archives Volume (in TB)



Arn

Archives Volume evolution (2000-2025)



Arn

<http://archives.esac.esa.int>

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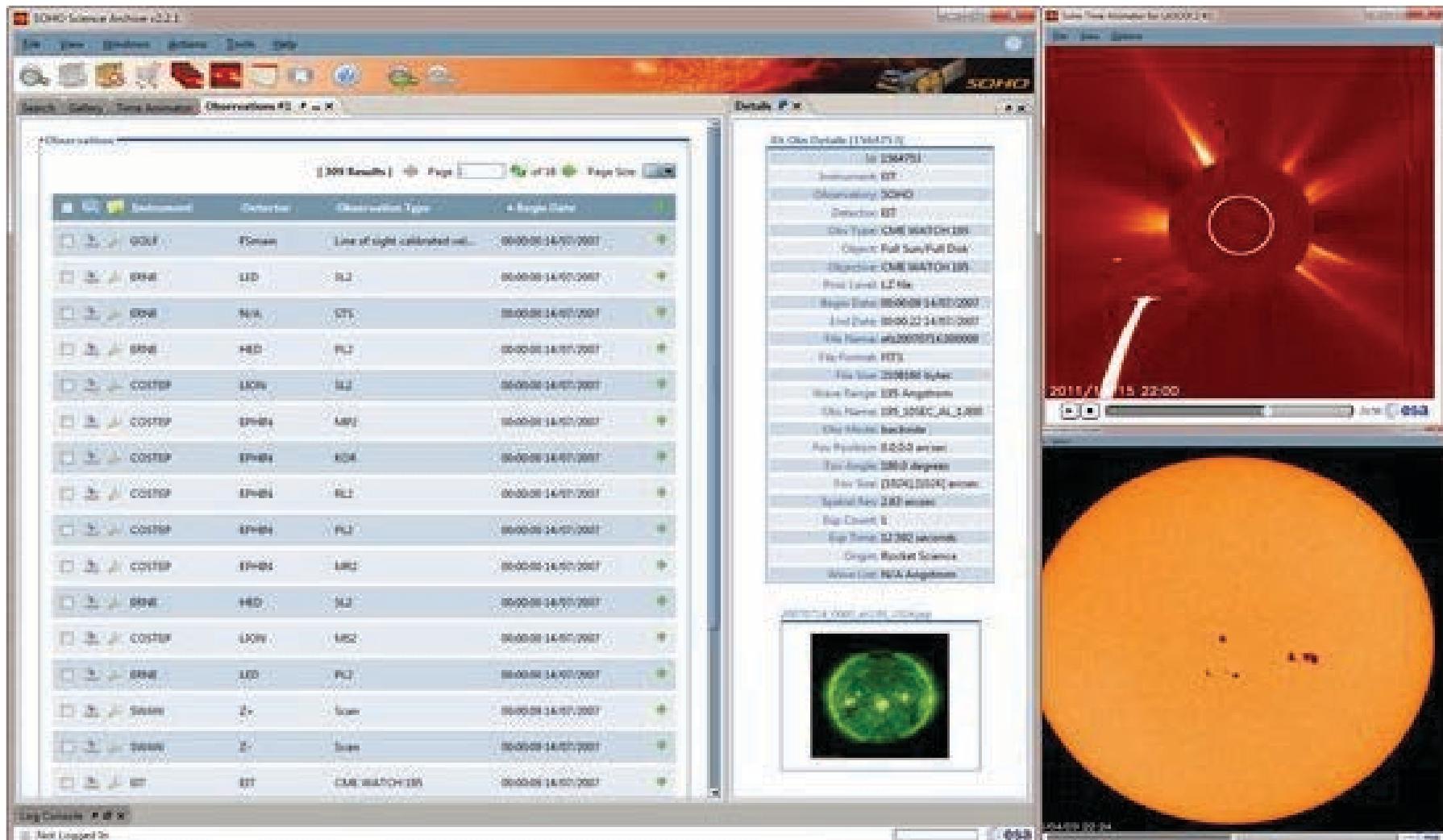
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European Space Agency

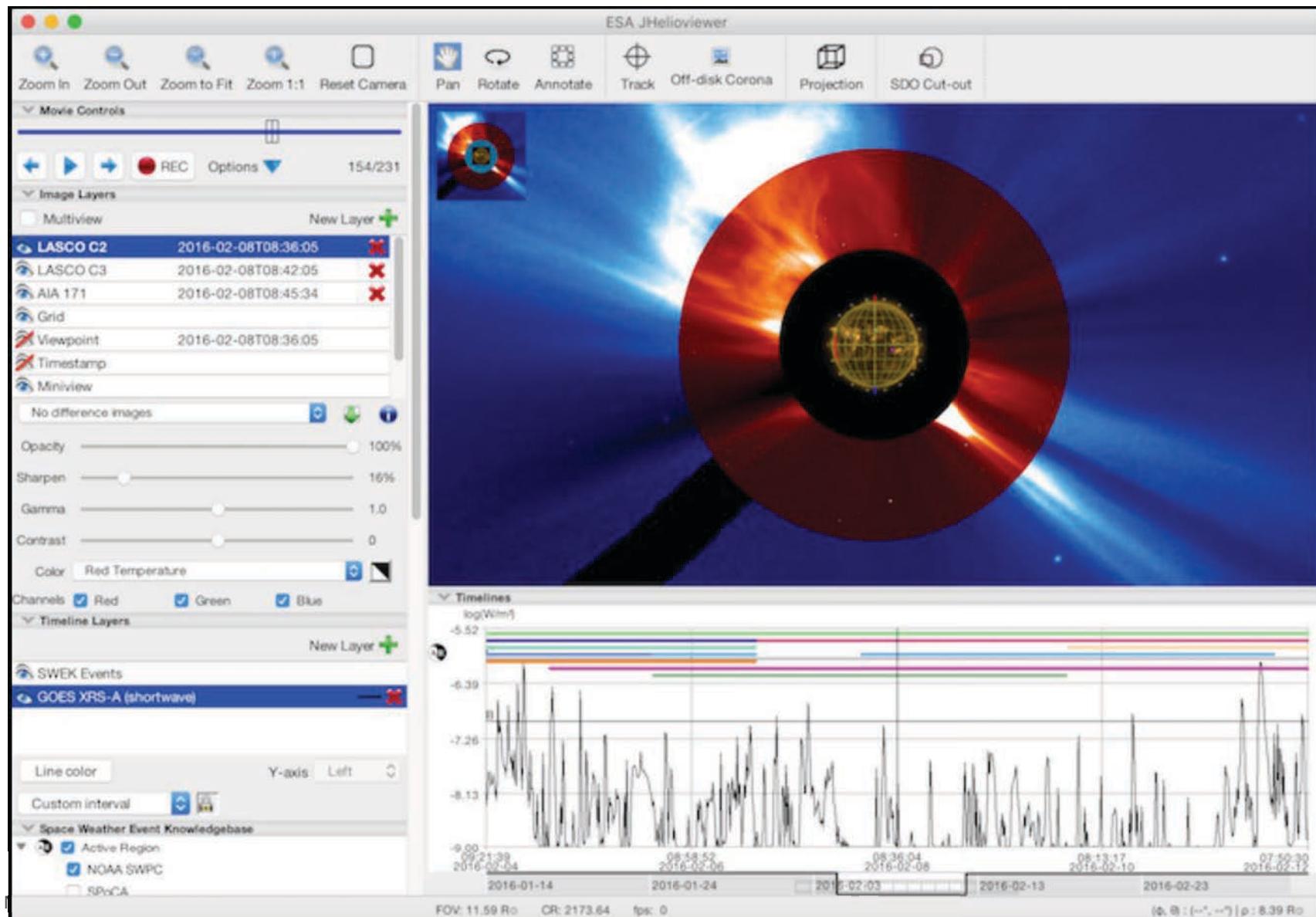
<http://archives.esac.esa.int>



The screenshot displays the Soho Science Archive interface. On the left, a list of observations is shown, including details like instrument (e.g., SOHO, EIT), detector (e.g., EIT, LASCO), observation type (e.g., Line of sight calibrated), and acquisition date (e.g., 00-00-00 14:00:00). In the center, a detailed view of an observation is provided, listing parameters such as Instrument (EIT), Observatory (SOHO), Detector (EIT), and various parameters related to the observation. On the right, two panels show solar imagery. The top panel shows a red solar image with a white circle highlighting a specific region, with a timestamp of 2010/11/15 22:00. The bottom panel shows a large orange solar image with a white circle highlighting a specific region, with a timestamp of 2010/11/15 22:00.

**Soho Science Archive
Including Proba-2 data**

<http://www.jhelioviewer.org>



Arnaud

<http://www.jhelioviewer.org>

JHelioviewer is an open-source application for visualization of solar images based on the JPEG 2000 standard.

It is part of the ESA/NASA Helioviewer Project and another client to the Helioviewer web services alongside the Helioviewer.org web application.

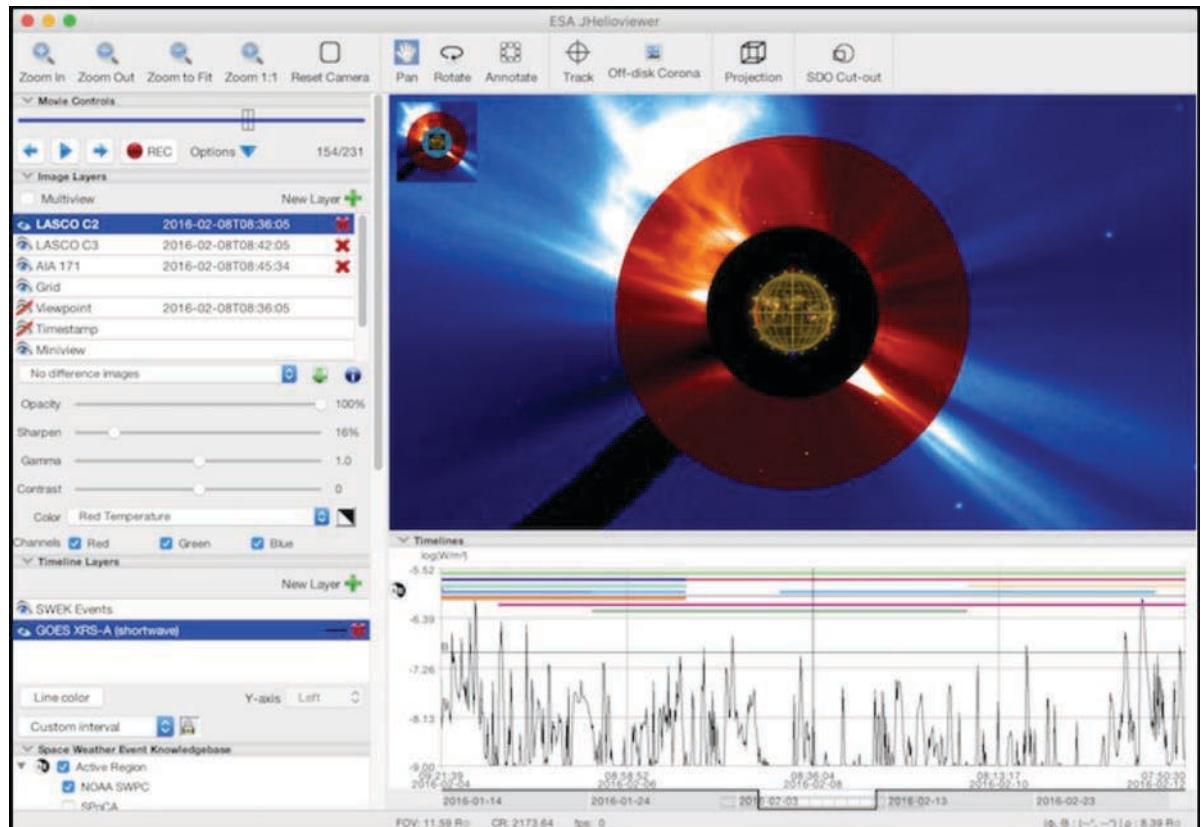
Commonly referred to as "Space Weather HelioViewer". Many of the new additions will be of particular interest to the Space Weather community.

JHelioviewer "2.10" released in March 2016 offers new features, including:
Combined view from different vantage points
(e.g. [STEREO](#) and [SDO](#)) for full sphere mapping

Timeline datasets viewing, synchronized with the time of the current image

Integration of features and events from the [Space Weather Events Knowledgebase \(SWEK\)](#), and alerts from the COMESEP system.

Magnetic field line extrapolations



<http://archives.esac.esa.int>

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Home

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ESASky

Videos

Scientific Tutorials

Publications

VOSpec

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European Space Agency

The Cluster Science Archive

<https://cosmos.esa.int/csa>



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cluster science archive

Cluster Science Archive » Access to the Archive



THE CLUSTER AND DOUBLE STAR SCIENCE ARCHIVE
VERSION 1.4.4, RELEASED 31 JULY 2017

→ START THE CSA GRAPHICAL USER INTERFACE [1.4.4]

To access the archive through its Java graphical user interface, please use the above link. This interface includes the following data services:

- Direct data download up to 1 GB or Schedule data download up to 40 GB
- Visualize key datasets (on-demand & pregenerated plots)
- Visualize particle distributions (on-demand only)
- Browse inventory information
- Store data profiles
- Automatic login

→ COMMAND-LINE WEBSITE

Since August 1, 2017, command line users must renew their **CSACOOKIE**, due to the installation of a new server on July 31. Use the above link to learn how to access the Cluster science archive data products via several ways including: wget, MATLAB, IDL and data streaming.

Check the [FAQ section](#) if any problem occurs, or [contact us](#).

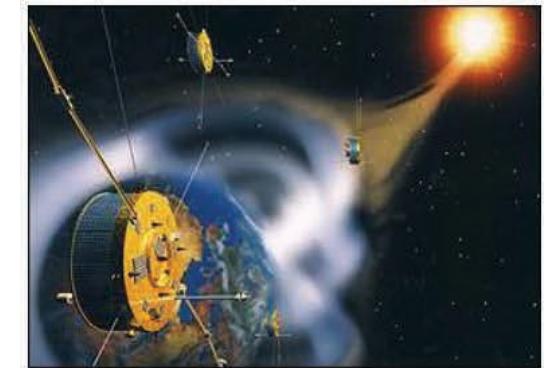
KEY NEW FEATURES

V1.4.4 Faster CSA server installed, minor update of inventory plots; for command line users, it **requires** their **CSACOOKIE** to be regenerated

Access to the Archive
Registration
Frequently Asked Questions (FAQ)
Multi-Spacecraft Analysis Techniques
Cluster quicklook plots
Double Star quicklook plots
Documentation
Software links
Other useful links
Workshops
CSA Helpdesk
Data Mining
People Editor
Self Registration Moderation
Find CSA Members

cluster science archive

- First 4-spacecraft magnetospheric constellation
- Polar orbit, probing most regions of the magnetosphere: solar wind, bow shock, Magnetosheath, MP, cusp, PSBL, plasmashell, AAR, outer plasmasphere, lobes...
- Various inter-spacecraft distances: from 3 km at the bow shock to 20,000 km along the magnetopause... over 17 years, under 2 very different solar cycles!
- Provide ground truth to test MHD-EPIC code for instance thanks to its magnetosphere coverage
- > 2500 publications and 120 PhD
- Still active with on-going coordinated observations campaign with MMS and THEMIS up to 2020
- > 100 TB of data



The Cluster Science Archive

<https://cosmos.esa.int/csa>



Journal of Geophysical Research: Space Physics

10.1002/2013JA019346

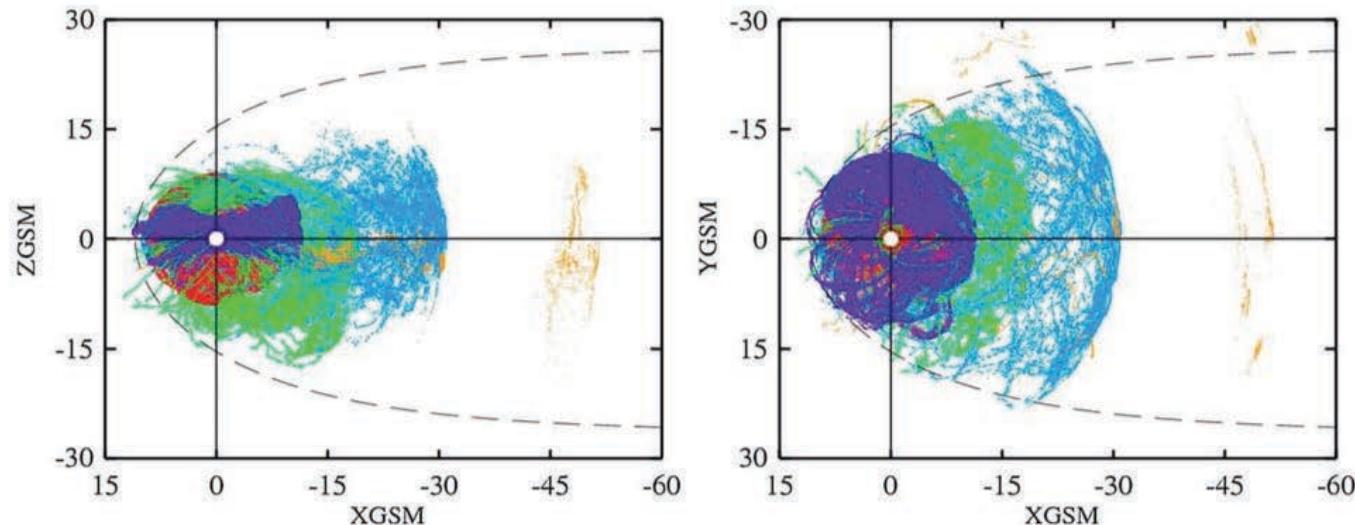
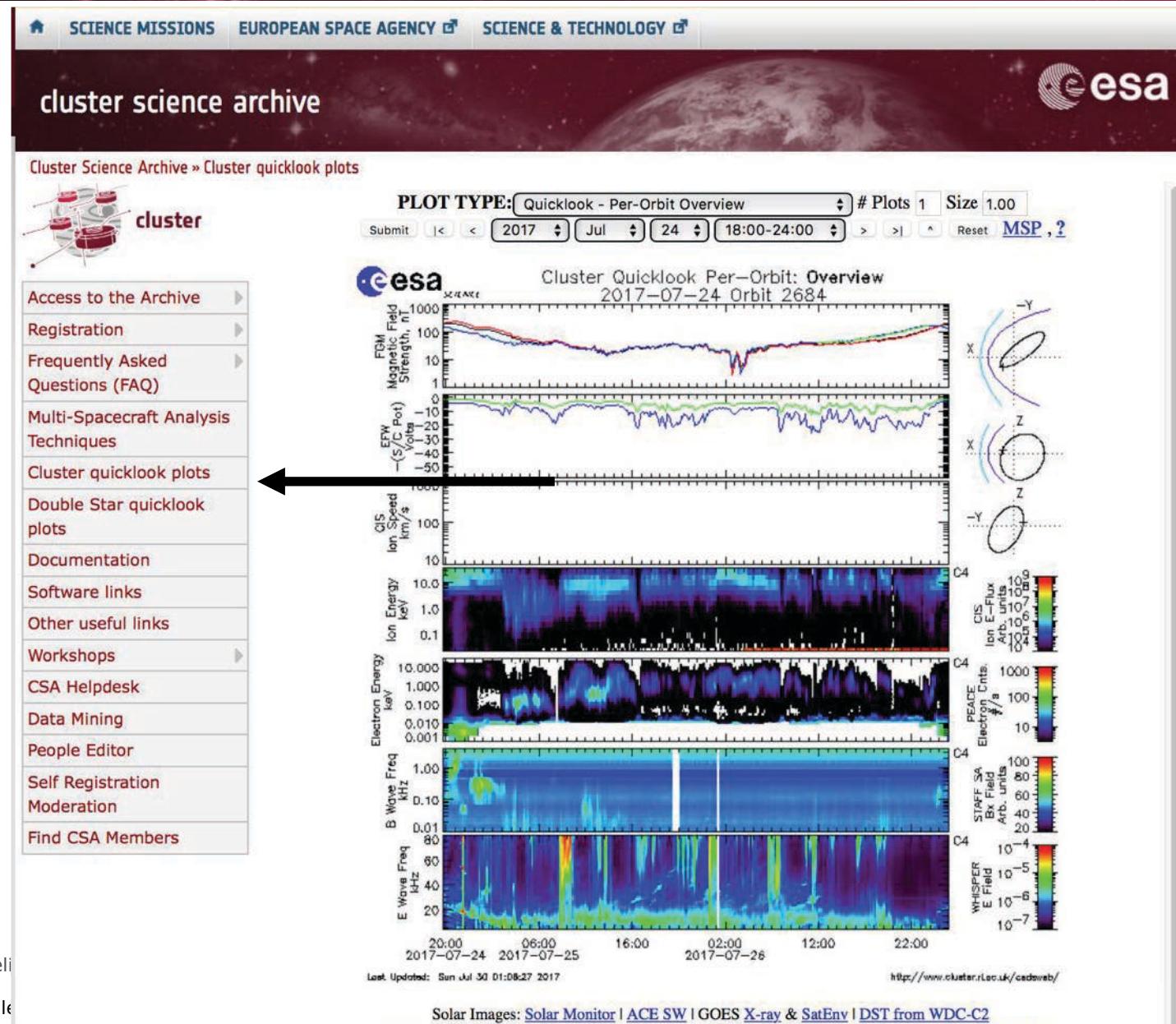


Figure 10. Spatial distribution of the data records in the modeling “grand” set, in projection on (left) XZ and (right) XY planes. Color coding: red = Polar, green = Cluster, light blue = Geotail, dark blue = THEMIS-A, -D, and -E, and light brown = THEMIS-B and -C. Average magnetopause location is shown by dashed lines.

Tsyganenko, 2014

The Cluster Science Archive

<https://cosmos.esa.int/csa>



The Cluster Science Archive

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DOCUMENTATION

Cluster Instrumentation

- [User guides, calibration reports, ICDs](#)
- [Other Documents](#)

Double Star Instrumentation

- [User guides, calibration reports, ICDs](#)
- [Other documentation](#)

CAA Documentation

- [General](#)
- [CEF format and Metadata Information](#)
- [Raw data media](#)
- [Technical Notes](#)

Other Documentation

- [ESOC Documentation](#)
- [ECLAT Documentation](#)

[View regenerate](#)

cluster science archive



Cluster Science Archive » Documentation



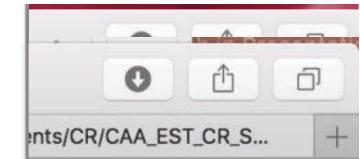
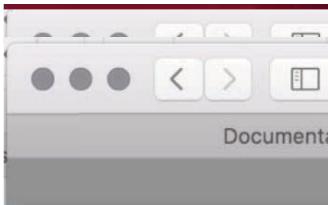
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INSTRUMENT USER GUIDES, CALIBRATION REPORTS AND ICD'S

Each instrument has produced the User Guide (UG) and the Calibration Report (CR) to help the CAA users in their CAA data exploitation activities. The Instrument Control Document (ICD) is a technical document between the instrument team and the CAA core team and is given here for the completeness. The purpose of these documents are:

- UG describes the key science datasets and their caveats and quality flags.
- CR describes the key results of the instrument calibration and cross-calibration activities
- ICD describes the detail syntax of the instrument datasets

User Guide	Calibration Report	ICD
ASPOC,v1.0	ASPOC,v1.0	ASPOC,v1.3
CIS,v3.7	CIS,v1.7	CIS,v3.4.3
DWP,v1.11	DWP,v1.3	DWP,v11.0
EDI,v3.5	EDI,v1.6	EDI,v2.7
EFW,v3.8	EFW,v2.4	EFW,v4.10
FGM,v6.2	FGM,v6.2	FGM,v6.2
PEACE,v2.6	PEACE,v1.7	PEACE,v6.4
RAPID,v5.1	RAPID,v5.0	RAPID,v5.4
STAFF,v3.5	STAFF,v4.2	STAFF,v5.5
WBD,v3.1	WBD,v3.0	WBD,v3.3
WHISPER,v1.6	WHISPER,v1.7	WHISPER,v1.12
AUX,v1.0		



Project

7.1 Comparison of STAFF-SC Spin plane DC field with FGM

7.1.1 Case studies

The STAFF/FGM comparison plots below have been presented at the *1st Cross-Calibration Workshop, 2006, ESTEC*, and at the *8th Cross-Calibration Workshop, Kinsale, Ireland, 28-30 October 2008*.

A first problem was identified after launch: the perpendicular DC-field measured by the spinning spacecraft at the spin frequency was not the same from S/C #1 than from the other S/C: The difference was of ~ 8 to 20 % with respect to FGM. S/C# 1 gave always lower values than other S/C by ~10%, and the difference was confirmed by FGM.

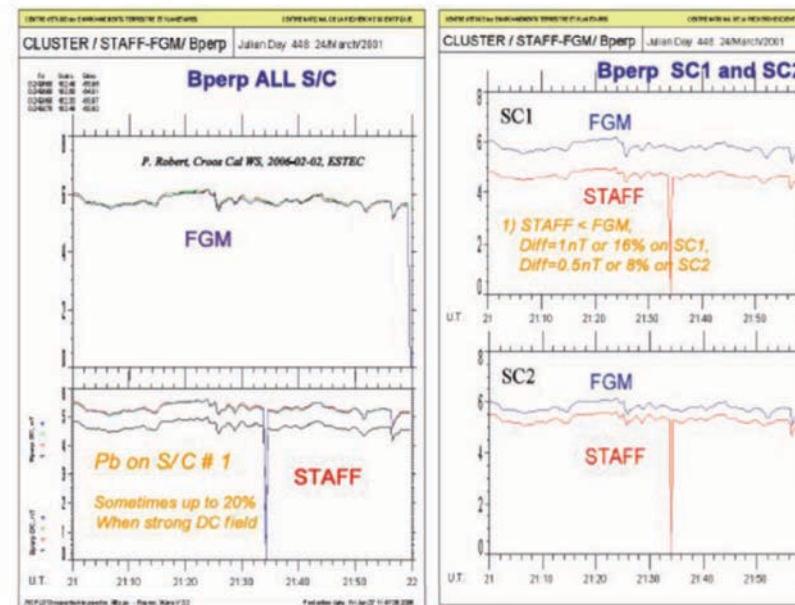


Figure 6: Comparisons of the modulus of the DC field in the plane perpendicular to the spin axis.

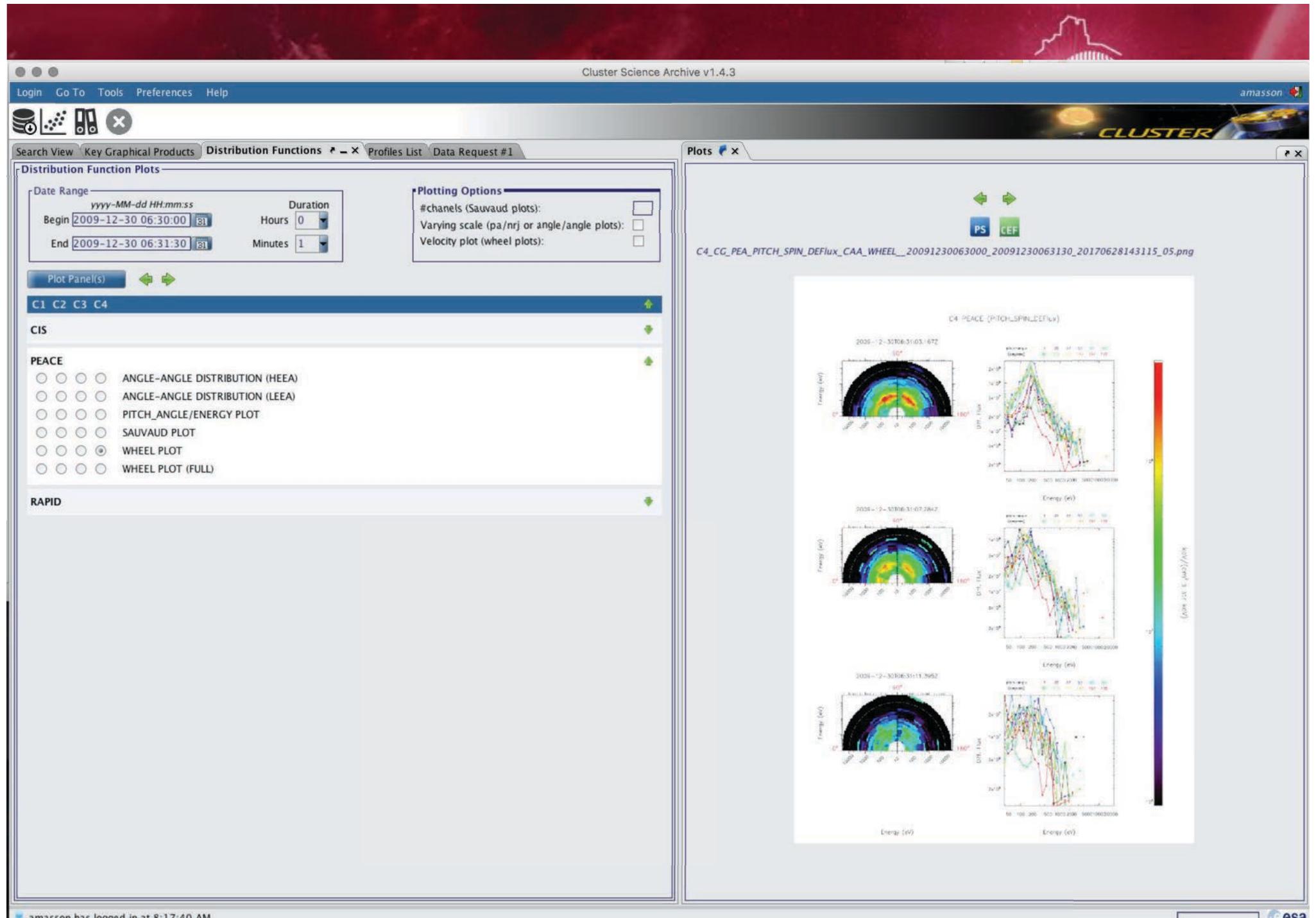
Left: comparison of the perpendicular components between the four spacecraft. Top panel is for FGM and bottom for STAFF-SC. Right: comparison of FGM and STAFF-SC for C1 and 2.

R-002
4.2
04-27
of 62

The Cluster Science Archive

II.5. Visualization tools

- Key graphical products (PI interaction)
- Pre-generated plots (1h, 6h, 24h)
- On-demand key graphical products visualization
- QL plots inc. cross-calibration plots
- Inventory plots



amasson has logged in at 8:17:40 AM

Arnaud Masson | ESA Heliophysics Archives: a key asset for ISWI | Boston College | 02/08/2017| Slide 23

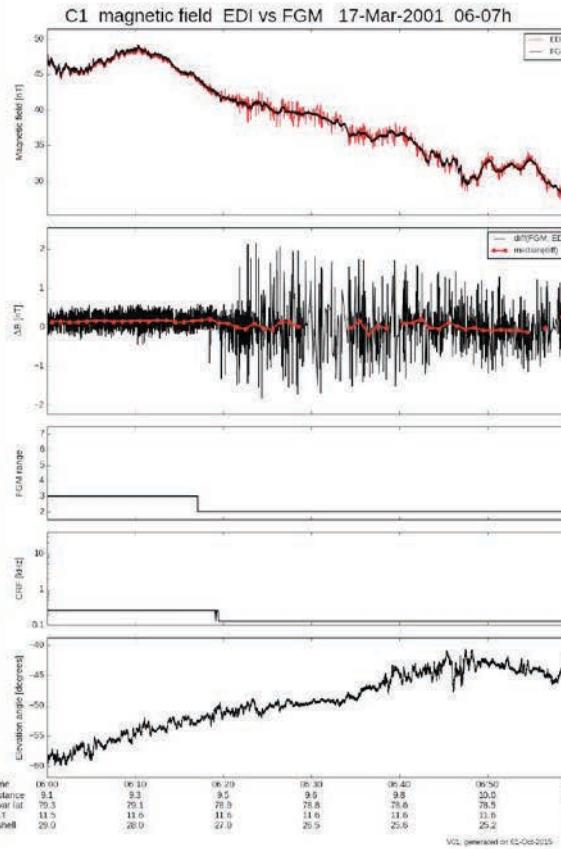
ESA UNCLASSIFIED - Releasable to the Public

European Space Agency

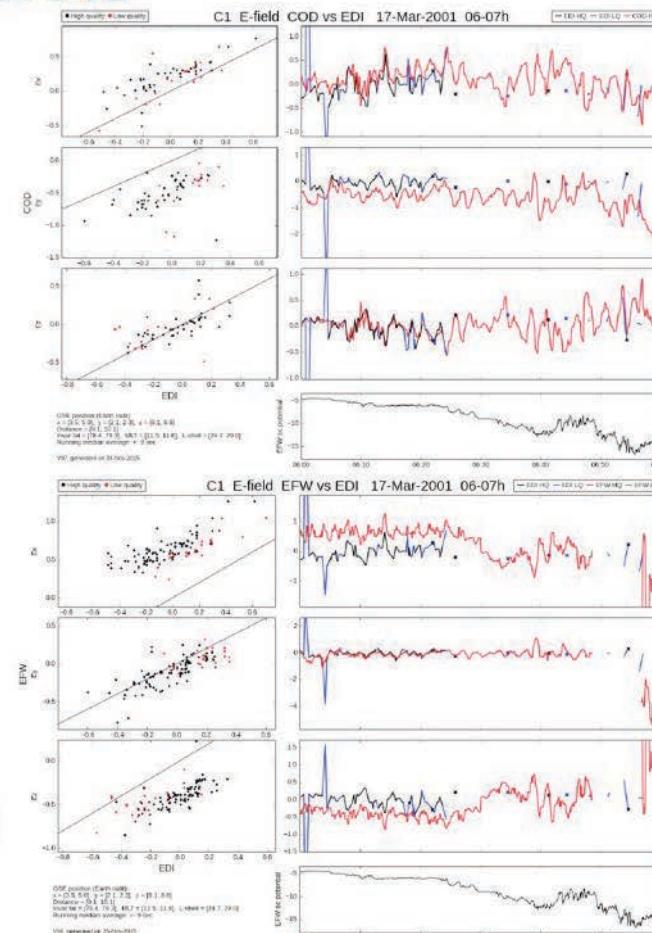
The Cluster Science Archive



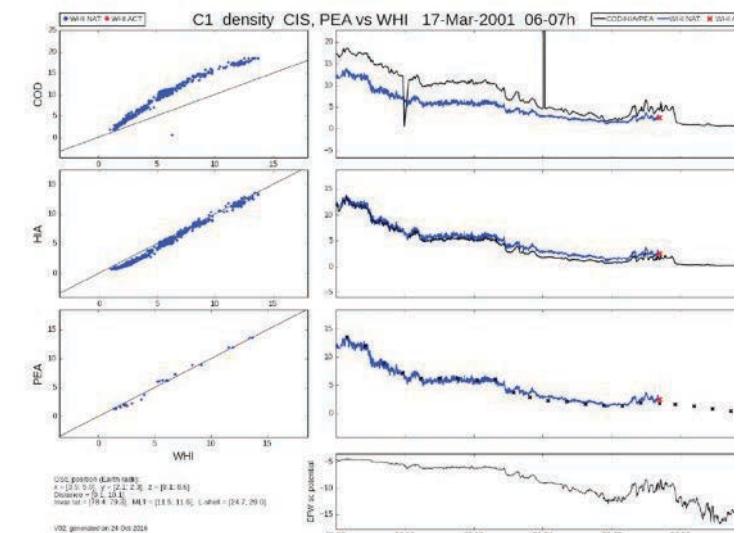
Cross-calibration plots



Magnetic field



Electric field



Density



Search View Quicklook Plots Inventory Plots Data Request #1

FP7/H2020 datasets included in the archive

- MAARBLE: IMAGE OUJ ground-based Pc3-5 power
- MAARBLE: CARISMA PINA ground-based Pc3-5 power
- MAARBLE: CARISMA RANK ground-based Pc3-5 power
- MAARBLE: IMAGE RVK ground-based Pc3-5 power
- MAARBLE: IMAGE SOD ground-based Pc3-5 power
- MAARBLE: Themis-A Pc 1-2 Wave Spectra and PP Parameters
- MAARBLE: Themis-A Pc 3-5 Wave Spectra and PP Parameters
- MAARBLE: Themis-A VLF Wave Spectra and PP Parameters
- MAARBLE: Themis-B Pc 1-2 Wave Spectra and PP Parameters
- MAARBLE: Themis-B Pc 3-5 Wave Spectra and PP Parameters
- MAARBLE: Themis-B VLF Wave Spectra and PP Parameters
- MAARBLE: Themis-C Pc 1-2 Wave Spectra and PP Parameters
- MAARBLE: Themis-C Pc 3-5 Wave Spectra and PP Parameters
- MAARBLE: Themis-C VLF Wave Spectra and PP Parameters
- MAARBLE: Themis-D Pc 1-2 Wave Spectra and PP Parameters
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- MAARBLE: Themis-D VLF Wave Spectra and PP Parameters
- MAARBLE: Themis-E Pc 1-2 Wave Spectra and PP Parameters
- MAARBLE: Themis-E Pc 3-5 Wave Spectra and PP Parameters
- MAARBLE: Themis-E VLF Wave Spectra and PP Parameters

Access the archive - Cosmos csa.esac.esa.int CAIO - Cluster Science Archive Inter-Operability subsystem

Now, in front of our request, we put --content-disposition. This is a wget command which *should* (see [Specifying the wget output filename](#)) give the output filename a suitable name.

```
--content-disposition "https://csa.esac.esa.int/csa/aio/product-action?  
DATASET_ID=C1_CP_FGM_SPIN&START_DATE=2003-03-03T00:00:00Z&END_DATE=2003-03-  
05T00:00:00Z&NON_BROWSER&CSACOOKIE=<cookie>"
```

And finally, put wget on the front and it's ready to run.

```
wget --content-disposition "https://csa.esac.esa.int/csa/aio/product-action?  
DATASET_ID=C1_CP_FGM_SPIN&START_DATE=2003-03-03T00:00:00Z&END_DATE=2003-03-  
05T00:00:00Z&NON_BROWSER&CSACOOKIE=<cookie>"
```

If everything works correctly, you will receive a .tar.gz file with your username and a number, e.g., hmiddlet1784379.tar.gz.

SPECIFYING THE FILENAME

You can specify the filename returned by replacing --content-disposition with -O <filename>, where -O is uppercase O for Oscar. The example above would become:

```
wget -O <filename>.tar.gz "https://csa.esac.esa.int/csa/aio/product-action?  
DATASET_ID=C1_CP_FGM_SPIN&START_DATE=2003-03-03T00:00:00Z&END_DATE=2003-03-  
05T00:00:00Z&NON_BROWSER&CSACOOKIE=<cookie>"
```

USING THE DIFFERENT TYPES OF REQUEST

The other data requests are used with wget in exactly the same way as above, using the https requests and parameters detailed on their respective tabs, but their outputs will be different. If you are supplying the filename, as detailed above, it makes working with the data easier if the correct file type is given:

DATA REQUEST

A data request, if successful, will supply a file package in **.tar.gz format**, as in the example above.

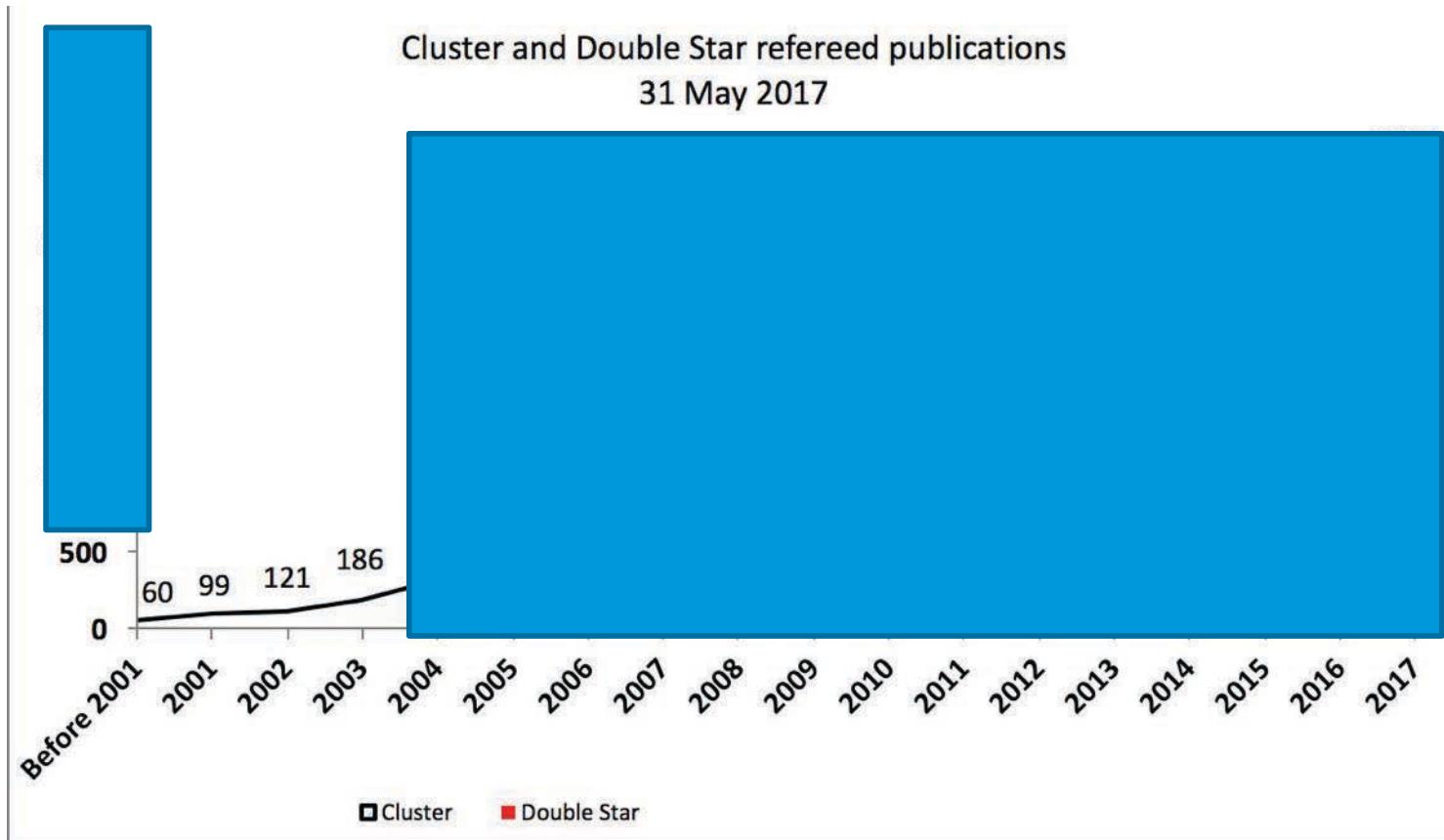
ASYNCHRONOUS DATA REQUEST

When you submit a successful asynchronous data request, the returned file will be a **text file** that should contain text similar to this:

The Cluster Science Archive



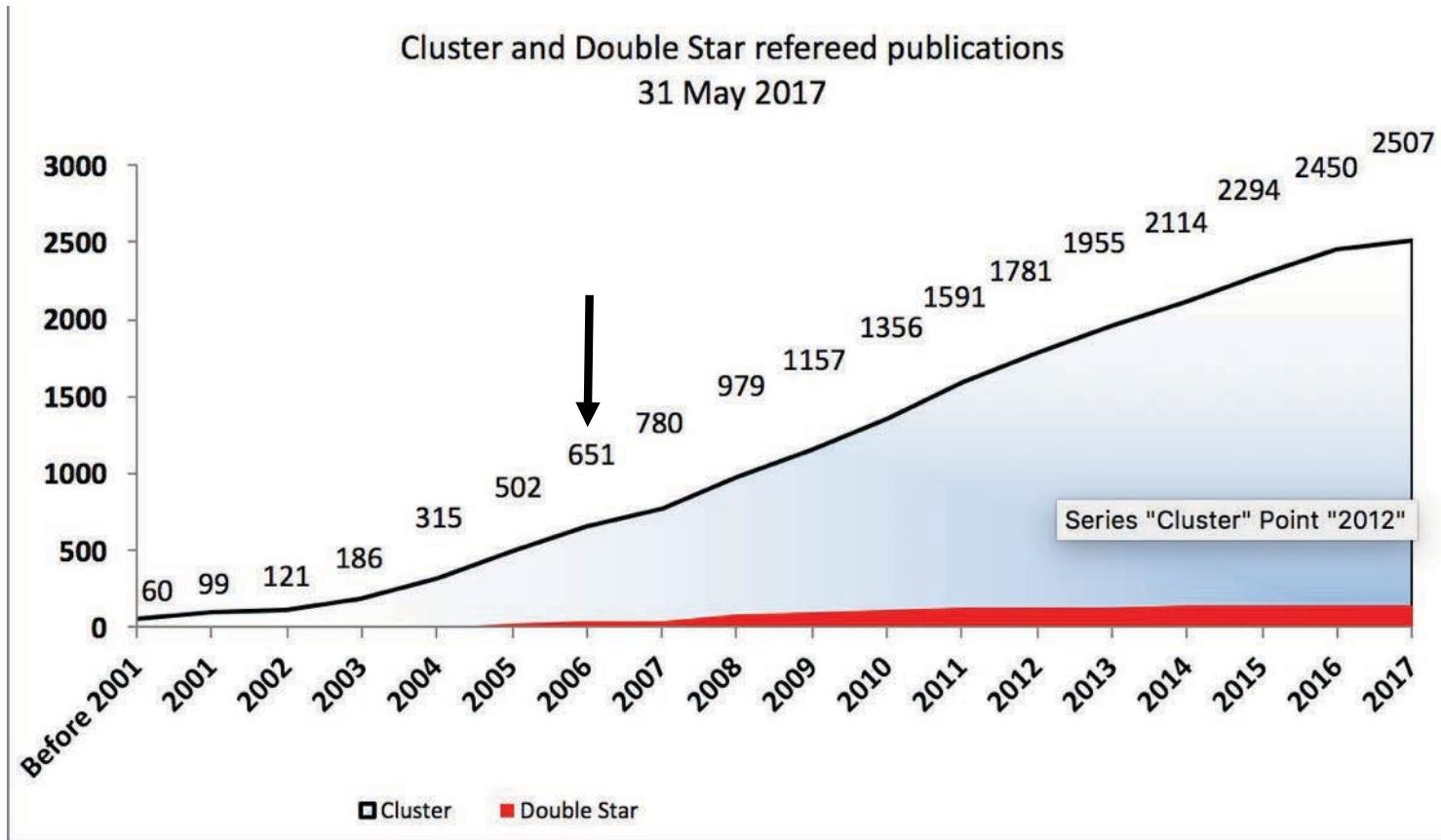
II.10. Wrap-up



The Cluster Science Archive



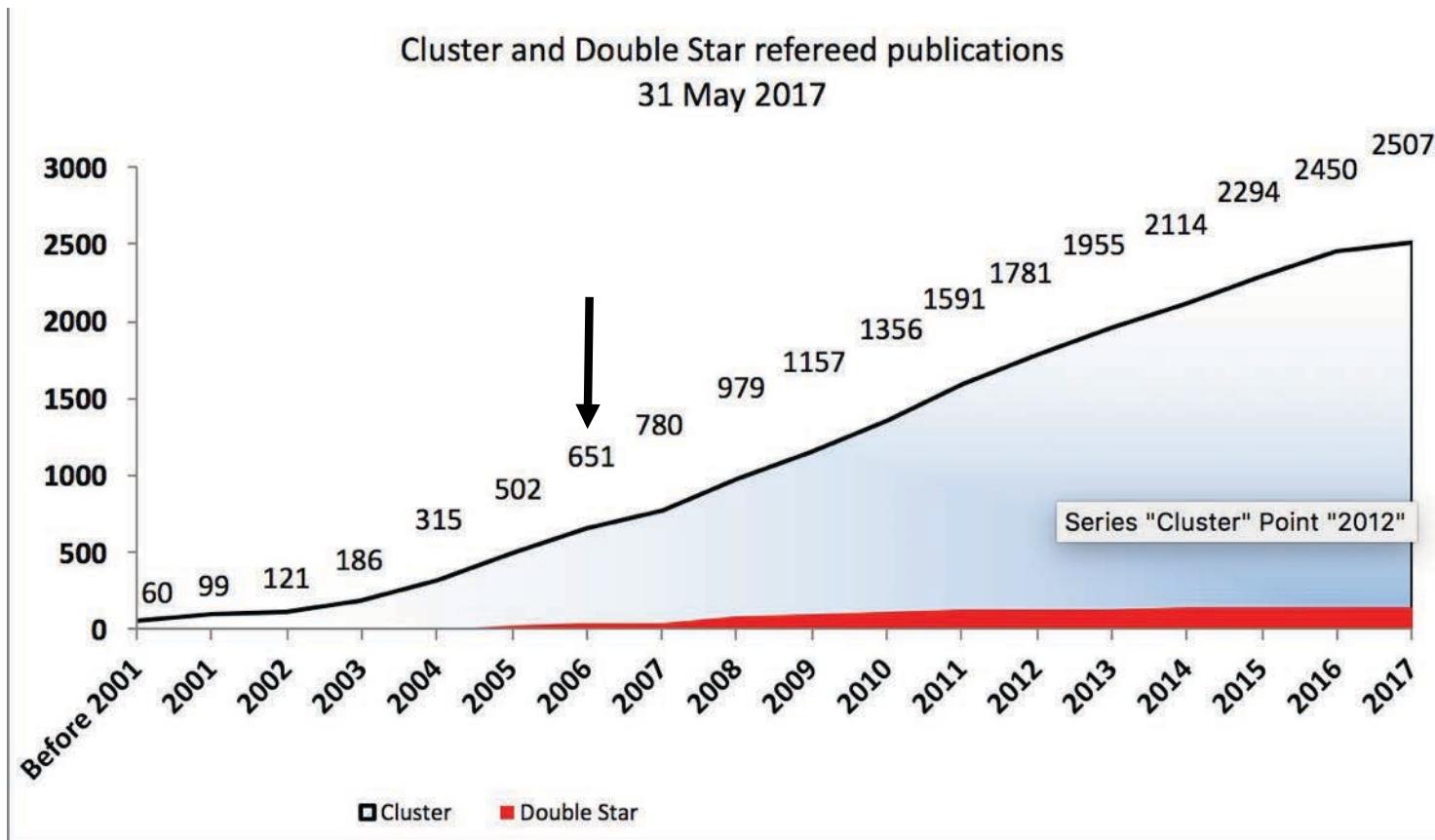
II.10. Wrap-up

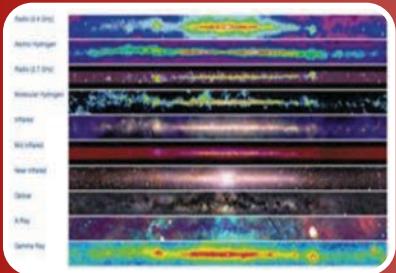


The Cluster Science Archive

Rank	Country
1	USA
2	China
3	UK
4	France
5	Sweden
6	Germany
7	Russia
8	Austria
9	Japan
10	Czech Rep.
11	Finland
12	Belgium
13	Italy
14	Netherlands
15	India
16	Canada
17	Norway
18	Hungary
19	Poland
20	Greece
21	Romania
22	Brazil
23	Australia
24	Mexico
25	Korea
26	Serbia
27	Israel
28	Switzerland
29	Palestinian T
30	Pakistan
31	Argentina
32	Armenia
33	South Africa
34	Thailand
35	Turkey
36	Chile
37	New Zealand
38	Spain
39	Ireland
40	Croatia
41	Iran

II.10. Wrap-up

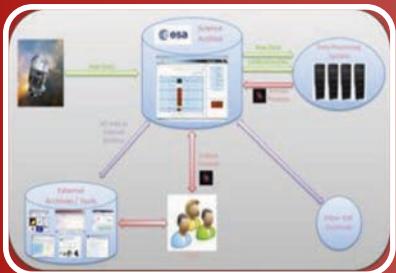




Enable maximum ***scientific exploitation*** of data sets



Enable efficient ***long-term preservation*** of data, software and knowledge, using modern technology



Enable cost-effective archive production by ***integration in, and across, projects***