

#### La Palma, Canary Islands, Spain 3 - 7. October, 2021

# Analytical simulations of the effect of satellite constellations on astronomical observations

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Based on https://arXiv.org/abs/2108.12335

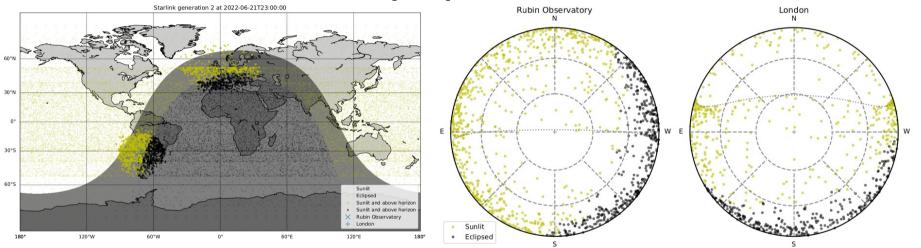
5 October 2021 | Bassa (ASTRON) | Analytical simulations



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# Discrete simulations of all-sky impact



- Constellation/satellite properties
- Observatory latitude

- Solar declination (time of year)
- Solar elevation (time of day)

See McDowell (2020), Hainaut & Williams (2020), Lawler et al. (2021)

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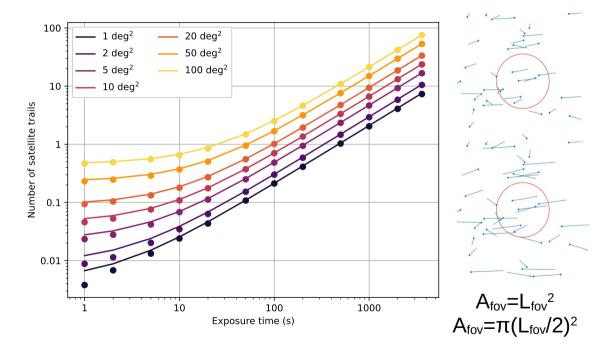
# **Discrete simulations of observational impact**

Count trails per exposure:

- Exposure time (t<sub>exp</sub>)
- Field-of-view (A<sub>fov</sub> and L<sub>fov</sub>)
- Satellite density (ρ<sub>sat</sub>)
- Angular velocity (ω<sub>sat</sub>)

$$N_{\text{trail}} = \rho_{\text{sat}}(A_{\text{fov}} + L_{\text{fov}}\omega_{\text{sat}}t_{\text{exp}})$$

Average over many discrete constellation realizations (slow for 65000 satellites!)

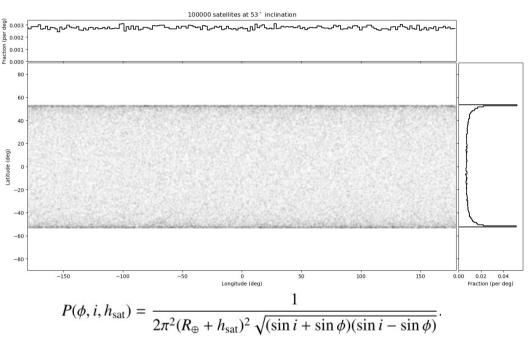




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# Analytical simulations of observational impact



Satellite density is uniform with longitude but strongly peaked with latitude (arcsine distribution)

Use analytical probability density to define orbital shell density (n<sub>sat</sub>/km<sup>2</sup>)

Transform to sky density  $\rho_{sat}$  ( $n_{sat}$ /deg<sup>2</sup>) for sight line at distance d, impact angle  $\alpha$ , field-of-view A, and  $N_{sat}$  objects in orbital shell

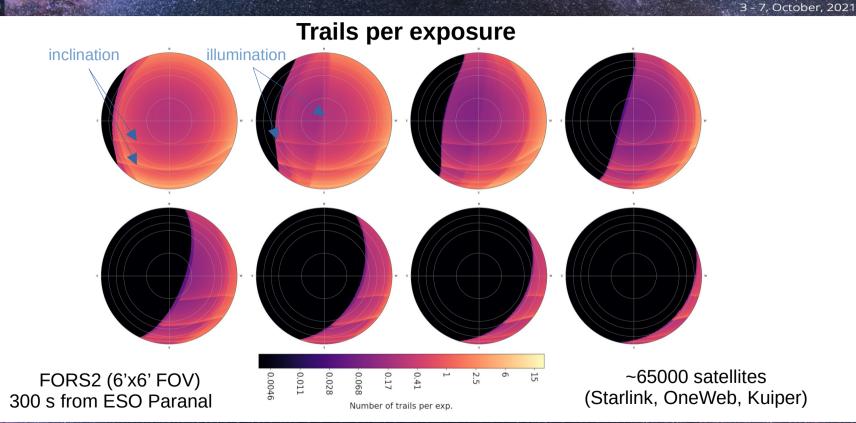
$$\rho_{\rm sat} = N_{\rm sat} P(\phi, i, h_{\rm sat}) \frac{d^2 A}{\cos \alpha}$$

Applicable to fully populated shells



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## Simple photometric modelling

• Apparent magnitude:

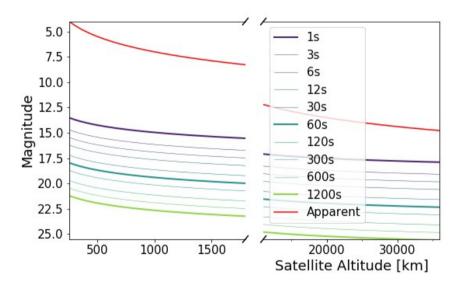
 $m_{\text{sat}} = m_{1000\text{km}} + 5\log_{10}(d_{\text{sat}}/1000) + kd_{\text{sat}}/h_{\text{sat}}$ 

Depends on distance  $d_{sat}$ , altitude  $h_{sat}$  and extinction k ( $m_{1000km}$ ~7 for Starlink and OneWeb)

• Effective magnitude:

$$m_{\rm eff} = m_{\rm sat} - 2.5 \log_{10} \frac{t_{\rm eff}}{t_{\rm exP}} = m_{\rm sat} - 2.5 \log_{10} \frac{r}{\omega_{\rm sat} t_{\rm exp}}$$

Depends on instrument resolution r, exposure time  $t_{\text{exp}}$  and angular velocity  $\omega_{\text{sat}}$ 



#### See afternoon talk by Olivier Hainaut!



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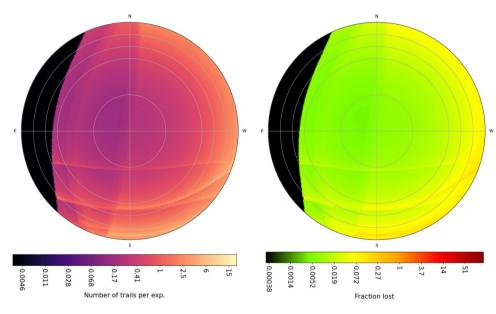
## Effect on observations

Definition:

- $m_{eff} < 1\sigma$  detection limit: no effect
- otherwise, area of 5 arcsec wide trail lost from detector or slit
- if m<sub>eff</sub> > heavy saturation limit, entire observation lost

Check of all orbital shells, all constellations.

*Fraction lost* is the fraction of observations or pixels in observation lost due to satellite trails



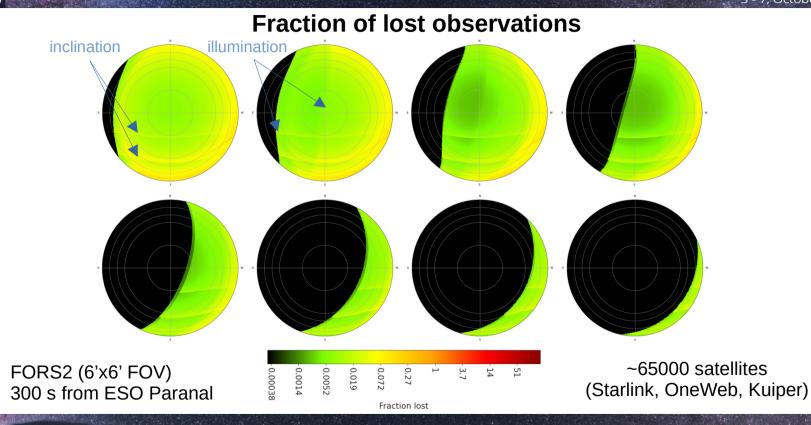
#### FORS2 (6'x6' FOV) 300 s from ESO Paranal

~65000 satellites (Starlink, OneWeb, Kuiper)



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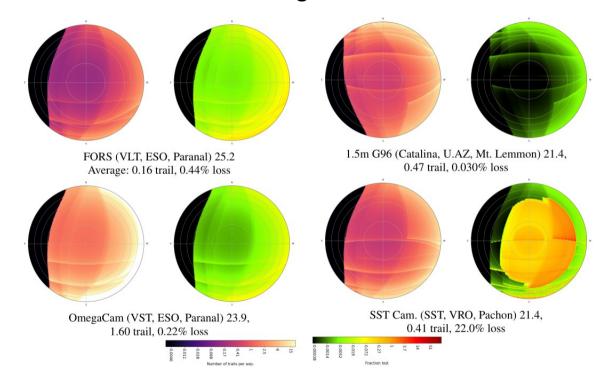


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Imagers

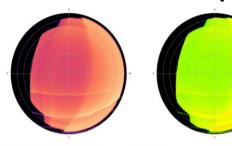




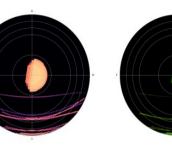
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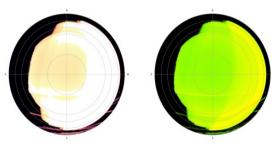
Spectrographs



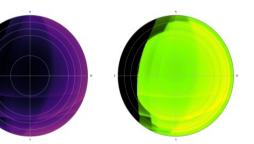
FORS (VLT, ESO, Paranal) 25.2 Average: 0.64 trail, 8.8% loss



4MOST-HiRes (VISTA, ESO, Paranal), 0.33 trail, 0.018% loss



4MOST-LowRes (VISTA, ESO, Paranal), 14.7 trails, 0.78% loss



HARMONI (ELT, ESO, Armazones), 0.007 trail, 0.70% loss



### Summary

- Derived analytical expressions to assess observational impact
- Analytical simulations allow for fast computation of observational impact for an observatory and satellite constellation(s)
- Impact on sky brightness: see afternoon talk by Olivier Hainaut
- Satellite trails will impact all observatories (both imagers and spectrographs)
- High sensitivity, large field-of-view instruments and/or long exposures are worst affected

# **Further information**

- Paper: arxiv.org/abs/2108.12335 (A&A under review)
- Online simulator: www.eso.org/~ohainaut/satellites/simulators.html
- Github repository: github.com/cbassa/satconsim

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# Thank you for your attention!



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#### Impact on observability of specific sources

