Ionospheric plasma fluctuation response to space weather events over the equatorial and low latitude region.

Patrick Essien, C.A.O.B Figueiredo, H. Takahashi, C.M. Wrasse,







Atmospheric Physics Lab Department of Physics University of Cape Coast Ghana

Outline

- Introduction
- Geomagnetic storms
- Ionospheric Plasma Fluctuation Response
- Conclusion

Equatorial and low latitude ionosphere

• In the equatorial and low latitude ionosphere, there are specific variabilities that are important for the ionospheric weather

Equatorial Ionization Anomaly (EIA)



Equatorial Plasma Bubbles (EPBs)



Travelling Ionospheric Disturbances (TIDs)



These irregularities/disturbances in the ionosphere can disrupt radio waves causing errors in radio wave GNSS positioning systems.

The lonosphere: a layered structure

Due to different ionization production and loss processes the electron density profile with altitude shows a layered structure that varies with:

- Time of day
- Day-to-Day
- Seasonal
- Location on Earth
- Solar activity and Space Weather conditions



Drivers of Ionosphere Thermosphere Weather



Courtesy J. Forbes

METHOD

■GNSS TEC MAP

1. Total Electron Content (TEC) is calculated from dual-frequency GNSS data.

 $TEC = \int_{r}^{s} N_{e} ds$

- 2. Slant TEC is converted to vertical TEC.
- 3. 1-hour running average is subtracted from the original TEC to obtain perturbation component of TEC.

 $dTEC(t) = TEC(t) - < TEC(t \pm 30 min) > .$

 Map the TEC data onto the ionospheric shell at an altitude of 350 km with grid of 0.25°x 0.25° in lat. and long.



 The ionospheric delay (I) on GPS signals can be described in terms of TEC, inversely proportional to a square of frequency

$$\Delta^{iono} = \frac{40.3}{f^2} TEC$$





Essien et al., 2021

Traveling Ionospheric Disturbances (TIDs)





Wright et al., 2017

March 2015 (St. Patricks Day) Storm

- Coronal images recorded during 00:00– 03:12 UT (a–h) on 15 March 2015
- Flare with partial halo CME in March 2015 (One was a magnitude C9.1).
- Velocity of larger CME calculated to be ~ 606 km/s corresponding to ~ 2-3 day arrival time



(Wu, CC., Liou, K., Lepping, R.P. et al.)



Geomagnetic Index (Dst)



- Strong (~50 nT) compression on 17th indicative of shock reaching magnetosphere.
- Magnetic storm took place on 17th with index as high as -223 nT showing generation of strong ring current.
- Main phase onset occurs for ~ 1 day.
- The recovery phase of the storm was slow long lasting ring current.

ΔTEC in Mar 2015



17-18

TIDs generated during 2015 St. Patricks dayGeomagnetic Storm0.25(5x5)2017/03/17



Longitude (°)

Longitude (°)

14:00:00 UT (11:00:00 BST)



Plasma Fluctuations begun at 14:00



Analysis of the fluctuations using 1D-FFT cross-spectrum





Essien et al., 2021



2017 September Storm

Active Region 12673

SDO HMI Magnetogram 6-Sep-2017 18:46:41.700



There was geomagnetic storm after the 2 days of the solar storm



Ionospheric Fluctuation Response to the Storm



Longitude (°)







Equatorial SpreadF

Equatorial Plasma Bubbles



2018 August Storm



TID characteristics on 26th August A_m=0.11 Λ_{H} =644.1km T=26.5 min C_H=344.0m/s Φ =344.0°

Seasonal variation of TIDs during Geomagnetic Quiet and disturbed Time



Conclusion

- Ionospheric plasma fluctuations are generated during geomagnetic storms as well as the recovery phase.
- Those TIDs can trigger EPBs which is the post sunset event.
- The statistics of the seasonal variation of the TIDs changes when series of geomagnetic storms occur within a particular month or season.

Storm	Dst Index (nT)	TIDs Velocity (m/s)	TIDs Wavelength (km)
March 2015	-233	306.0	633.5
September 2017	-110	263.0	641.6
August 2018	-190	344.0	644.1

Acknoweledgement



UNITED NATIONS Office for Outer Space Affairs















Mathematical Sciences

RWANDA



Conselho Nacional de Desenvolviment Científico e Tecnológico

