

# A Novel Ionospheric Diagram for Understanding Ionospheric Characteristics and Behaviour

## United Nations Workshop On the International Space Weather Initiative: The Way Forward

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## 1. Ellingham Diagram

## 2. Poynting Vector

## 3. Electron Densities

3.1. D – Layer

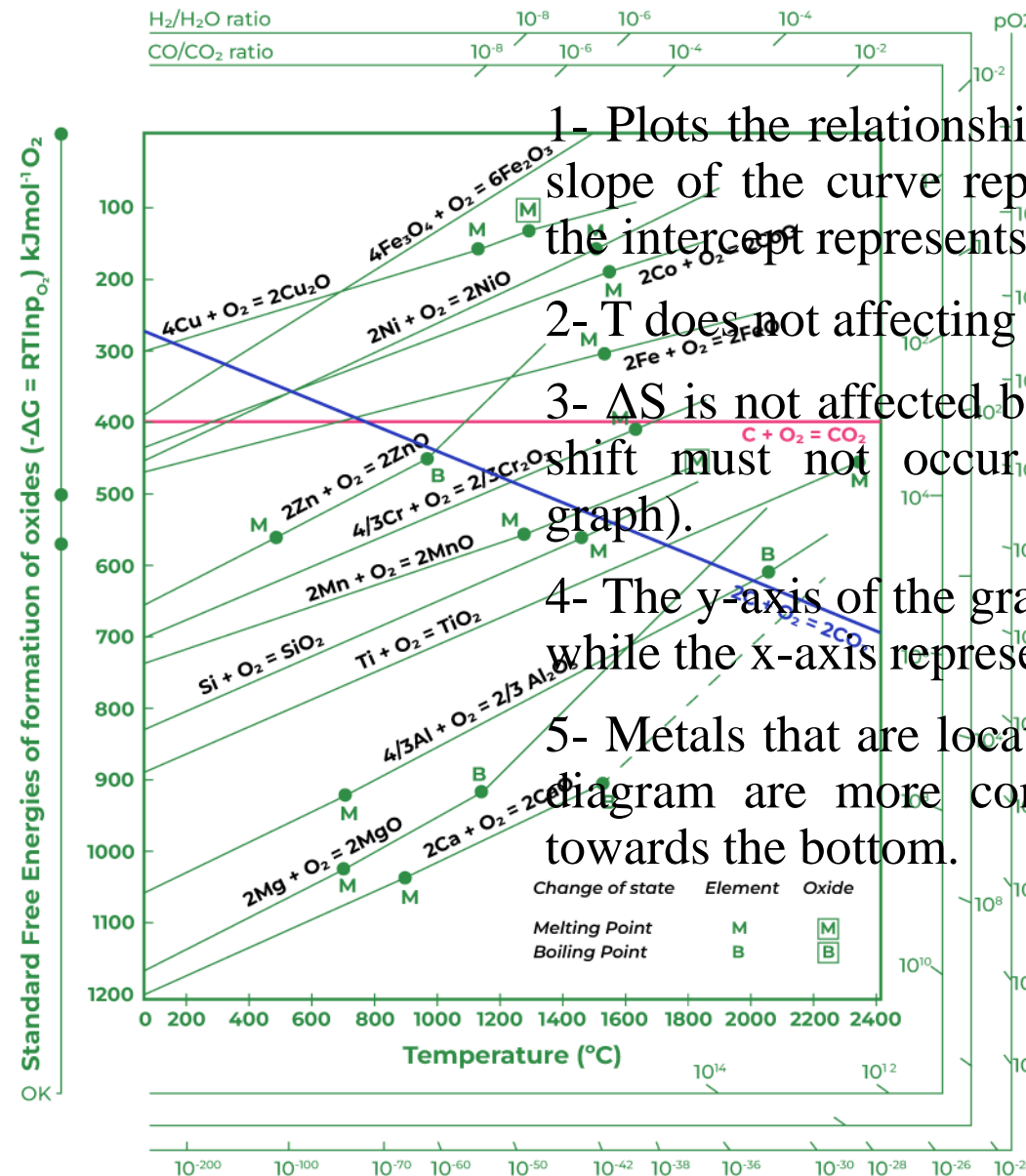
3.2. E – Layer

3.3. F1 – Layer

3.4. F2 – Layer

## 4. Draft Diagrams

# Ellingham Diagram?!?



- 1- Plots the relationship between T and  $\Delta G$ . The slope of the curve represents the entropy, while the intercept represents the H.
- 2- T does not affecting  $\Delta H$  (slopes).
- 3-  $\Delta S$  is not affected by temperature, but a phase shift must not occur (sudden changes in the graph).
- 4- The y-axis of the graph represents temperature, while the x-axis represents  $\Delta G$ .
- 5- Metals that are located closer to the top of the diagram are more common than those located towards the bottom.

$$\Delta G = \Delta H - T\Delta S$$

$$\Delta G = RT \ln K_p$$

$$V = \frac{RT}{p}$$

$$A + B \rightarrow C$$

$$K_p = \frac{p_C}{p_A p_B}$$

Ellingham, H. J. T. (1944)

# Poyting Vector

$$\vec{S} = \frac{1}{\mu_0} (\vec{E} \times \vec{B})$$

$$\vec{S} = \frac{\sigma}{\mu_0} (\vec{J} \times \vec{B})$$

Dipole magnetic field

$$\vec{B} = B_r \hat{r} + B_\theta \hat{\theta} + B_\phi \hat{\phi}$$

$$B_r = -\frac{2\mu_0 M \cos \theta}{r^3}$$

$$B_\theta = -\frac{\mu_0 M \sin \theta}{r^3}$$

$$B_\phi = 0$$

$$\vec{v} = v_r \hat{r} + v_\theta \hat{\theta} + v_\phi \hat{\phi}$$

# Poynting Vector

$$\vec{S} = \frac{Mq_e^3 n_e^2}{m_e v_e r^3} [(v_\phi \sin \theta) \hat{r} - [(2v_\phi \cos \theta) \hat{\theta}] + [(2v_\theta \cos \theta - v_r \sin \theta) \hat{\phi}]]$$

# Collision Frequency and Electron Density

$$\nu_e = \nu_{e0} e^{-\frac{(h-h_0)}{H}}$$

$$n_e = n_{e0} e^{-\frac{(h-h_0)}{H}}$$



# Poynting Vector

$$\vec{S} = \frac{Mq_e^3 n_{e0}^2 e^{-\frac{(h-h_0)}{H}}}{m_e v_{e0} r^3} \left[ (v_\phi \sin \theta) \hat{r} - [(2v_\phi \cos \theta) \hat{\theta}] + [(2v_\theta \cos \theta - v_r \sin \theta) \hat{\phi}] \right]$$

$$S_r = \frac{Mq_e^3 v_\phi n_{e0}^2 \sin \theta e^{-\frac{(h-h_0)}{H}}}{m_e v_{e0} r^3} \hat{r}$$

# Drift Velocity and Kinetic Energy

$$v_{\phi} = -\frac{3cW}{q_e B_0 r} \hat{\phi}$$

35%  $\epsilon$

$$\epsilon = u_{sw} B_{IMF}^2 l_0^2 \sin^4 \left( \frac{\theta}{2} \right),$$

Strangeway, R. J., Johnson, R. G. (1983)

Wang, C., et al. (2014)

Kamide, Y., et al. (1998)

Akasofu, S. I. (1981)

Perreault, P. And Akasofu, S. I. (1978)

# Epsilon Parameter and Kp

$$\epsilon = u_{sw} B_{IMF}^2 l_0^2 \sin^4 \left( \frac{\theta}{2} \right)$$

$$n_{sw} = 7/cm^3$$

$$u_{sw} = 450 km/s$$

$$\epsilon \sim \begin{cases} (9.3 - 3.7Kp) \times 10^4 & Kp \leq 2 \\ (3.7Kp - 9.3) \times 10^4 & Kp > 2 \end{cases}$$

Goncharova, M., et al. (2000)

Anil Bhardwaj, et al. (2014)

# Kinetic Energy and Drift Velocity

$$1 \text{ erg} = 6.242 \times 10^{11} \text{ eV}$$

$$W = 2.23\varepsilon \times 10^{-22}$$

$$W = |8.3Kp - 20.7| \times 10^{-18}$$

$$v_{\phi} = -\frac{3cW}{q_e B_0 r} \hat{\phi} = -\frac{10^{-17}}{\hat{\phi}} \hat{\phi}$$

# Kinetic Energy and Drift Velocity

$$S_r = -\frac{\mu_0 c q_e^2}{\pi m_e} \frac{|6.3 K p - 16| n_{e0}^2 \sin \theta e^{-\frac{(h-h_0)}{H}}}{v_{e0} r^4} \times 10^{-18} \hat{r}$$

$$-\ln |S_r| = \ln \left( \frac{\mu_0 c q_e^2}{\pi m_e R_E^4} \right) + \ln \left( \frac{|6.3 K p - 16| n_{e0}^2 \sin \theta}{v_{e0}} \right) - \frac{(h-h_0)}{H} - 18 \ln(10)$$

# Number Densities

$$n_e(D) = 1.43 \times 10^7 e^{-0.15h' + (\beta - 0.15)(h - h')}$$

$$h' = 74.4 - 8.1 \cos \chi + 5.8 \cos \theta - 1.2 \cos \phi - 6\lambda$$

$$\beta = 0.5 - 0.2 \cos \chi - 0.1 \cos \phi + 0.1\lambda$$

Wait and Spies (1964)

Thomson (1993)

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Wait and Spies (1964)

Thomson (1993)

$$n_e(h) = (NmE) e^{\frac{(h-(hmE))}{H}}$$

$$(NmE) = 1.7 \times 10^4 (F10.7) (\cos \chi)$$

Davies (1990)

Nava et al., (2008)



$$n_e(h) = (NmF1) e^{1-z-e^z}$$

$$z = \left( \frac{h - hmF1}{H} \right)$$

$$NmF1 = 10^4 \times (24.1 + 1.24R12^2 \times 10^{-4} + 0.1R12)$$

$$n_e(h) = (NmF2) e^{\frac{1}{2}(1-z-e^z)}$$

$$z = \left( \frac{h - hmF2}{H} \right)$$

$$H = 70km \text{ and } hmF2 = 350km$$

$$foF1 = 1.4foF2$$

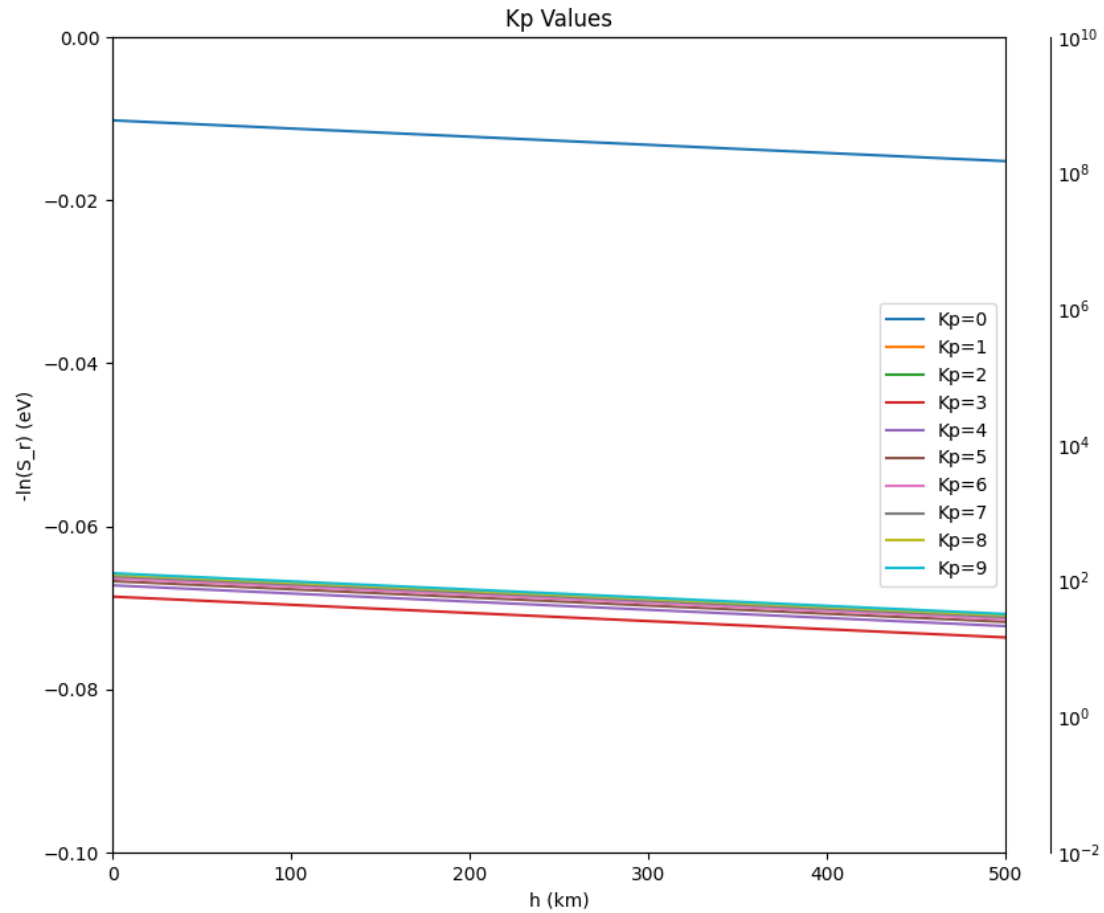
Jakowski, (2005)

Hoque and Jakowski, 2008

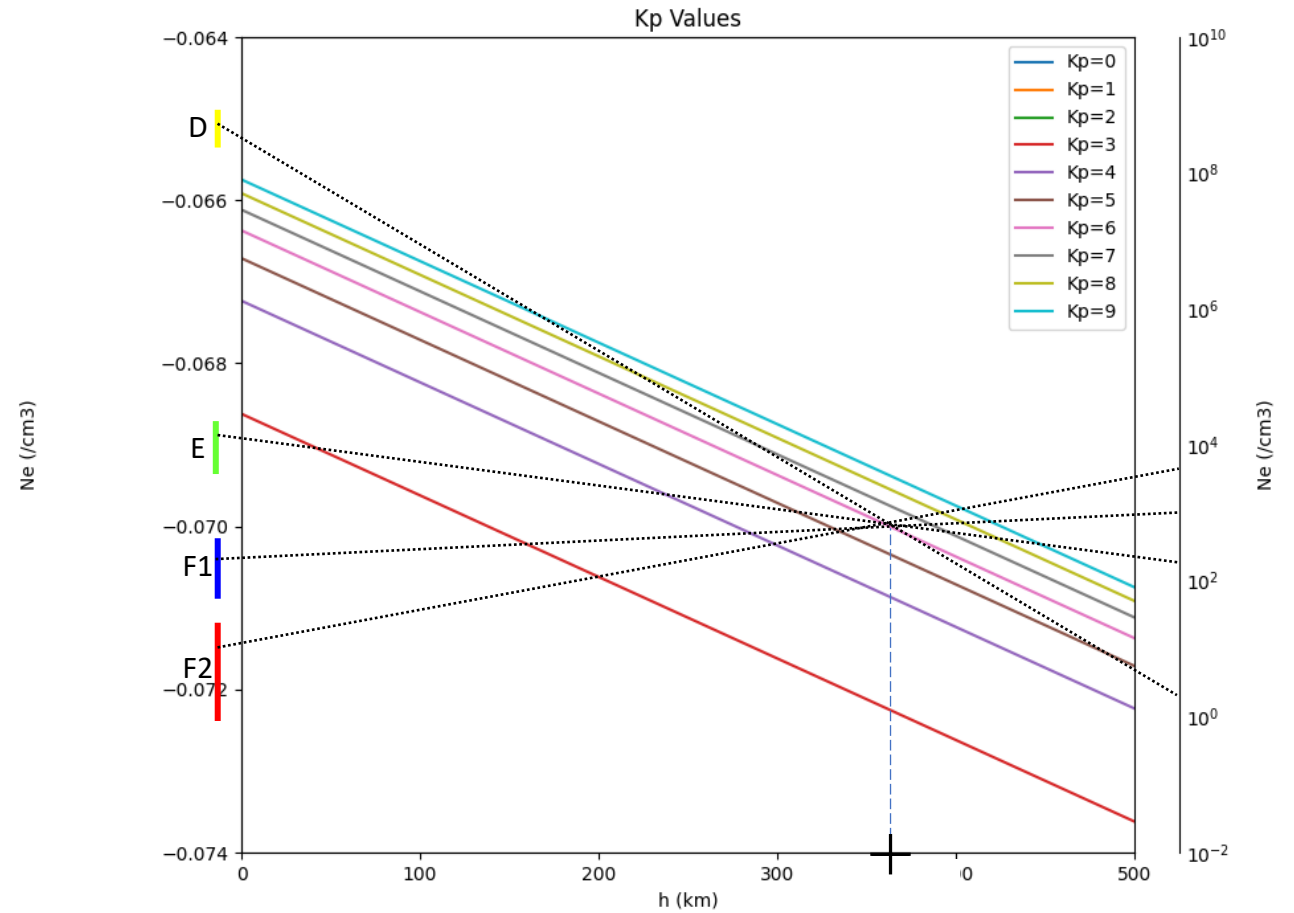
Leitinger et al., 2005

# Draft Diagrams

Novel Ionospheric Diagram for theta = 30

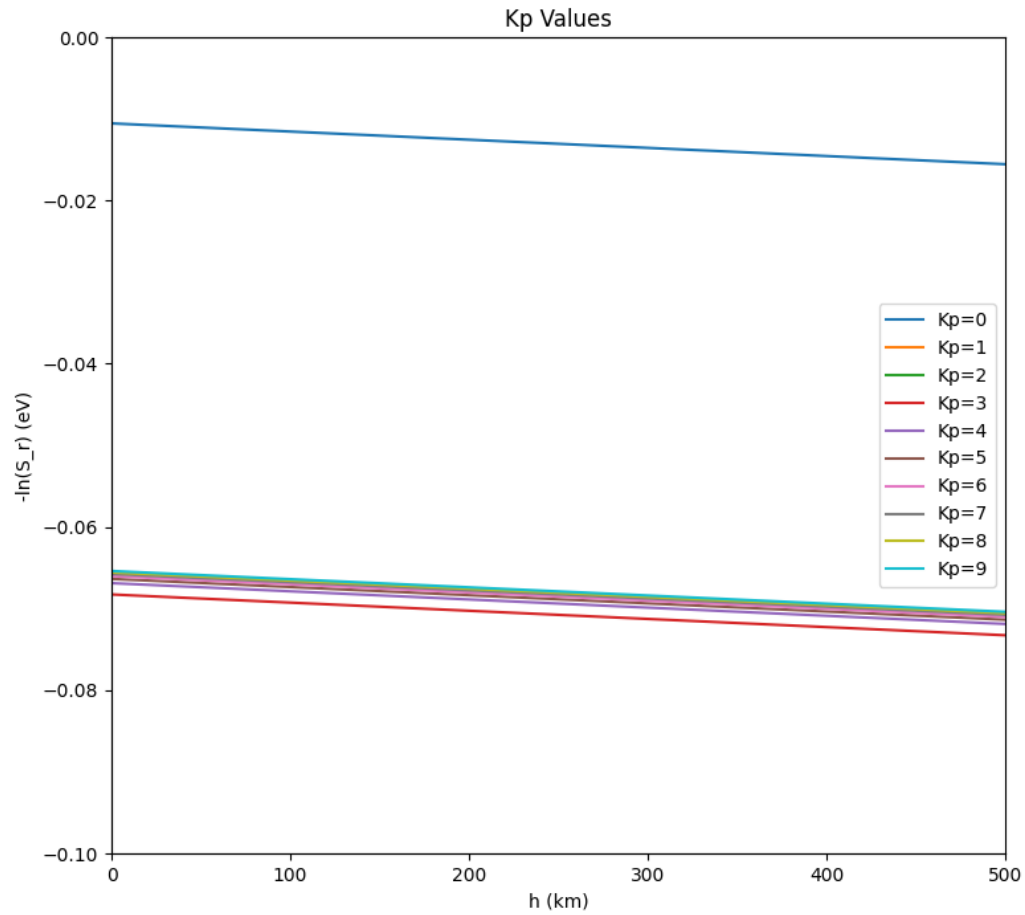


Novel Ionospheric Diagram for theta = 30

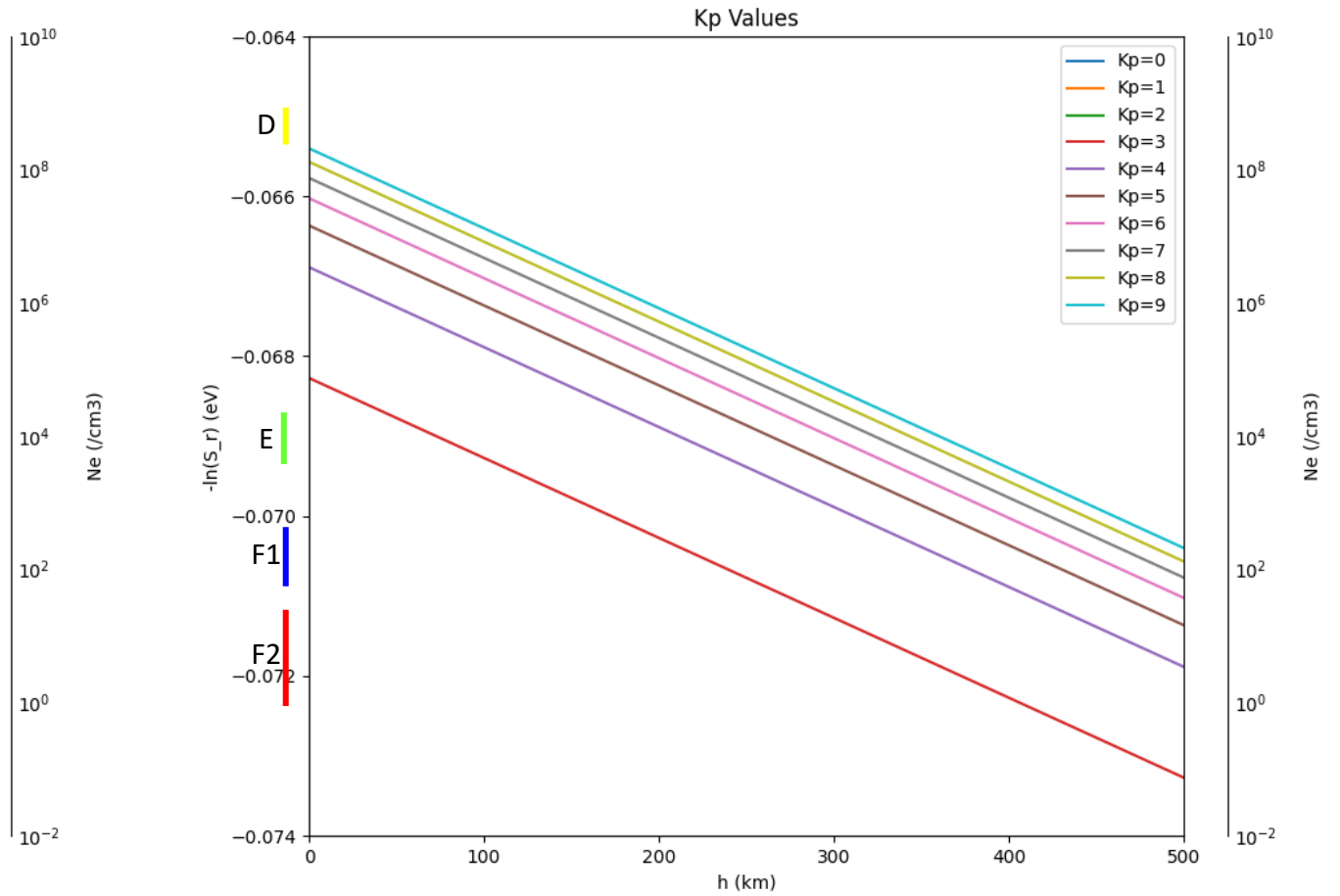


# Draft Diagrams

Novel Ionospheric Diagram for theta = 45

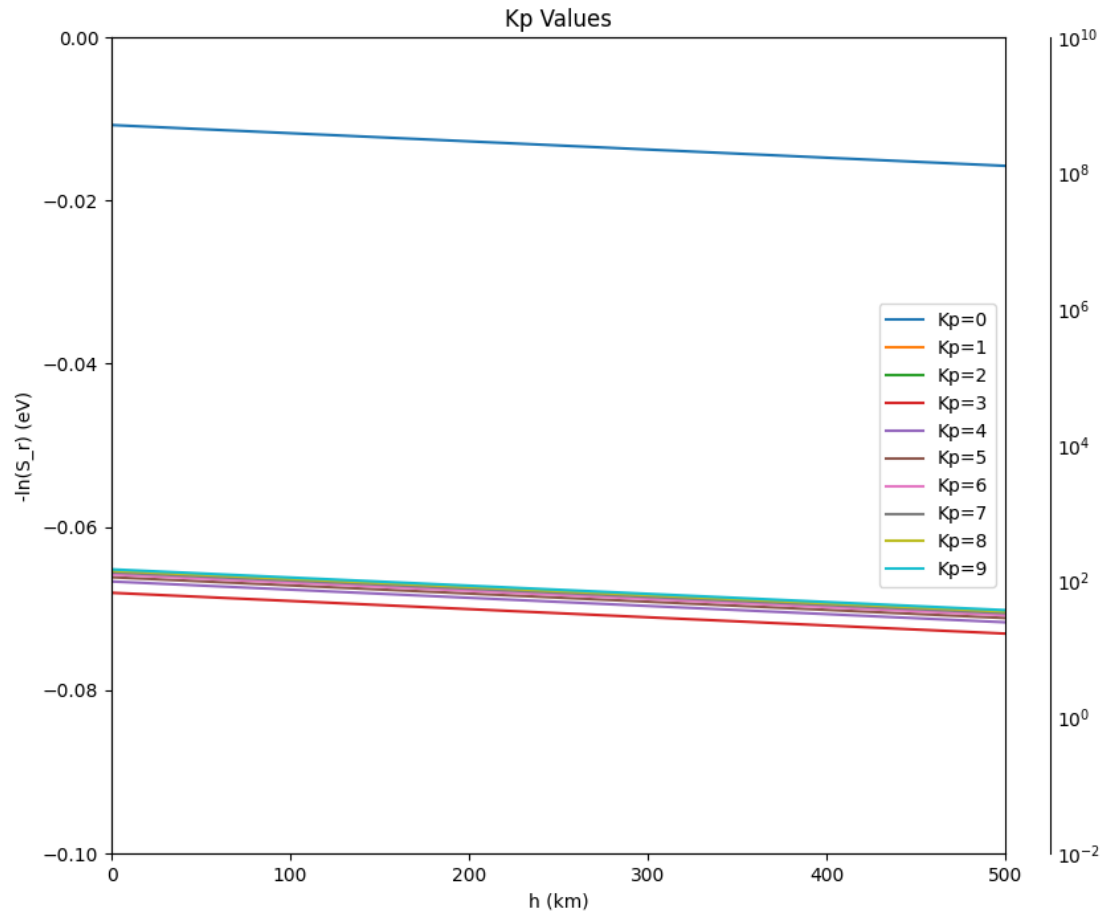


Novel Ionospheric Diagram for theta = 45

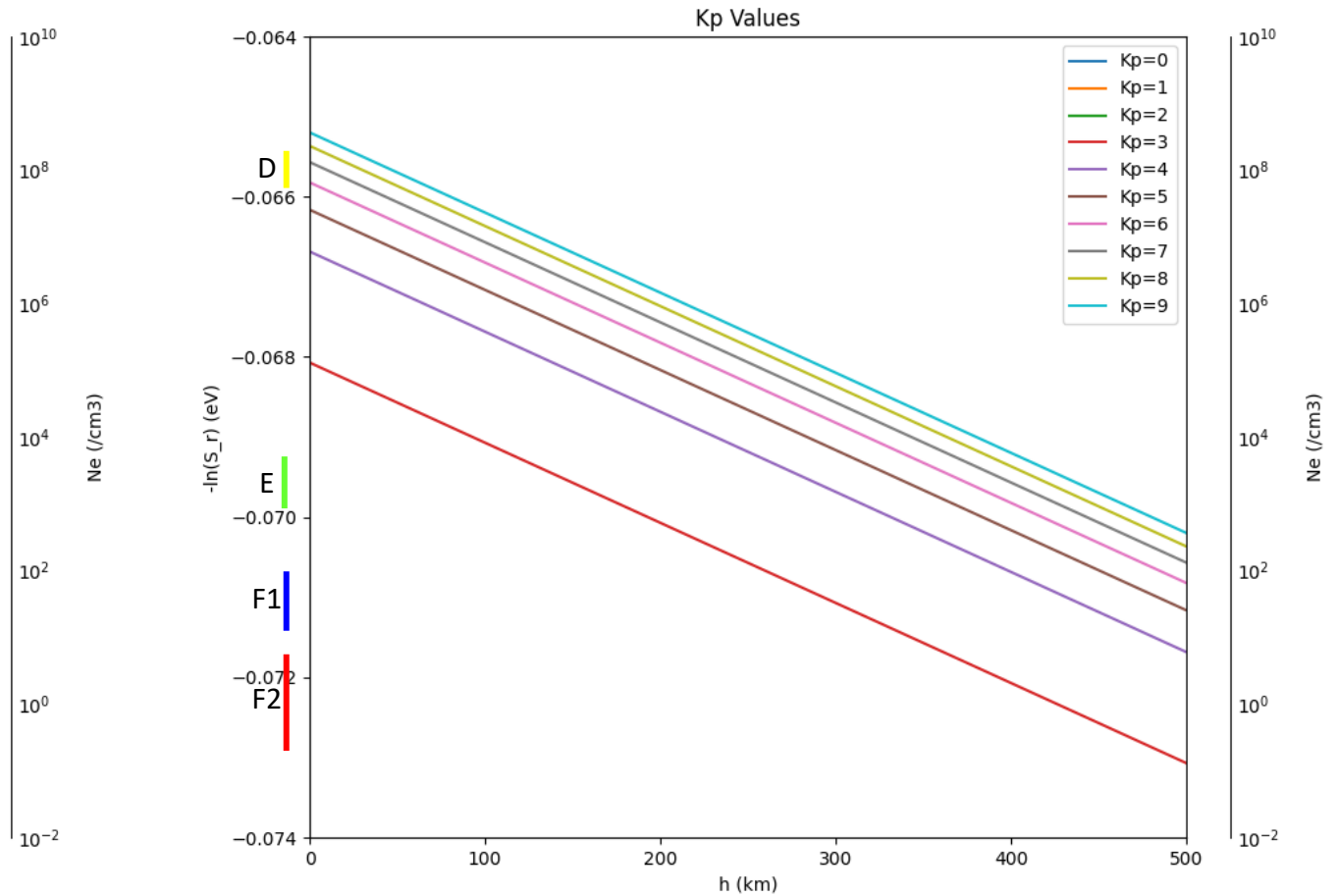


# Draft Diagrams

Novel Ionospheric Diagram for theta = 60

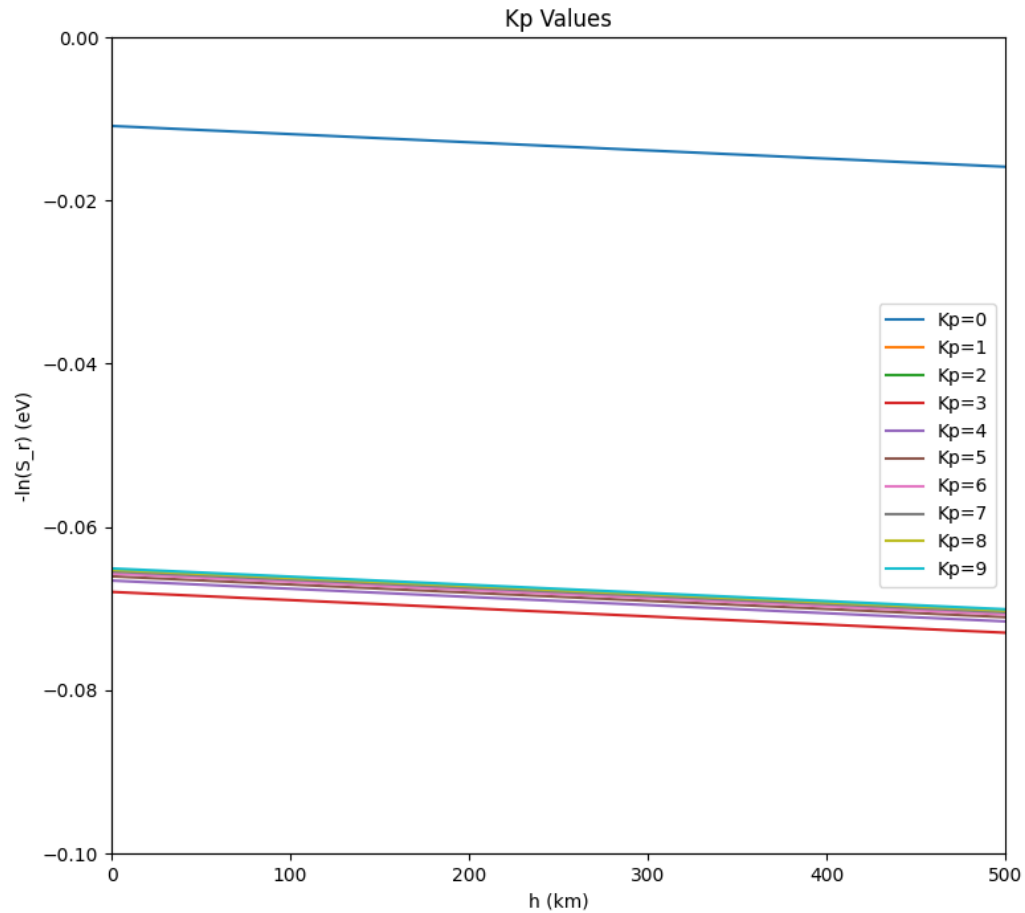


Novel Ionospheric Diagram for theta = 60

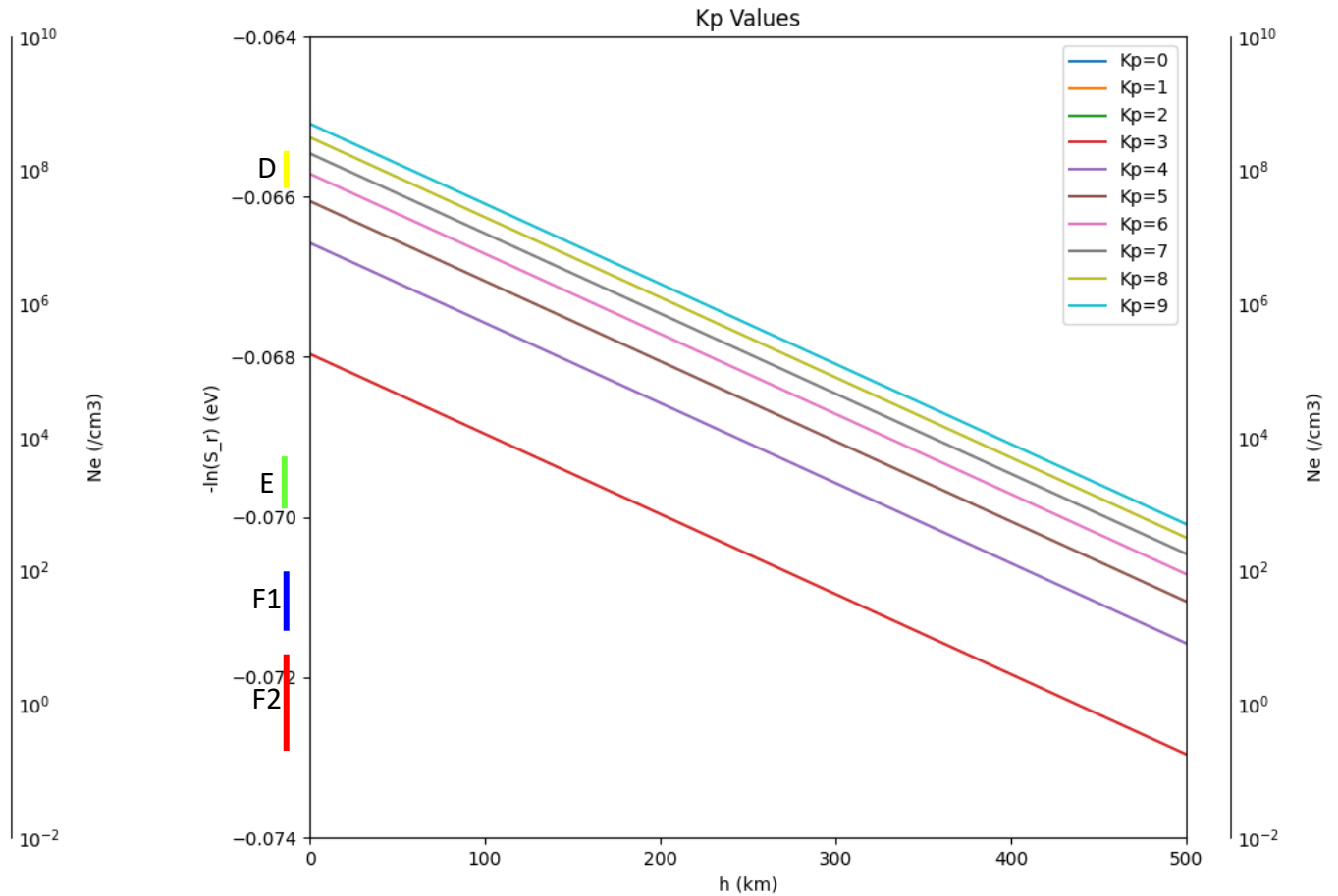


# Draft Diagrams

Novel Ionospheric Diagram for theta = 75

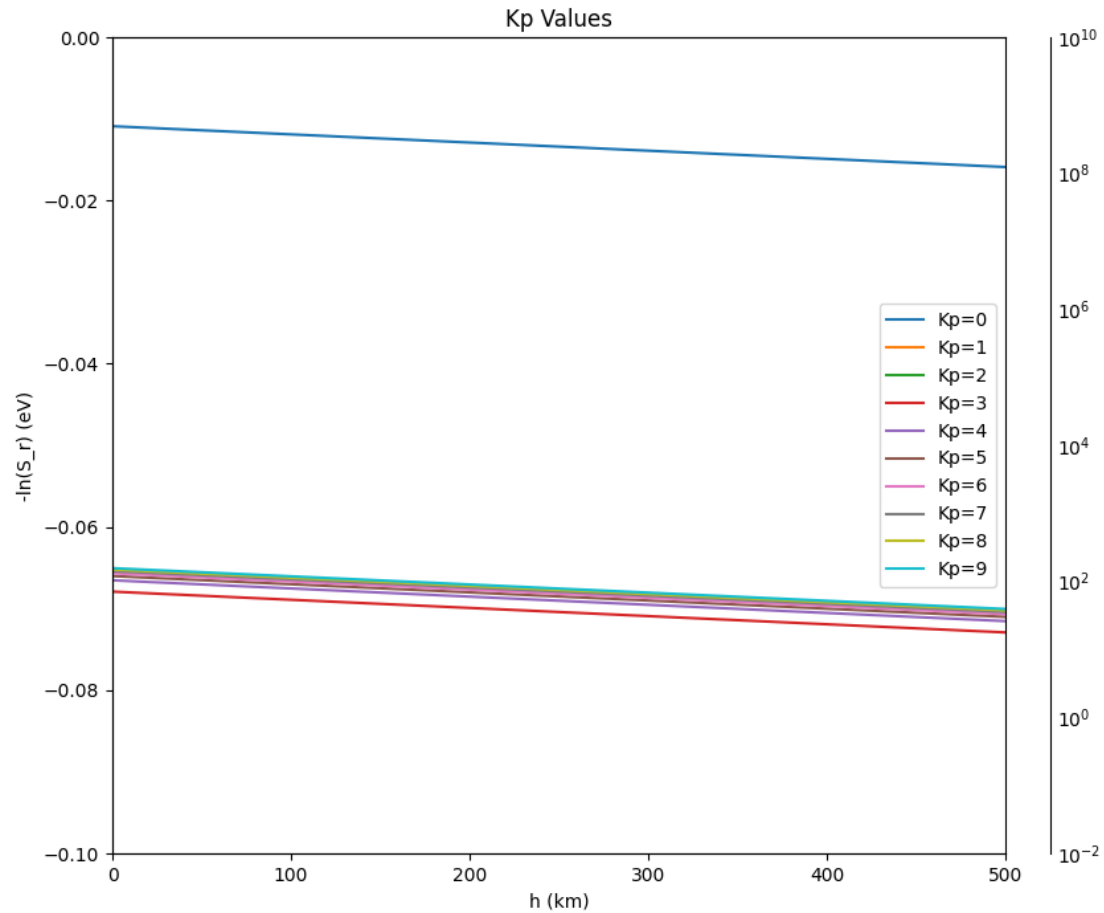


Novel Ionospheric Diagram for theta = 75

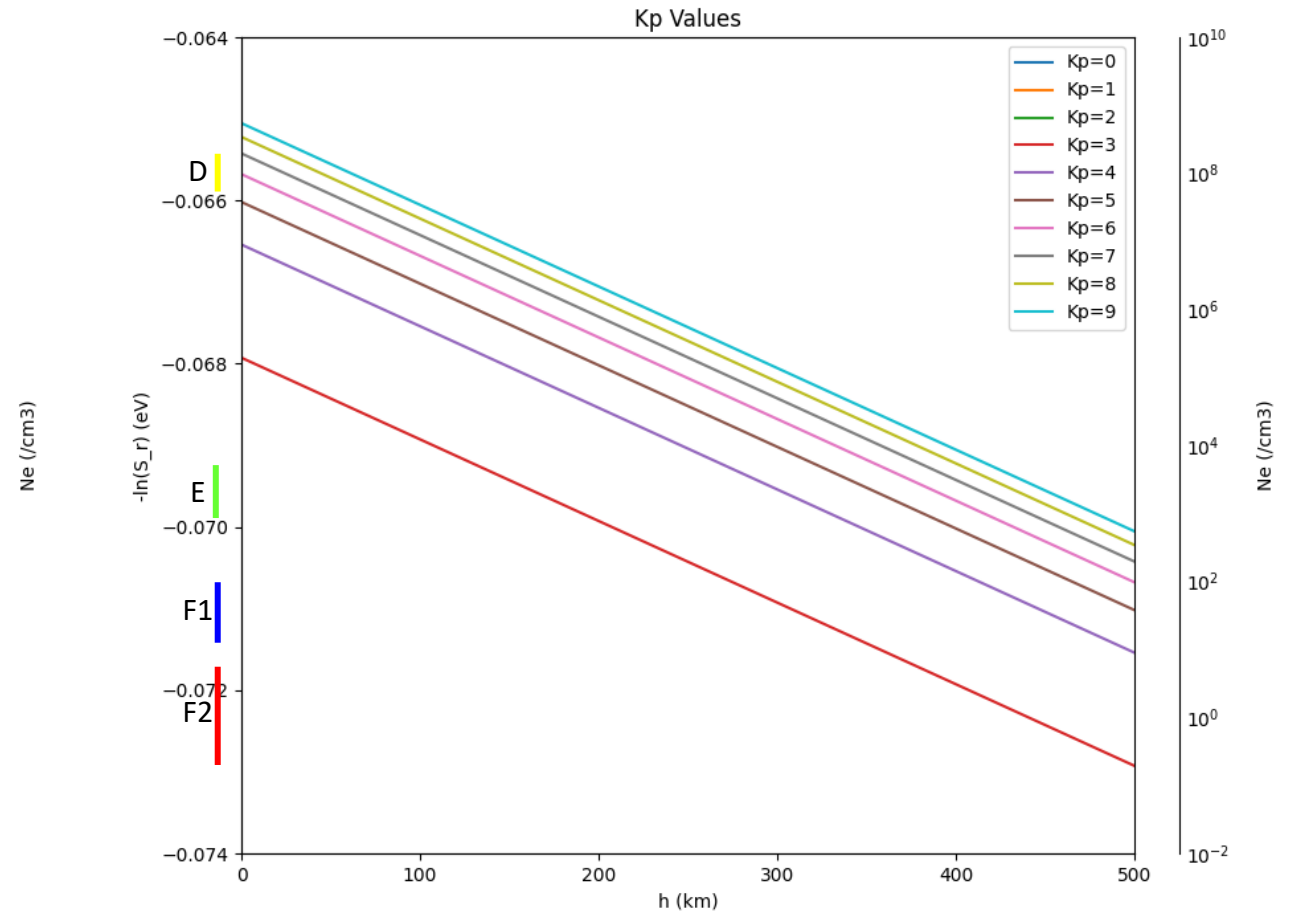


# Draft Diagrams

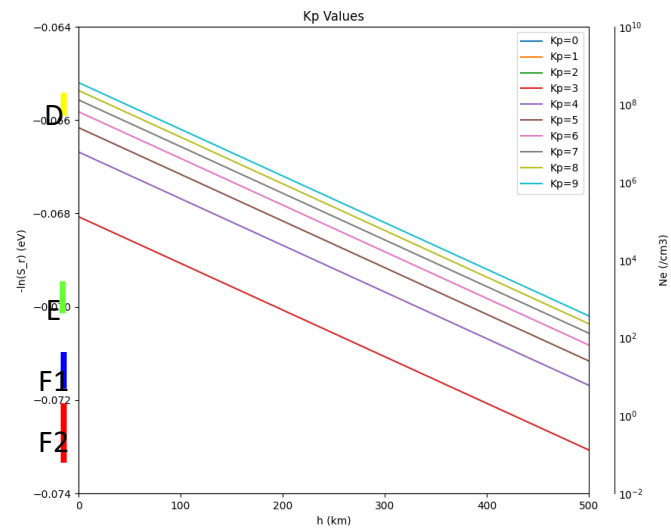
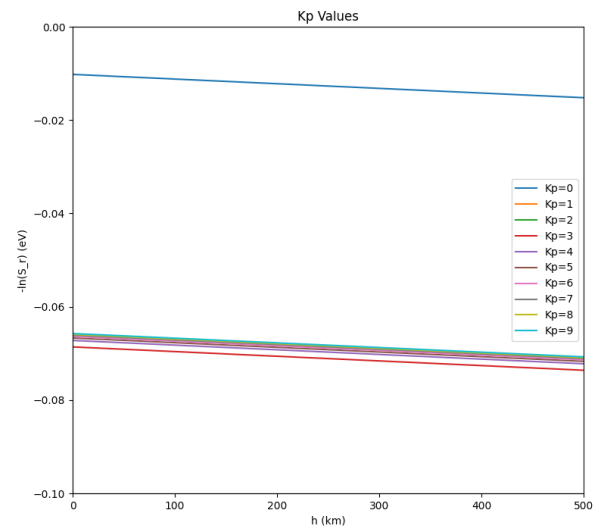
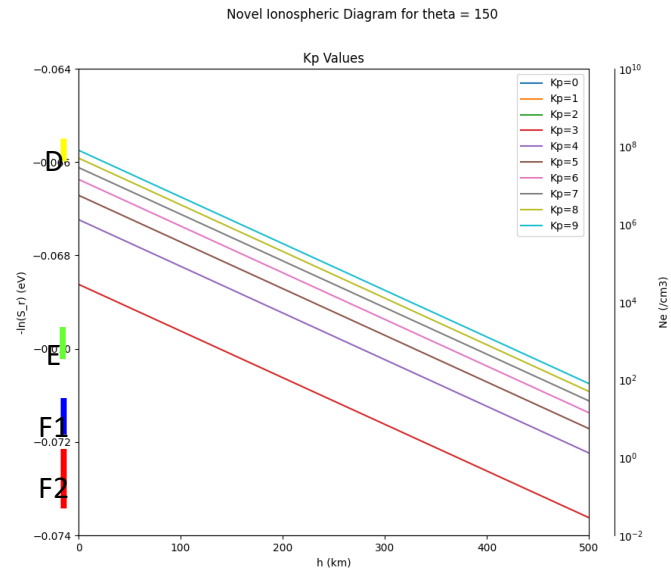
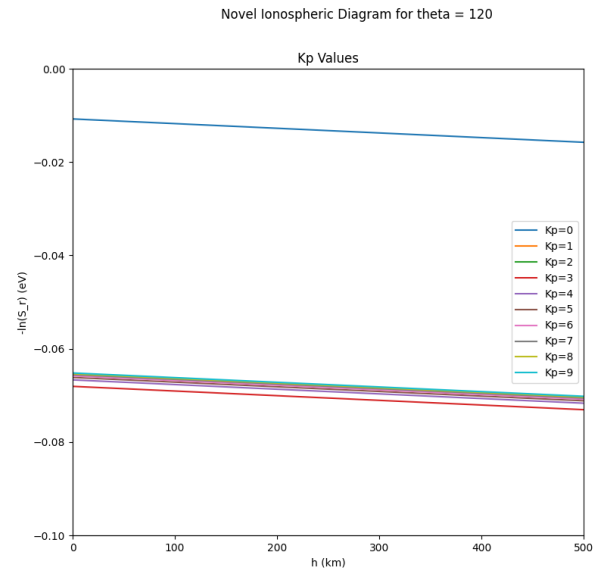
Novel Ionospheric Diagram for theta = 90



Novel Ionospheric Diagram for theta = 90



# Draft Diagrams





# Acknowledgements

We would like to express our gratitude to Dr. Y. Bayrak for his invaluable assistance with the Ellingham diagram, which has been a key inspiration for this study. We would also like to thank, B. Öztürk and A. G. Çim, for their insightful comments and valuable contributions to the literature review process. We would also like to express our gratitude to the UN Office for Outer Space Affairs for providing us with a travel grant, which has enabled us to attend this conference and share our research findings with the international scientific community.

Thank you for your attention, and waiting for your critics...