

International Space Weather Initiative: The way Forward

Austria (Vienna), 26-30 June 2023

TOPIC :

Effect of Turbulence of High-Speed Solar Winds upstream of the Earth's Magnetosphere: Case of the Outer Minima of Solar Cycles 20, 21, 22, 23 and 24.

Presented by

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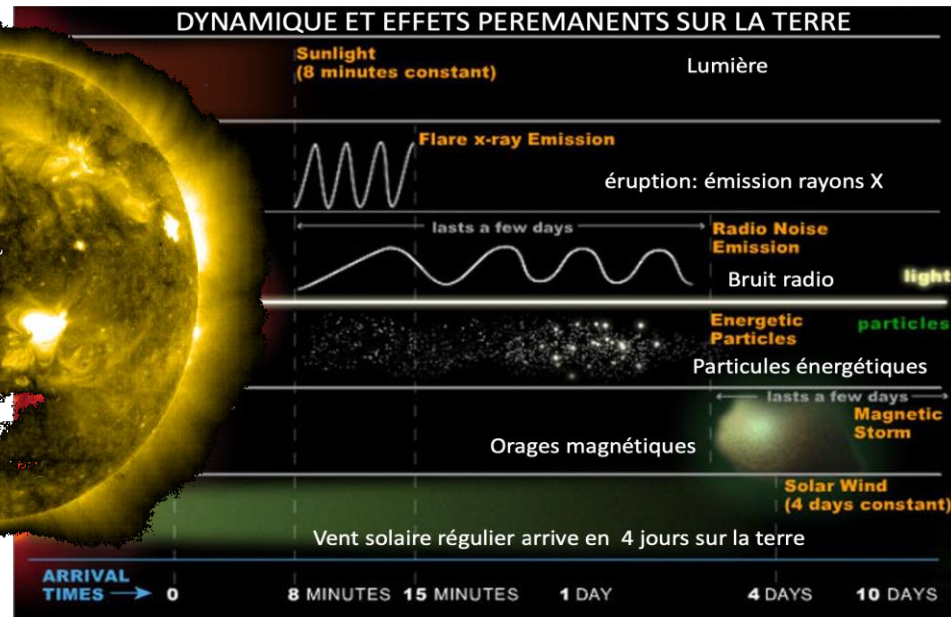
Generalities, Earth-Sun relationship

Data and methodology

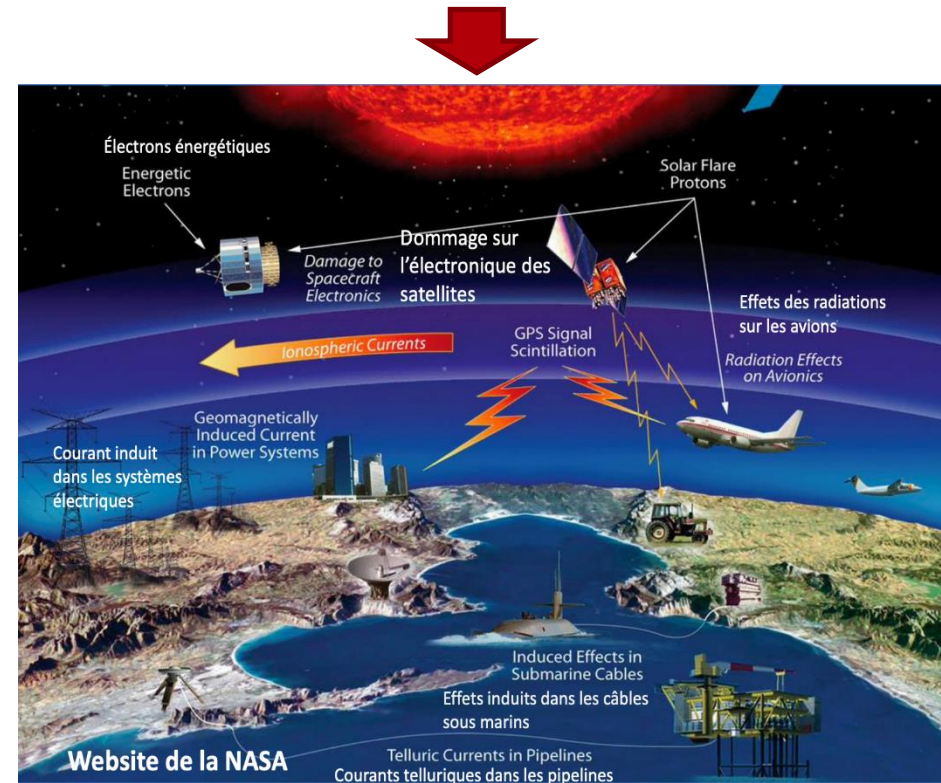
Results and discussion

Conclusion

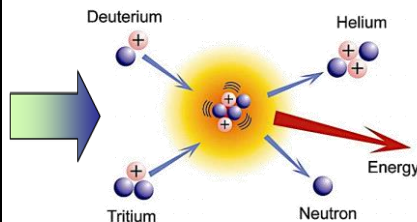
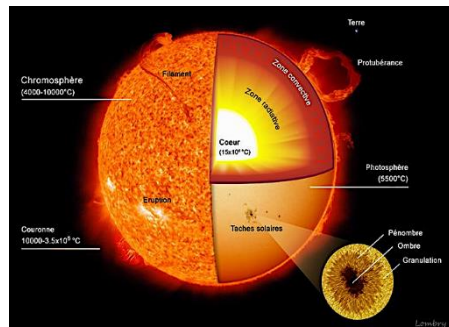
THE SUN: EMISSIONS & IMPACTS



Impacts of solar events on the Earth.

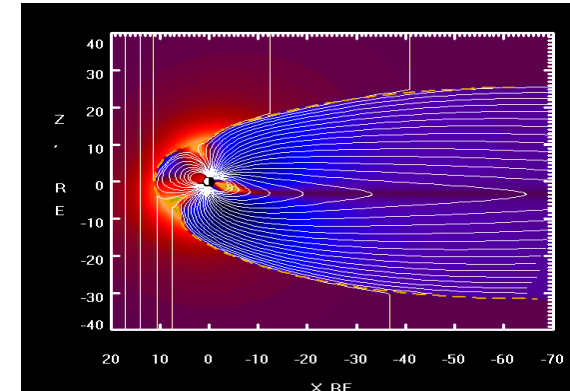
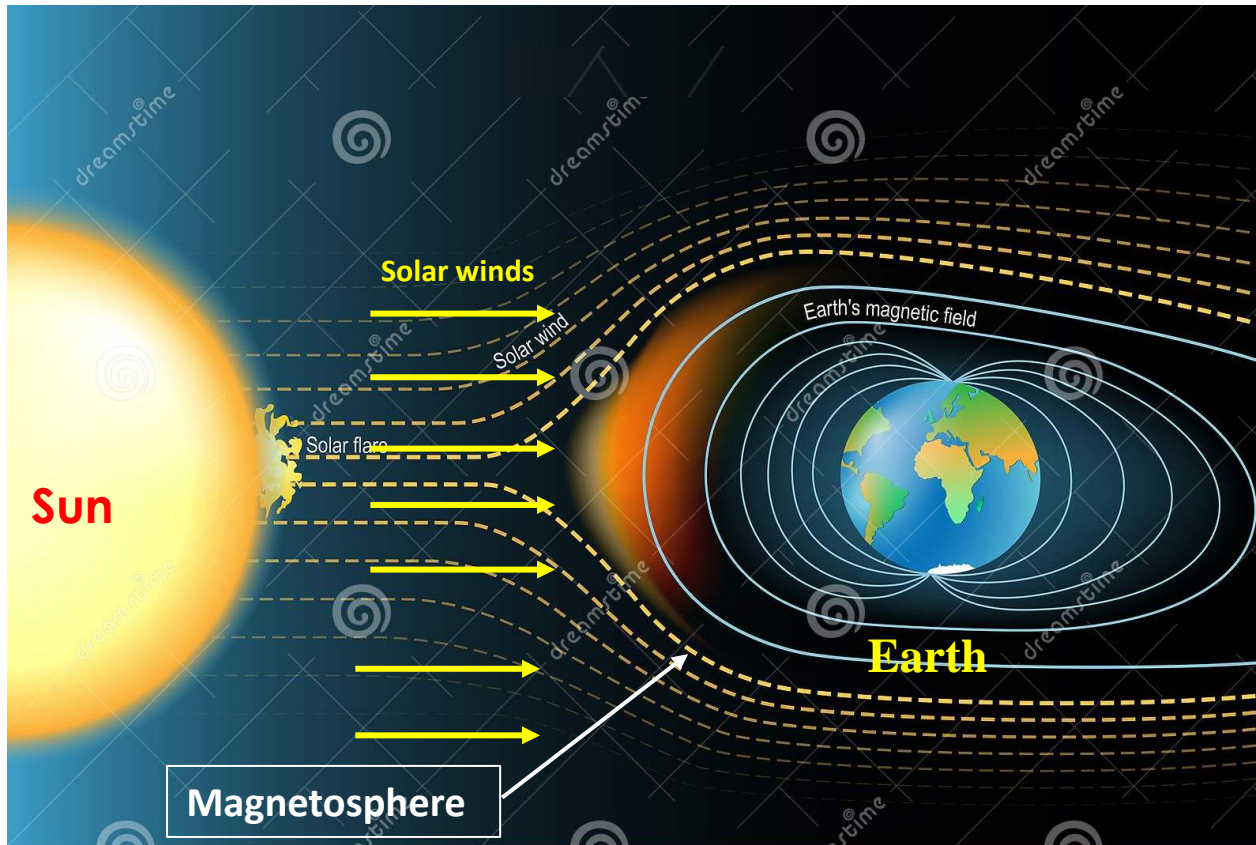


Emissions from the Sun



The energy of the Sun comes from nuclear reactions that occur at the center by fusion of H atoms into He. The average surface temperature is 5770 K. That of the nuclear core is estimated to be $\approx 15 \cdot 10^6$ °C, perhaps a little more. The Sun loses about 10^9 kg of plasma per second.

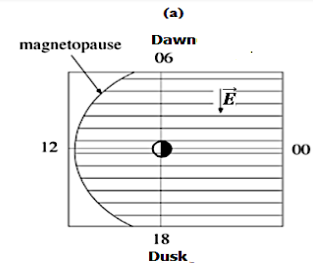
■ THE EARTH'S MAGNETOSPHERE :



SW movement is converted into electrical energy.



Total electric field : $E = E_M + E_{cor}$



Magnetosphere :
Earth shield against solar energy ramparts, controlled by the Earth **B**.



OBJECTIVES OF THE STUDY



Laboratoire de Recherche en Énergétique et Météorologie de l'Espace

■ GENERAL OBJECTIVE:

- Understanding the dynamics and structure of the Earth's magnetosphere via High-speed solar winds (HSSW) during the outer minima (descending phases) of solar cycles (SC) 20 to 24.

■ SPECIFIC OBJECTIVES:

- Extracting HSSW from 1964 to 2019 \mapsto **05 SC** (20, 21, 22, 23 & 24).
- Determine the outer minima of SC 20-24 from the Wolf number R_z
- Quantify solar fields and power upstream of the Earth's magnetosphere during the 05 peaks.



■ DATA :

- ❖ **SIDC** : «<https://www.sidc.be/SILSO/datafiles>» \mapsto Rz
- ❖ **ISGI** : «http://isgi.unistra.fr/oi_data_download.php» \mapsto Aa
- ❖ **OMNIWeb** : «<http://omniweb.gsfc.nasa.gov/form/dx1.html>» \mapsto V_x , Ey, B, Bz, n.

■ QUANTITY CONTROLLING EARTH'S MAGNETOSPHERE:

(Wu Lei and al., 1981;
Revah and Bauer, 1982)

$$E_M = 0.13E_y + 0.09 \quad \{1\}$$

(Wang et al., 2014)

$$E_{in} = 3.78 \times 10^7 n^{0.24} V^{1.47} B_T^{0.86} \left[\sin^{2.70} \left(\frac{\theta}{2} \right) + 0.25 \right] \quad \{2\}$$

(Milan et al., 2012)

$$\Phi_D = 3.3 \times 10^5 V_x^{4/3} B_{yz} \sin^{9/2} \left(\frac{\theta}{2} \right) \quad \{3\}$$

PIXEL DIAGRAMS (BARTEL'S DIAGRAMS)

Current class

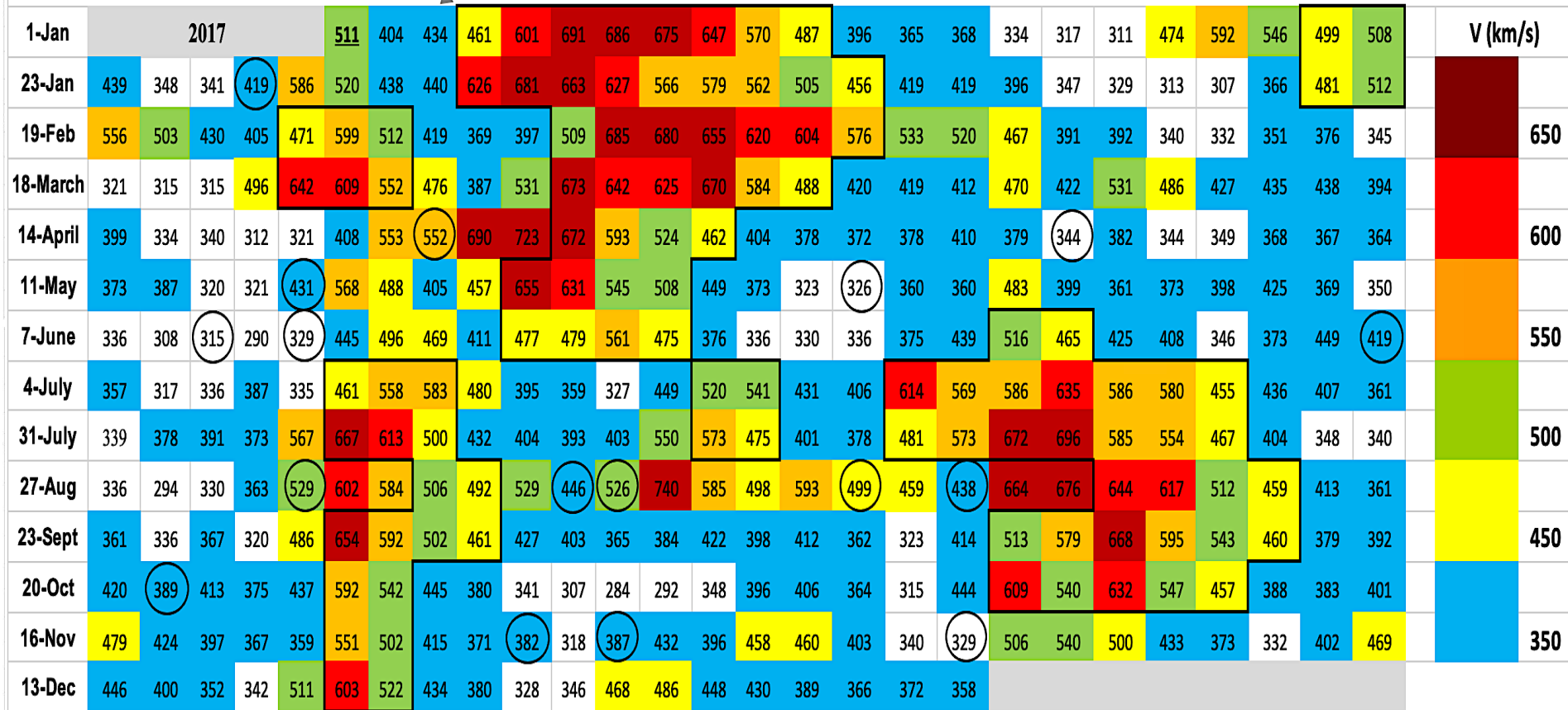


Figure 1: Pixel diagrams of 1974, 1986, 1994, 2003 and 2017 peaks

■ Solar flux structure and magnetospheric activities

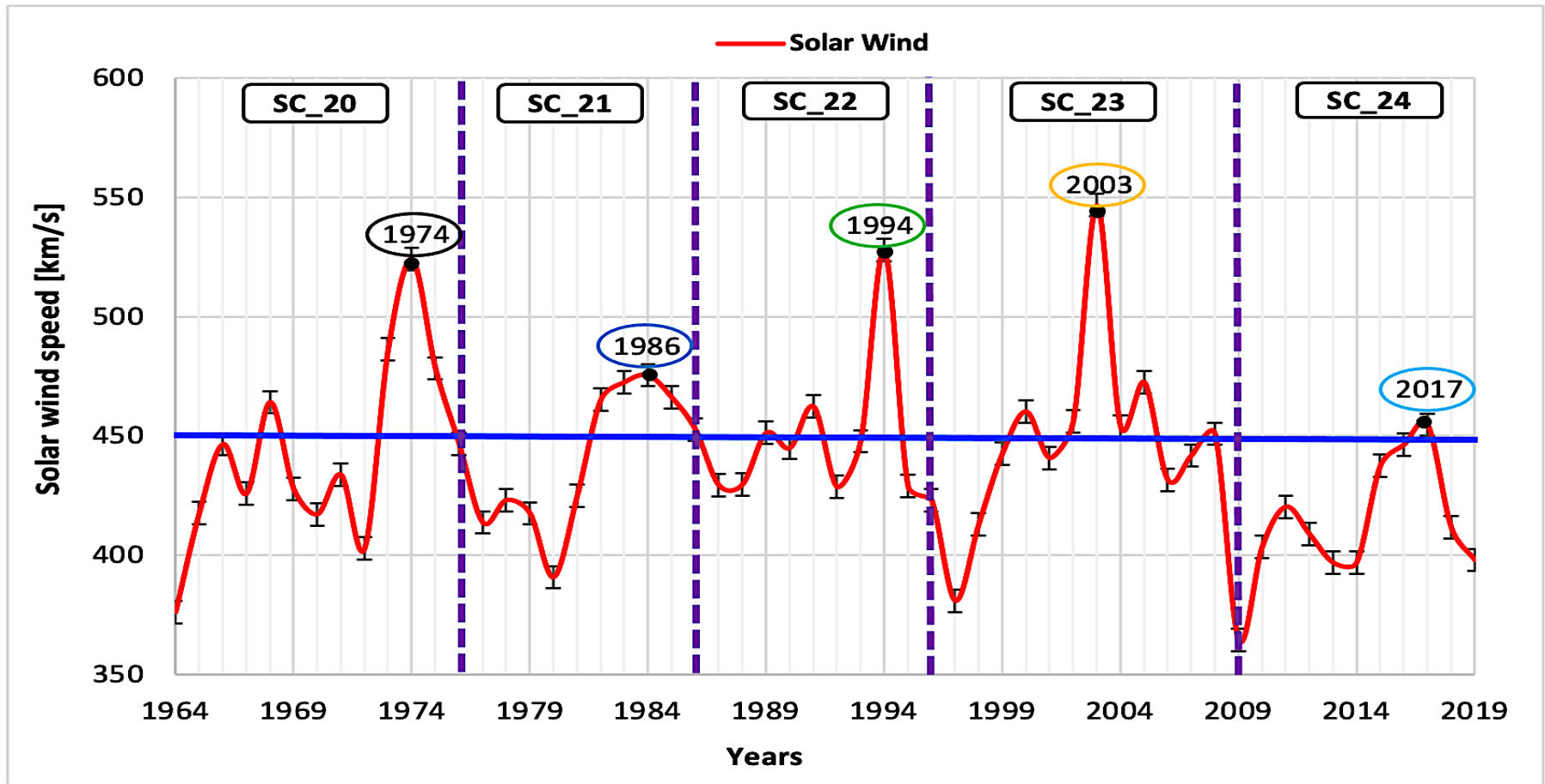


Figure 2: Annual evolution of the daily average solar wind speed from 1964-2019.

■ Solar flux structure and magnetospheric activities

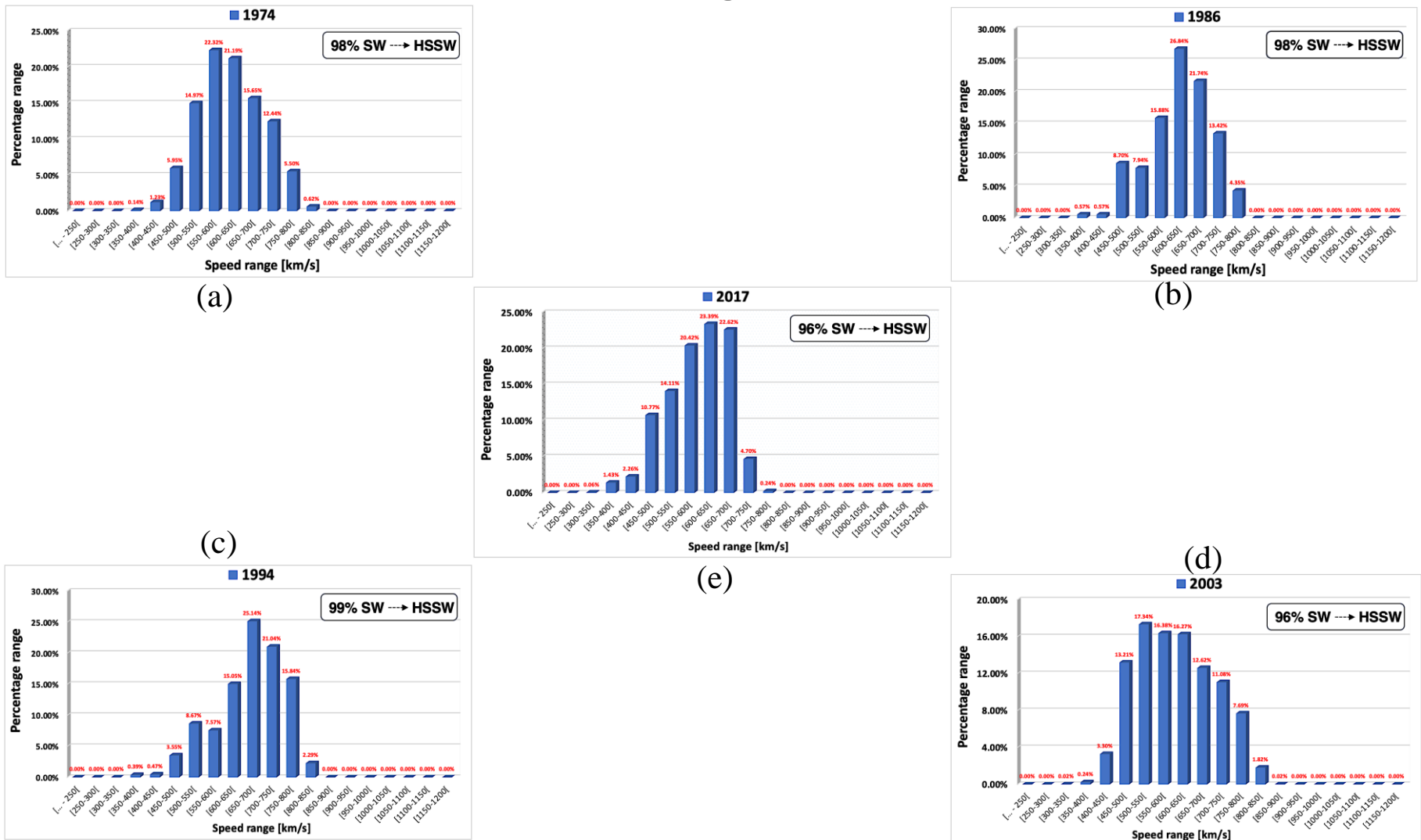


Figure 3: Annual evolution of the daily average solar wind speed from 1964-2019.

■ Solar flux structure and magnetospheric activities

Table 1: Summary of solar events of outer minima

	Outer Minima [year]				
	1974	1986	1994	2003	2017
Number of shocks	15	16	8	14	17
Number of recurring days	148	40	84	190	70
V_{sw} [km/s]	616.16	622.88	667.37	604.36	598.44
Dst [nT]	-16.31	-22.88	-32.70	-25.07	-19.96
Aa [nT]	37.27	33.03	45.42	41.12	34.06
By [nT]	-0.72	-1.15	-0.40	+0.12	+0.33
Bz	South	North	South	South	South

<https://www.sidc.be/SILSO/monthlyhemisphericplot>

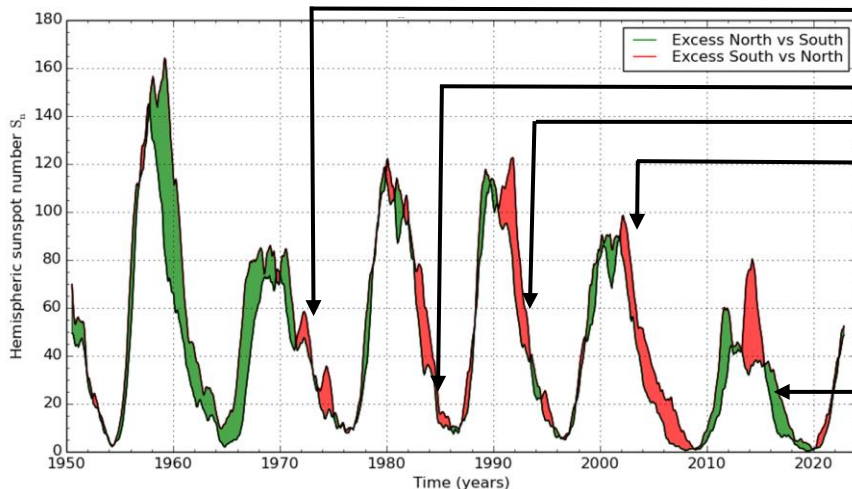


Figure 4: North-South fluctuations in sunspot numbers

■ Geoeffectiveness of the outer minimum of SC 20-24.

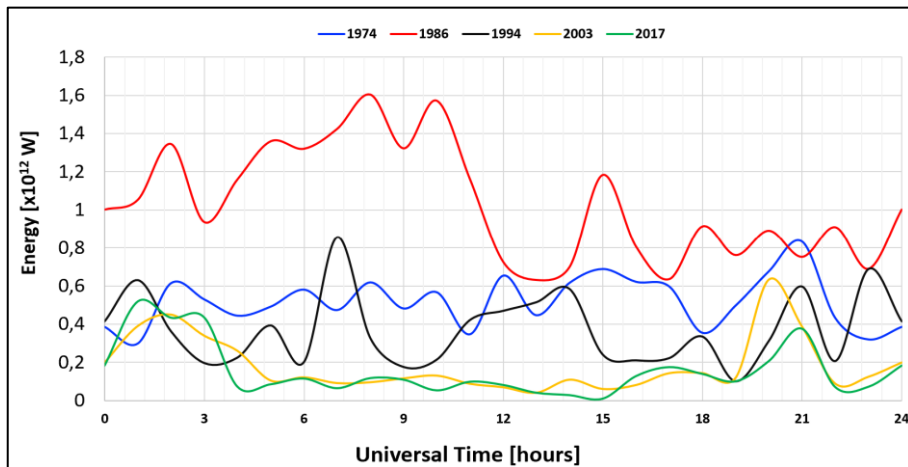


Figure 5: Power upstream of the Earth's magnetosphere

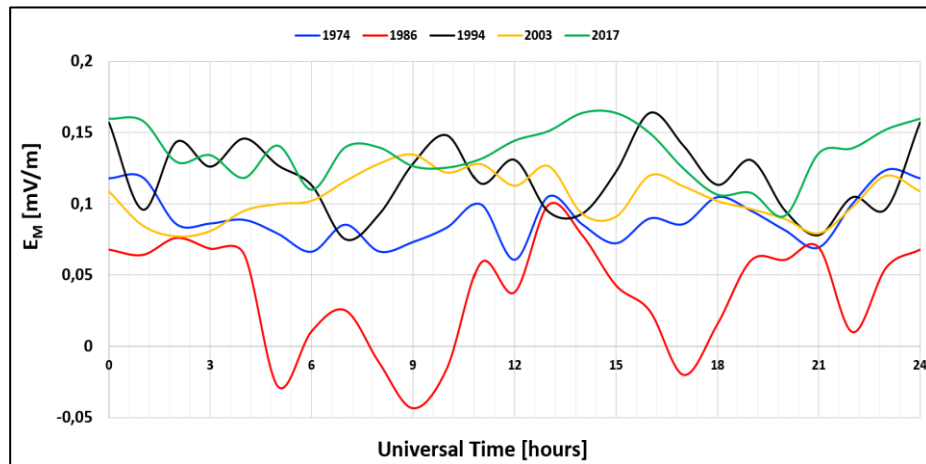


Figure 6: Daily variation in E_M field

Table 2: Correlation of some solar wind parameters

	Outer Minima [year]				
	1974	1986	1994	2003	2017
B_z [nT]	-0.02	0.65	-0.33	-0.22	-0.60
E_M [mV/m]	0.09	0.04	0.12	0.10	0.12
E_{in} [$\times 10^{12}$ W]	1.58	1.52	1.88	1.70	1.65
E_{in} & E_M	60%	76%	-59%	64%	56%

■ Geoeffectiveness of the outer minimum of SC 20-24.

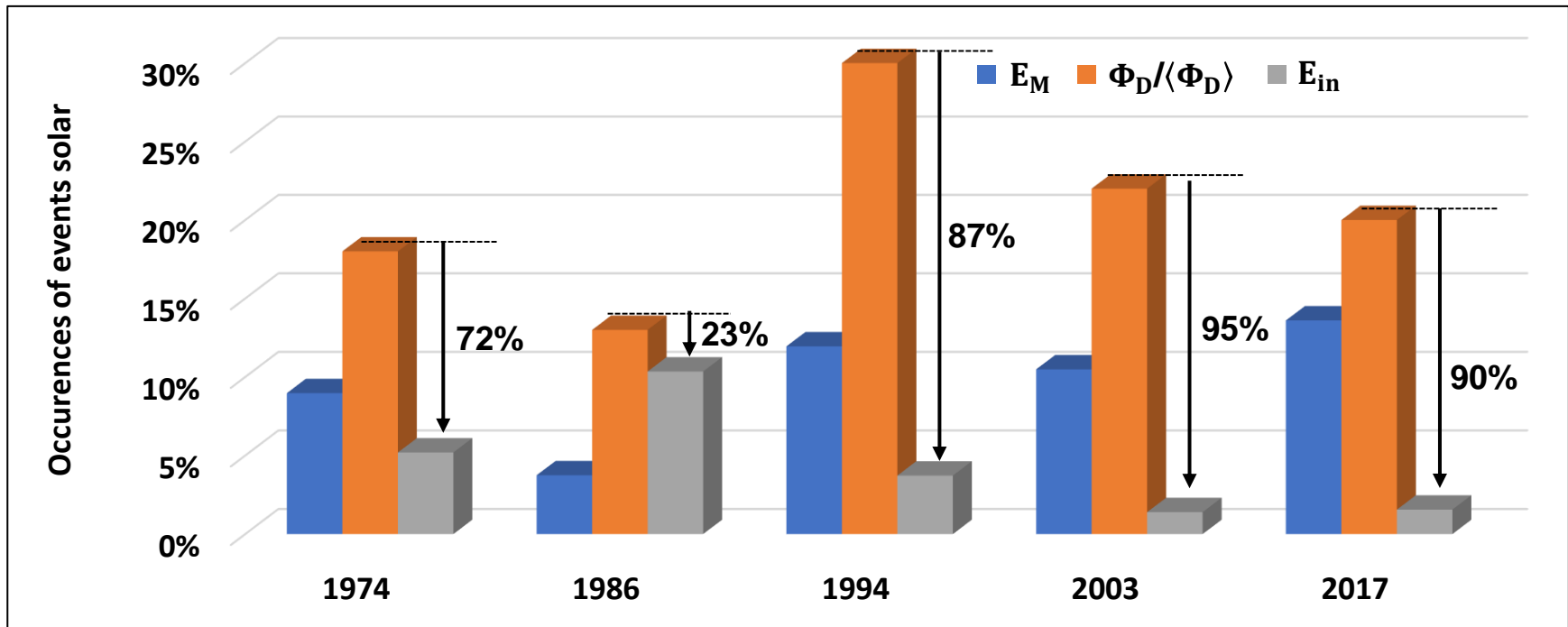
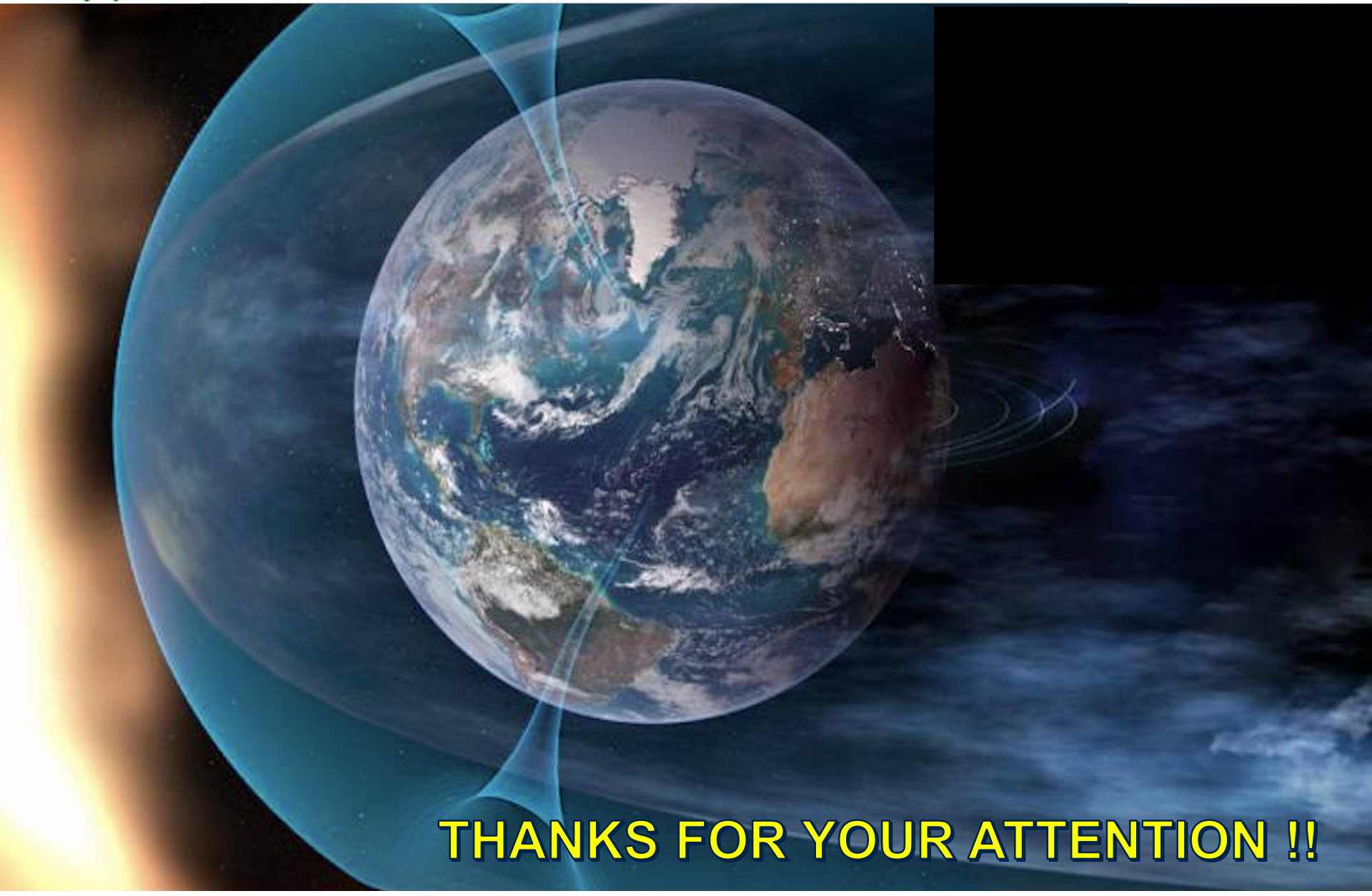


Figure 7: Occurrence E_M field, E_{in} and $\Phi_D / \langle \Phi_D \rangle$

Table 3: Correlation of some solar wind parameters.

	Outer Minima [year]				
	1974	1986	1994	2003	2017
$\Phi_D / \langle \Phi_D \rangle$	18%	13%	30%	22%	20%
E_{in} & E_M	60%	76%	-59%	64%	56%

- During the peaks of the outer minima of SC :
 - ❖ At high speeds, the normalized daytime reconnection rate is likely to be improved when the IMF-Bz is antiparallel to the geomagnetic field.
 - ❖ Significant geomagnetic activity is sometimes present even in the absence of such important ICMEs.
 - ❖ HSSW represent the stability criterion for areas that are particularly close to the outer minimum of solar cycles.
 - ❖ For high B_y intensities with a north-pointing CMI-Bz, the trapping and energization of HSSW particles in the Earth's magnetic cavity has an enhanced influence on the magnetospheric convection electric field.



THANKS FOR YOUR ATTENTION !!