## **International Space Weather Initiative: The way Forward**

Austria (Vienna), 26-30 June 2023



Effect of Turbulence of High-Speed Solar Winds upstream of the Earth's Magnetosphere: Case of the Outer Minima of Solar Cycles 20, 21, 22, 23 and 24.

Presented by

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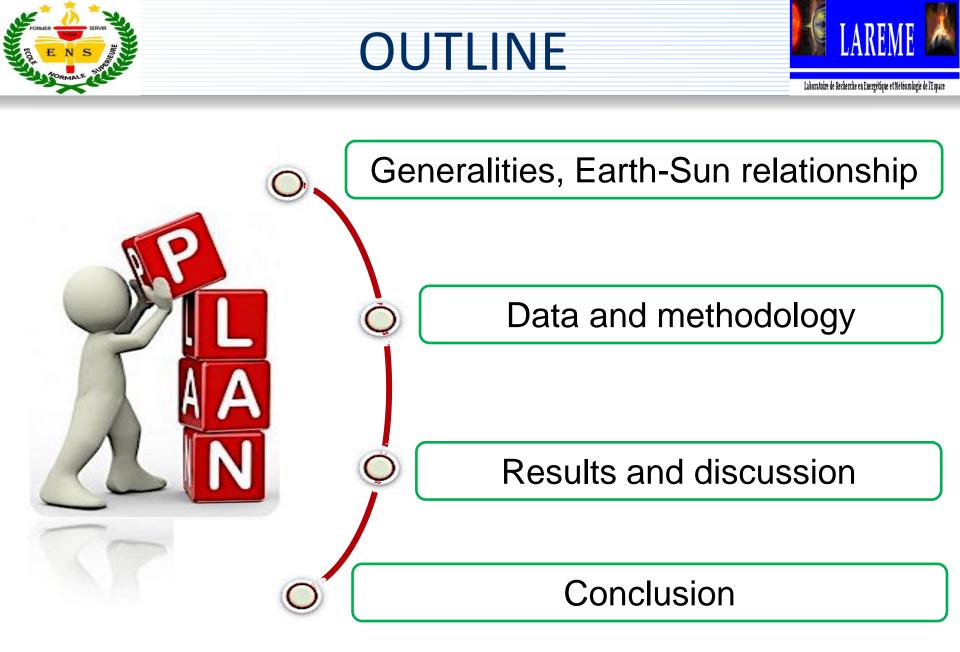
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nternational Committee on Global Navigation Satellite Systems







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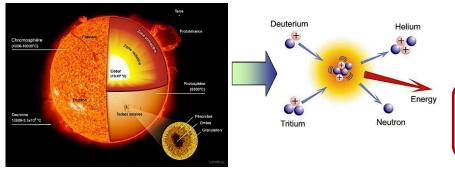
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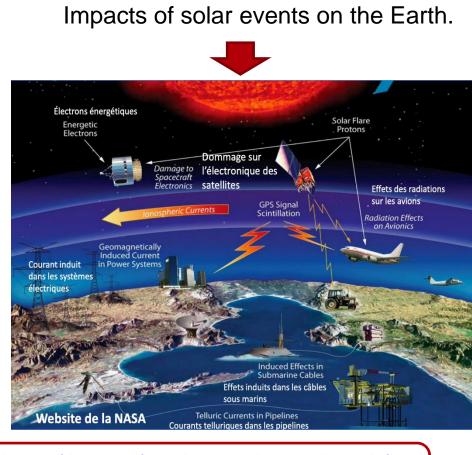
#### THE SUN: EMISSIONS & IMPACTS

**GENERALITIES** 

DYNAN	AIQUE ET EFFETS PEREN	ANENTS SUR LA TERI	RE
	Sunlight <sub>(</sub> 8 minutes constant)	Lumière	
AT A A	Flare x-ray En	iission éruption: émission ray	vons X
	lasts a few da	ys Radio Noise Emission Bruit radio	
		Energetic Particles Particules én	particles ergétiques
	Orage	s magnétiques	few days Magnetic Storm
	Vent solaire régulier arr		r Wind ays constant
	8 MINUTES 15 MINUTES	1 DAY 4 DAYS	10 DAYS

#### Emissions from the Sun





The energy of the Sun comes from nuclear reactions that occur at the center by fusion of H atoms into He. The average surface temperature is 5770 K. That of the nuclear core is estimated to be  $\approx 15.10^{6}$  °C, perhaps a little more. The Sun loses about 10<sup>9</sup> kg of plasma per second.



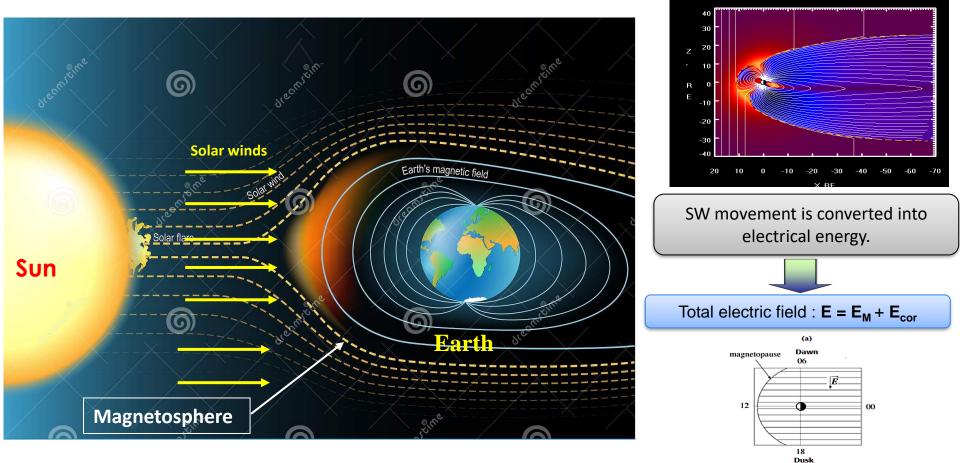
GENERALITIES



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#### THE EARTH'S MAGNETOSPHERE :



#### Magnetosphere :

Earth shield against solar energy ramparts, controlled by the Earth B.



### GENERAL OBJECTIVE:

 Understanding the dynamics and structure of the Earth's magnetosphere via High-speed solar winds (HSSW) during the outer minima (descending phases) of solar cycles (SC) 20 to 24.

### • SPECIFIC OBJECTIVES:

- Extracting HSSW from 1964 to 2019 → **05 SC** (20, 21, 22, 23 & 24).
- Determine the outer minima of SC 20-24 from the Wolf number Rz
- Quantify solar fields and power upstream of the Earth's magnetosphere during the 05 peaks.





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#### DATA :

- ♦ SIDC :  $(https://www.sidc.be/SILSO/datafiles) \mapsto Rz$
- ISGI: «<u>http://isgi.unistra.fr/oi\_data\_download.php</u>» → Aa
- ↔ OMNIWeb : «<u>http://omniweb.gsfc.nasa.gov/form/dx1.html</u>»  $\mapsto$  V<sub>x</sub>, Ey, B, Bz, n.

## QUANTITY CONTROLLING EARTH'S MAGNETOSPHERE:

(Wu Lei and al., 19 Revah and Bauer, 19		<b>{1</b> }
(Wang et al., 2014)	$E_{in} = 3.78 \times 10^7 n^{0.24} V^{1.47} B_T^{0.86} \left[ \sin^{2.70} \left( \frac{\theta}{2} \right) + 0.25 \right]$	<b>{2</b> }
(Milan et al., 2012)	$\Phi_{\rm D} = 3.3 \times 10^5  V_x^{4/3}  \mathrm{B}_{yz} \sin^{9/2} \left(\frac{\theta}{2}\right)$	{ <mark>3</mark> }



DATA & METHODOLOGY

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### PIXEL DIAGRAMS (BARTEL'S DIAGRAMS)

				Curre	ent cl	ass																							
1-Jan			2017			<u>511</u>	404	434	461	601	691	686	675	647	570	487	396	365	368	334	317	311	474	592	546	499	508	V (km	/s)
23-Jan	439	348	341	(419)	586	520	438	440	626	681	663	627	566	579	562	505	456	419	419	396	347	329	313	307	366	481	512		
19-Feb	556	503	430	405	471	599	512	419	369	397	509	685	680	655	620	604	576	533	520	467	391	392	340	332	351	376	345		650
18-March	321	315	315	496	642	609	552	476	387	531	673	642	625	670	584	488	420	419	412	470	422	531	486	427	435	438	394		
14-April	399	334	340	312	321	408	553	552	690	723	672	593	524	462	404	378	372	378	410	379	(344)	382	344	349	368	367	364		600
11-May	373	387	320	321	(431)	568	488	405	457	655	631	545	508	449	373	323	(326)	360	360	483	399	361	373	398	425	369	350		
7-June	336	308	(315)	290	(329)	445	496	469	411	477	479	561	475	376	336	330	336	375	439	516	465	425	408	346	373	449	(419)		550
4-July	357	317	336	387	335	461	558	583	480	395	359	327	449	520	541	431	406	614	569	586	635	586	580	455	436	407	361		
31-July	339	378	391	373	567	667	613	500	432	404	393	403	550	573	475	401	378	481	573	672	696	585	554	467	404	348	340		500
27-Aug	336	294	330	363	(529)	602	584	506	492	529	(446)	(526)	740	585	498	593	(499)	459	(438)	664	676	644	617	512	459	413	361		
23-Sept	361	336	367	320	486	654	592	502	461	427	403	365	384	422	398	412	362	323	414	513	579	668	595	543	460	379	392		450
20-Oct	420	(389)	413	375	437	592	542	445	380	341	307	284	292	348	396	406	364	315	444	609	540	632	547	457	388	383	401		
16-Nov	479	424	397	367	359	551	502	415	371	382	318	(387)	432	396	458	460	403	340	(329)	506	540	500	433	373	332	402	469		350
13-Dec	446	400	352	342	511	603	522	434	380	328	346	<mark>468</mark>	486	448	430	389	366	372	358										

Figure 1: Pixel diagrams of 1974, 1986, 1994, 2003 and 2017 peaks



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#### Solar flux structure and magnetospheric activities

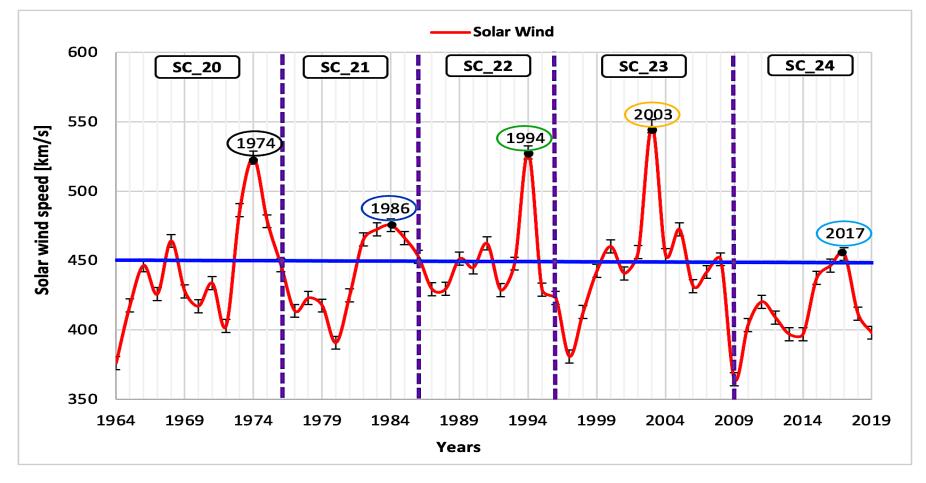


Figure 2: Annual evolution of the daily average solar wind speed from 1964-2019.



### Solar flux structure and magnetospheric activities

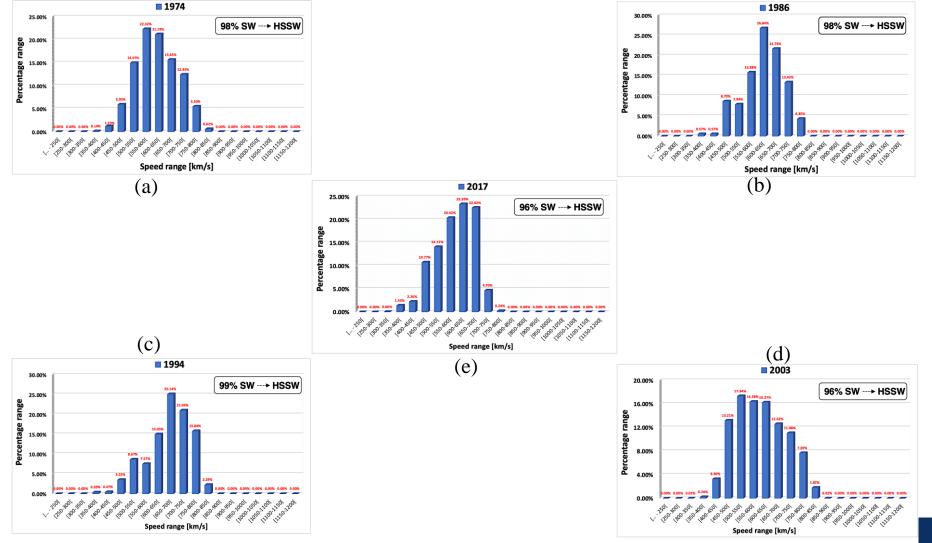


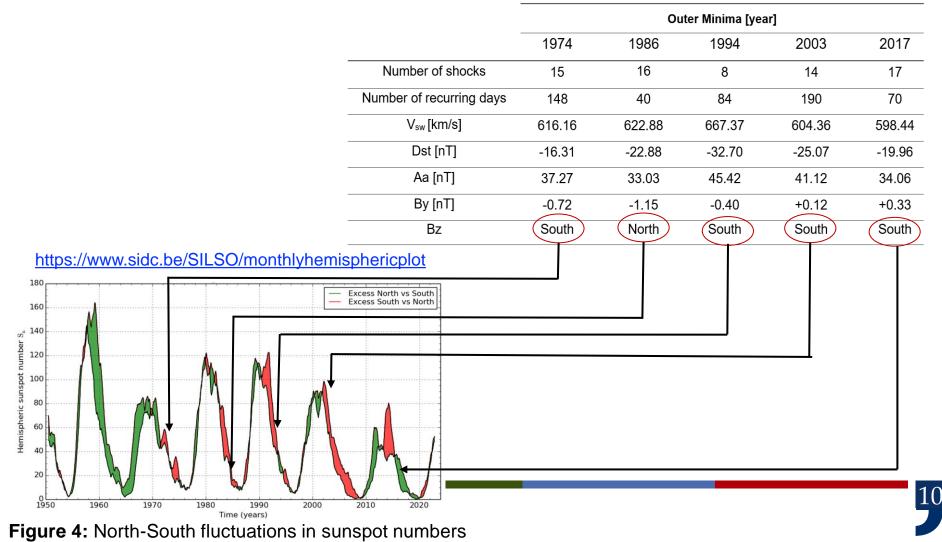
Figure 3: Annual evolution of the daily average solar wind speed from 1964-2019.





#### Solar flux structure and magnetospheric activities

Table 1: Summary of solar events of outer minima





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#### Geoeffectiveness of the outer minimum of SC 20-24.

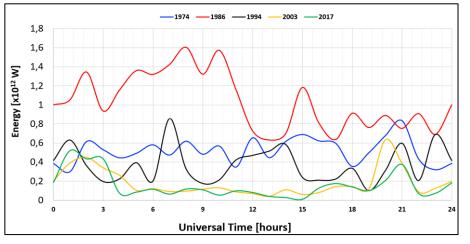


Figure 5: Power upstream of the Earth's magnetosphere

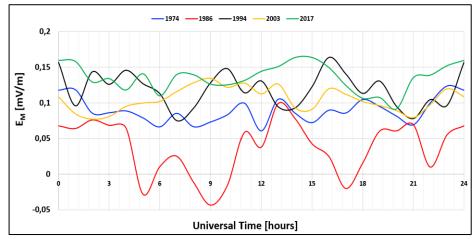


Figure 6: Daily variation in E<sub>M</sub> field

		Outer Minima [year]											
	1974	1986	1994	2003	2017								
Bz[nT]	-0.02	0.65	-0.33	-0.22	-0.60								
E <sub>M</sub> [mV/m]	0.09	0.04	0.12	0.10	0.12								
E <sub>in</sub> [x10 <sup>12</sup> W]	1.58	1.52	1.88	1.70	1.65								
Ein & EM	60%	76%	-59%	64%	56%								

#### Table 2: Correlation of some solar wind parameters





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#### Geoeffectiveness of the outer minimum of SC 20-24.

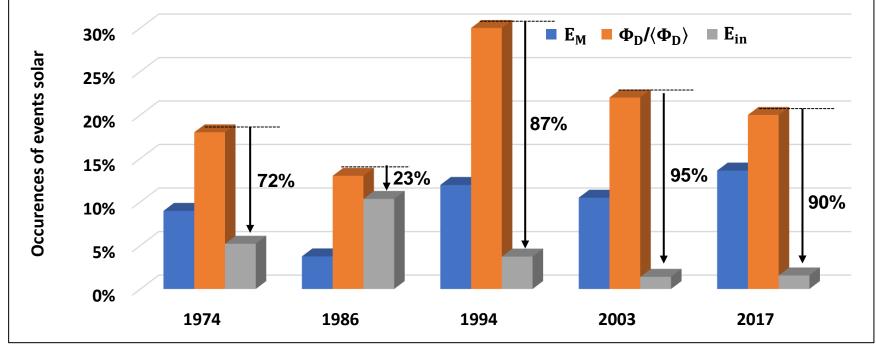


Figure 7: Occurrence  $E_M$  field,  $E_{in}$  and  $\Phi_D/\langle \Phi_D \rangle$ 

Table 3: Correlation of some solar wind parameters	5.
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	Outer Minima [year]										
	1974	1986	1994	2003	2017						
$\Phi_{ m D}/\langle\Phi_{ m D} angle$	18%	13%	30%	22%	20%						
Ein & EM	60%	76%	-59%	64%	56%						







During the peaks of the outer minima of SC :

- At high speeds, the normalized daytime reconnection rate is likely to be improved when the IMF-Bz is antiparallel to the geomagnetic field.
- Significant geomagnetic activity is sometimes present even in the absence of such important ICMEs.
- HSSW represent the stability criterion for areas that are particularly close to the outer minimum of solar cycles.
- For high By intensities with a north-pointing CMI-Bz, the trapping and energization of HSSW particles in the Earth's magnetic cavity has an enhanced influence on the magnetospheric convection electric field.





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## **THANKS FOR YOUR ATTENTION !!**