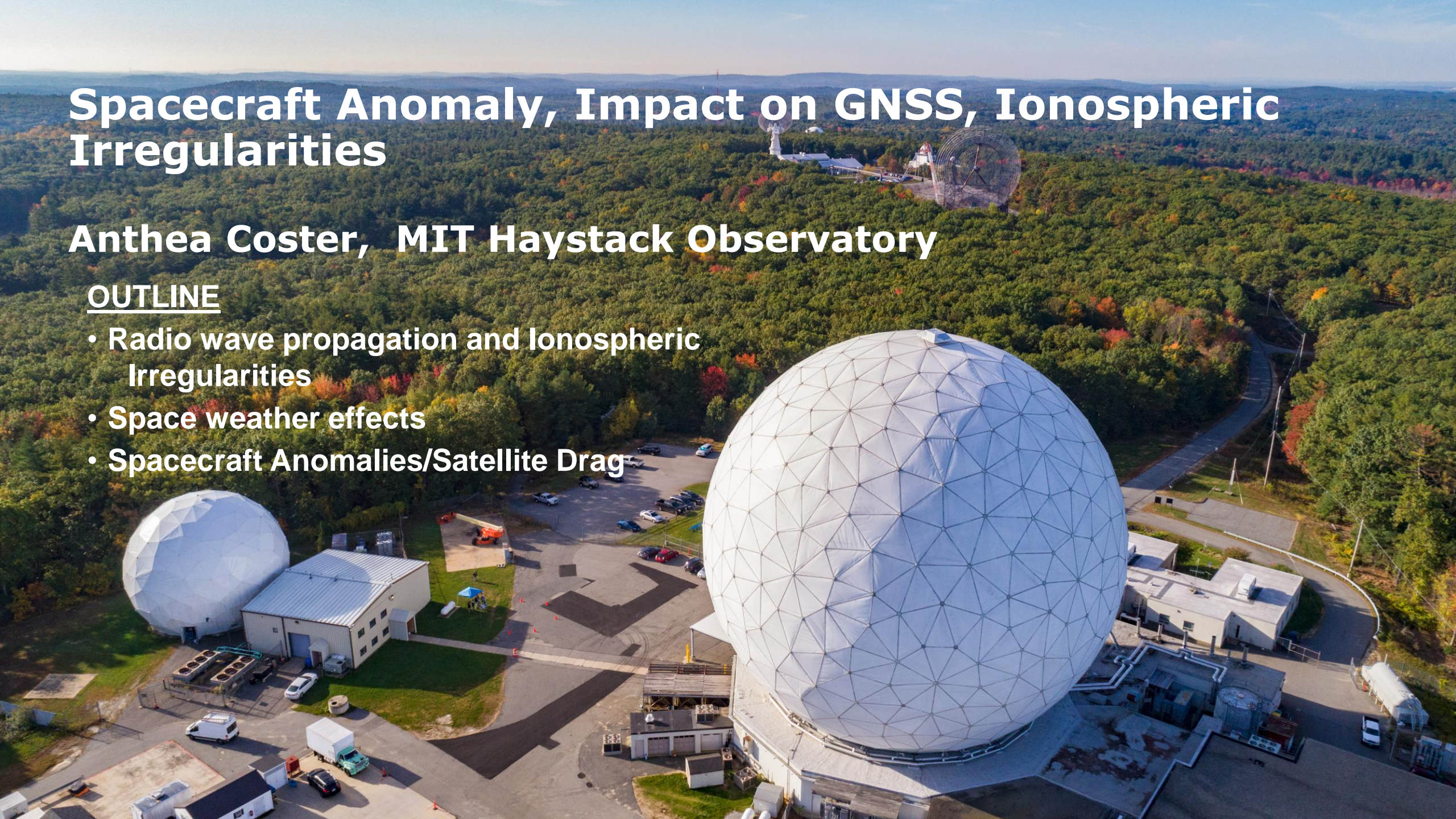


Spacecraft Anomaly, Impact on GNSS, Ionospheric Irregularities

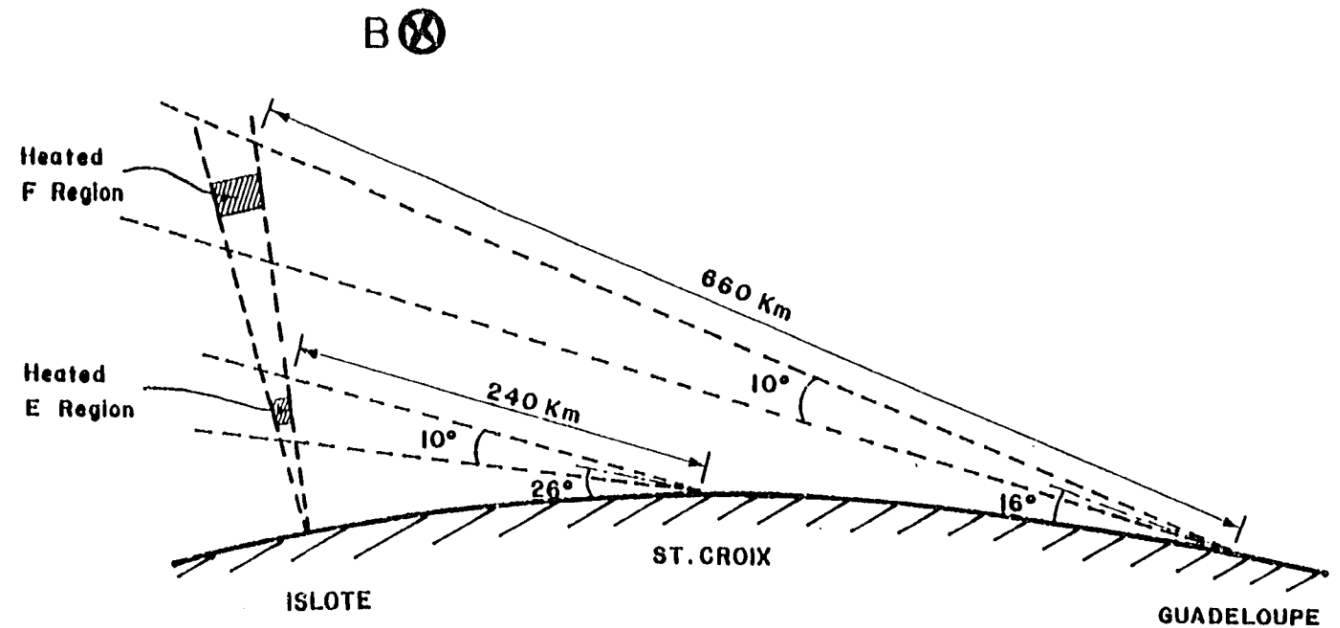
Anthea Coster, MIT Haystack Observatory

OUTLINE

- Radio wave propagation and Ionospheric Irregularities
- Space weather effects
- Spacecraft Anomalies/Satellite Drag

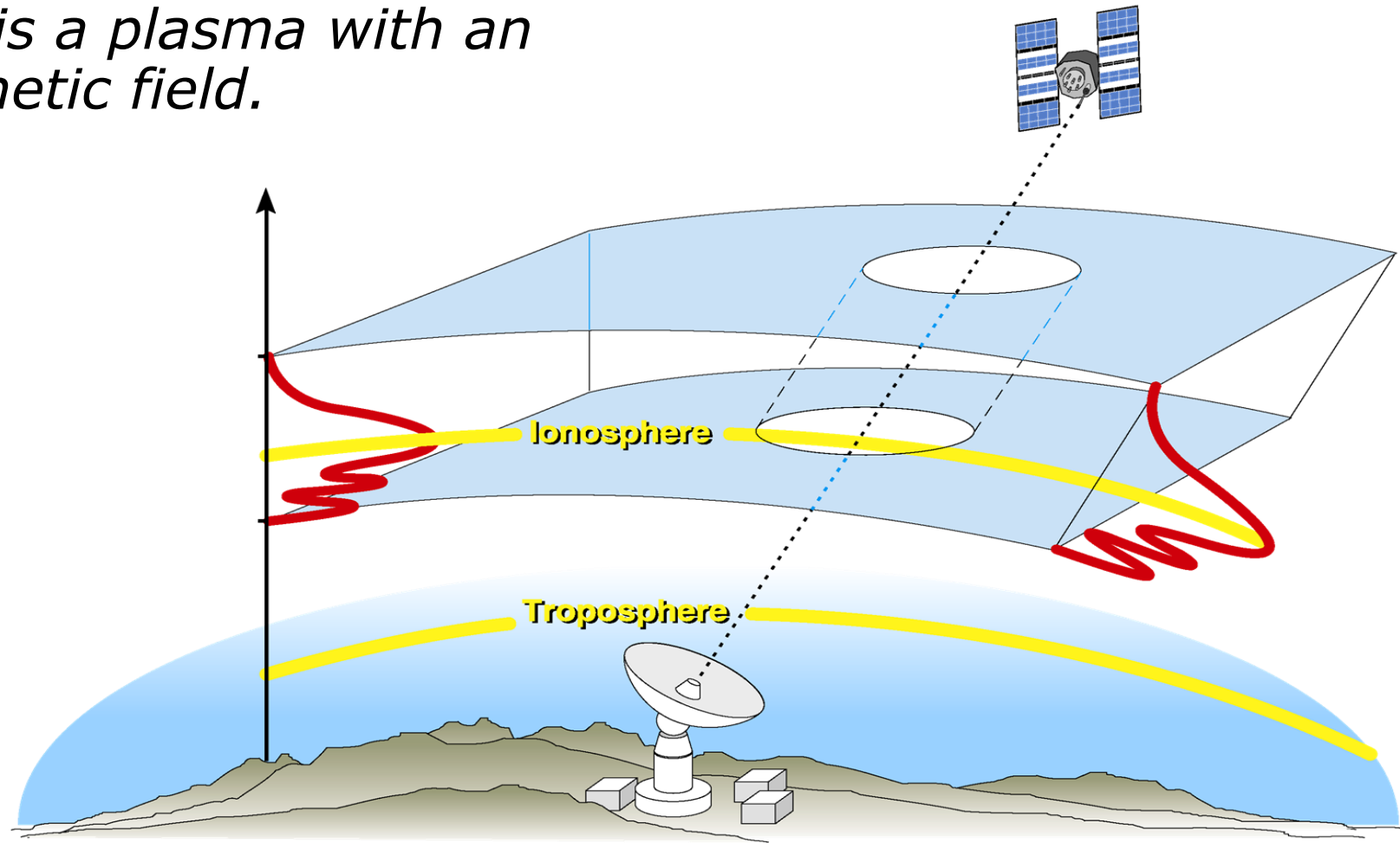


PhD Research in 1980 : Time Synchronization Hard



Radio Wave Propagation

The ionosphere is a plasma with an embedded magnetic field.



Appleton-Hartree Equation

$$n^2 = 1 - \frac{X}{1 - iZ \cdot \frac{\frac{1}{2}Y^2 \sin^2 \theta}{1 - X - iZ} \pm \frac{1}{1 - X - iZ} \left(\frac{1}{4}Y^4 \sin^4 \theta + Y^2 \cos^2 \theta (1 - X - iZ)^2 \right)^{1/2}}$$

or, alternatively^[4]:

$$n^2 = 1 - \frac{X(1 - X)}{1 - X - \frac{1}{2}Y^2 \sin^2 \theta \pm \left(\left(\frac{1}{2}Y^2 \sin^2 \theta \right)^2 + (1 - X)^2 Y^2 \cos^2 \theta \right)^{1/2}}$$

n = complex refractive index

$$i = \sqrt{-1}$$

$$X = \frac{\omega_0^2}{\omega^2}$$

$$Y = \frac{\omega_H}{\omega}$$

$$Z = \frac{\nu}{\omega}$$

ϵ_0 = permittivity of free space

μ_0 = permeability of free space

B_0 = ambient magnetic field strength

e = electron charge

m = electron mass

θ = angle between the ambient magnetic field vector and the wave vector

ν = electron collision frequency

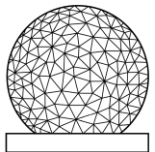
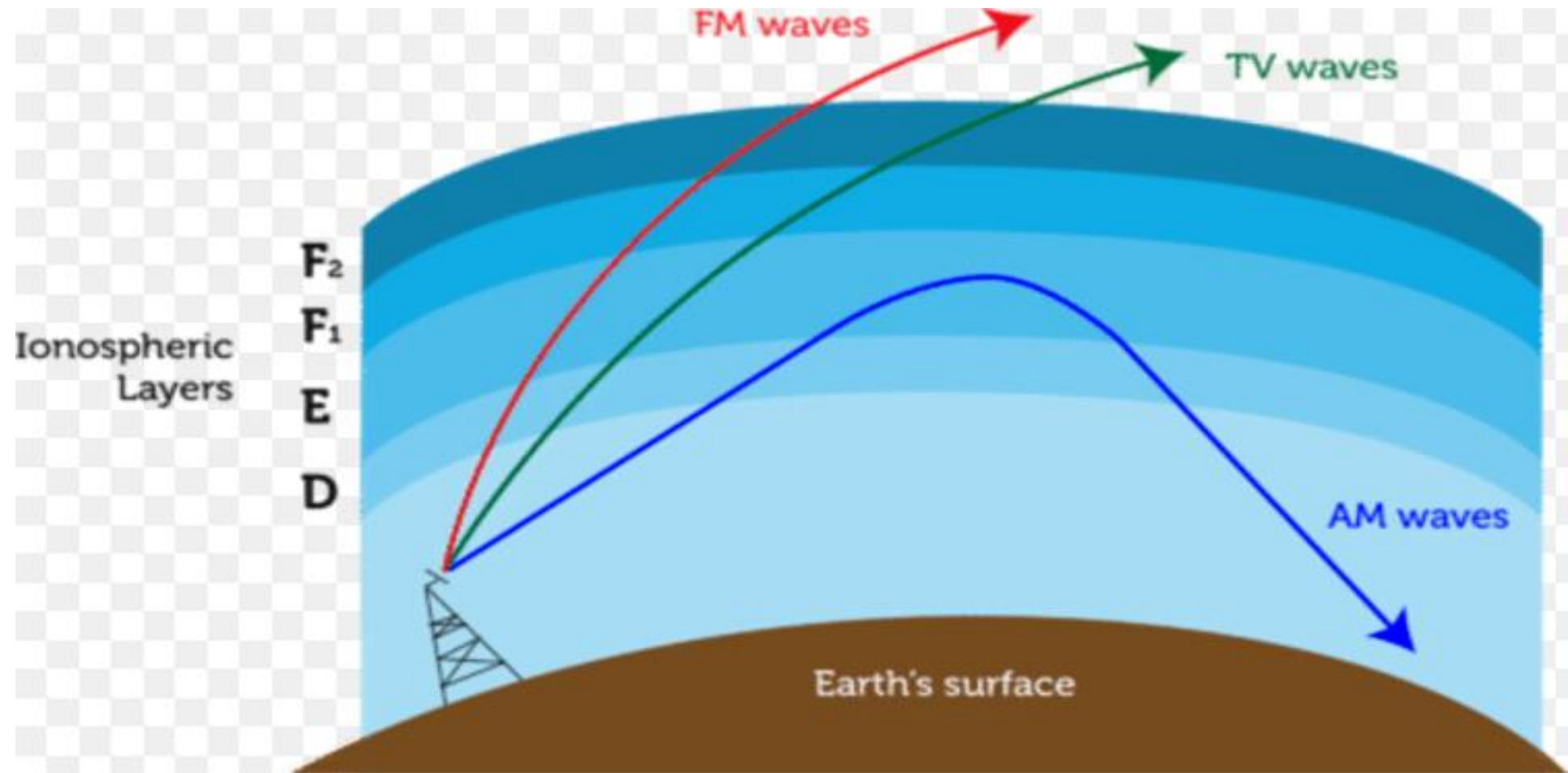
$\omega = 2\pi f$ (radial frequency)

f = wave frequency (cycles per second, or Hertz)

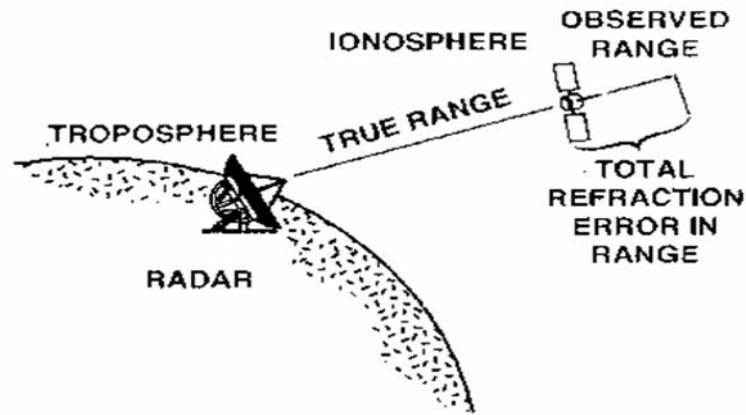
$$\omega_0 = 2\pi f_0 = \sqrt{\frac{Ne^2}{\epsilon_0 m}} = \text{electron plasma frequency}$$

$$\omega_H = 2\pi f_H = \frac{B_0 |e|}{m} = \text{electron gyro frequency}$$

Some Radio Waves are Reflected



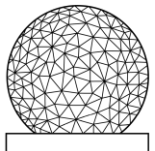
Ionospheric Range Delay



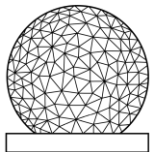
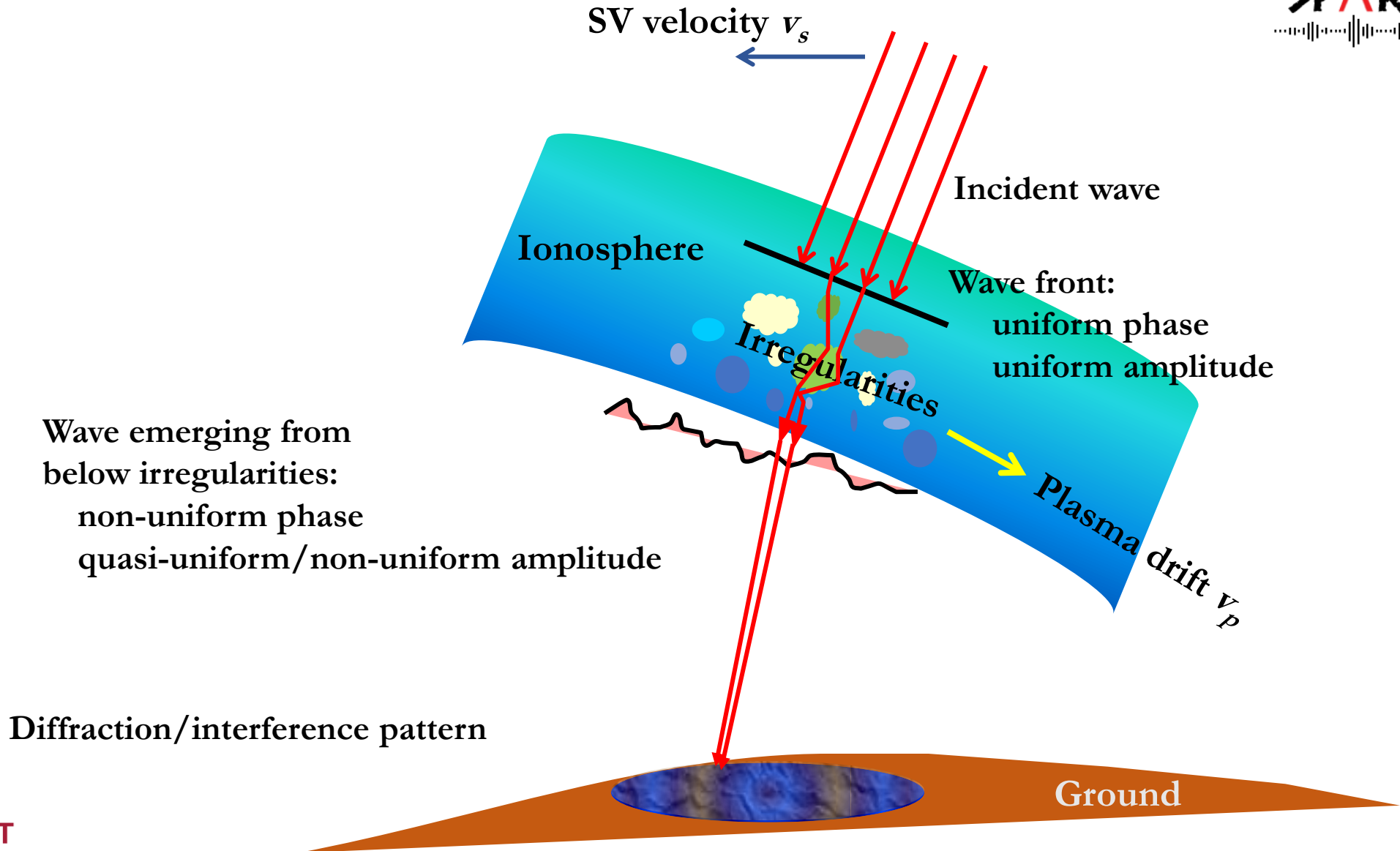
$$\Delta R_{ion}(meters) = \frac{40.3}{f^2} \underbrace{\int_0^R N_e dr}_{\text{TEC}}$$

Range Delay

<u>TEC</u>	<u>S-Band</u>	<u>L-Band</u>	<u>UHF</u>	<u>VHF</u>	<u>Elev</u>	<u>Mapping Function</u>
50	2.4 m	12 m	104 m	787 m	90 °	x 1
110	5.1 m	26 m	223 m	1.7 km	20 °	x 2.12

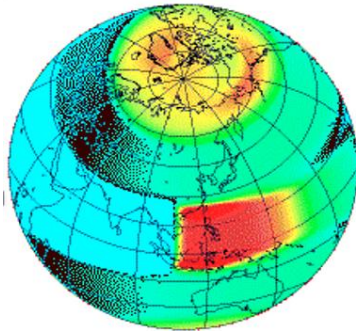


Space Weather - Scintillation



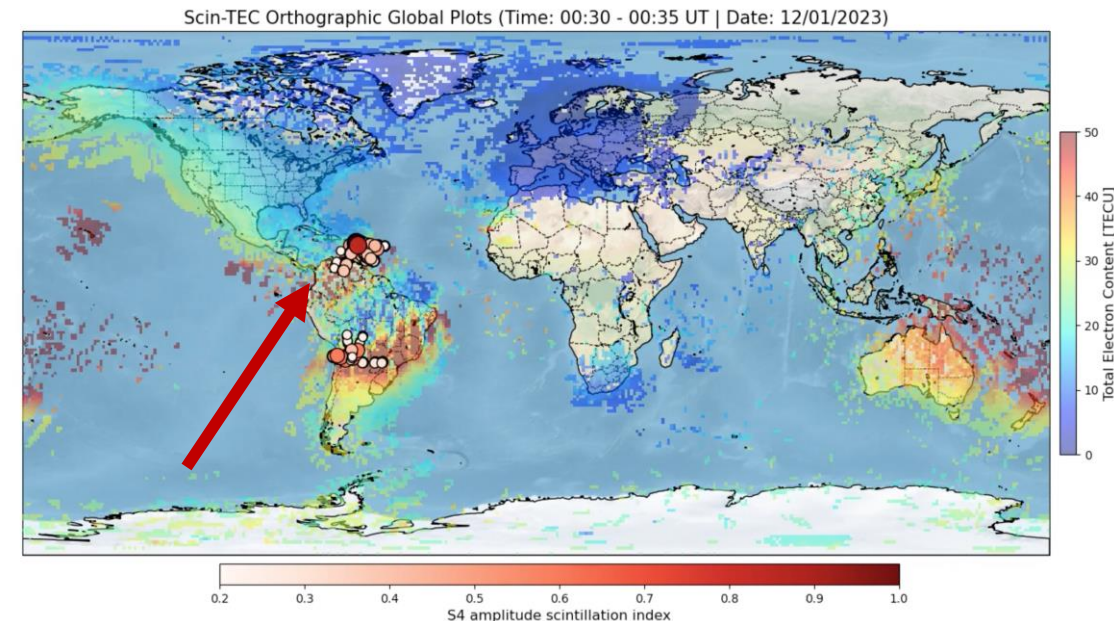
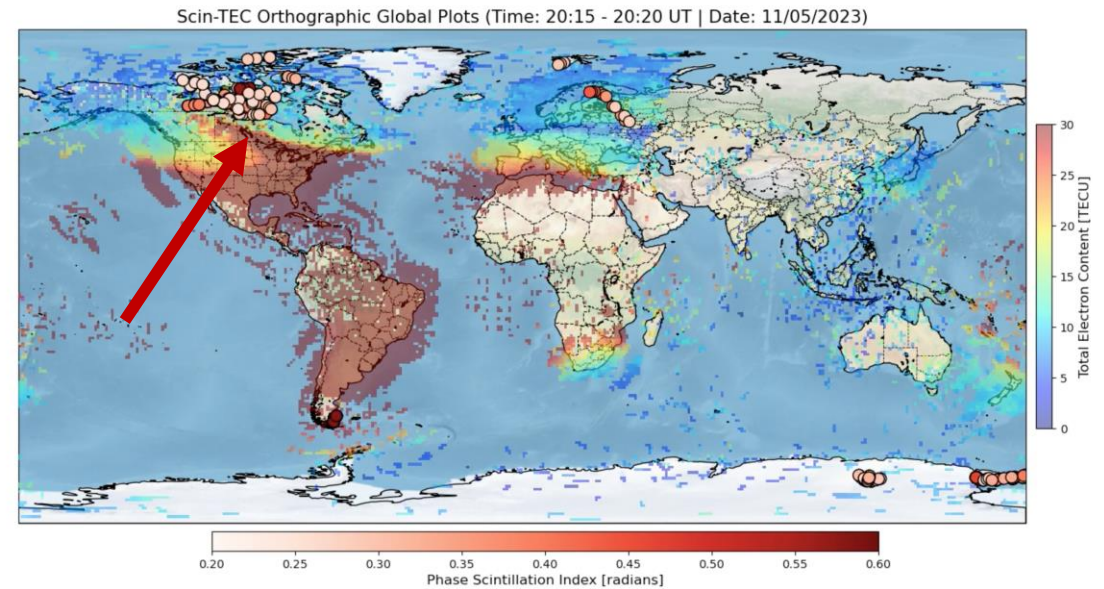
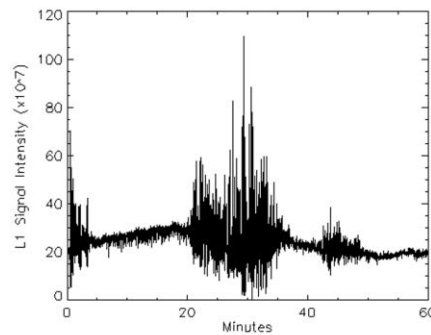
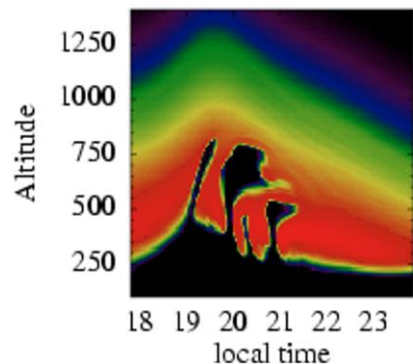
High Latitude GNSS Scintillation

Associated with aurora and precipitation,
Storm-enhanced density features (SED),
polar cap patches, Tongues of Ionization (TOI)



Low-Latitude GNSS Scintillation

Associated with Equatorial Spread-F, “bubbles”



Scintillation Impacts

- GNSS loss of lock due to phase or amplitude scintillation
- Radars lose ability to coherently track objects in space
- Loss of or difficulty in receiving communication channels

OUTLINE

- Radio wave propagation and ionospheric irregularities



- Space weather effects
- Spacecraft Anomalies/Satellite Drag

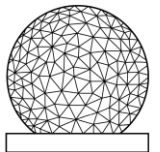
Three Agents of Space Weather

J.J.J. Kunches, NOAA

Electromagnetic Emission
8 minutes to Earth

Charged Particle Radiation
Tens of minutes to several hours to Earth

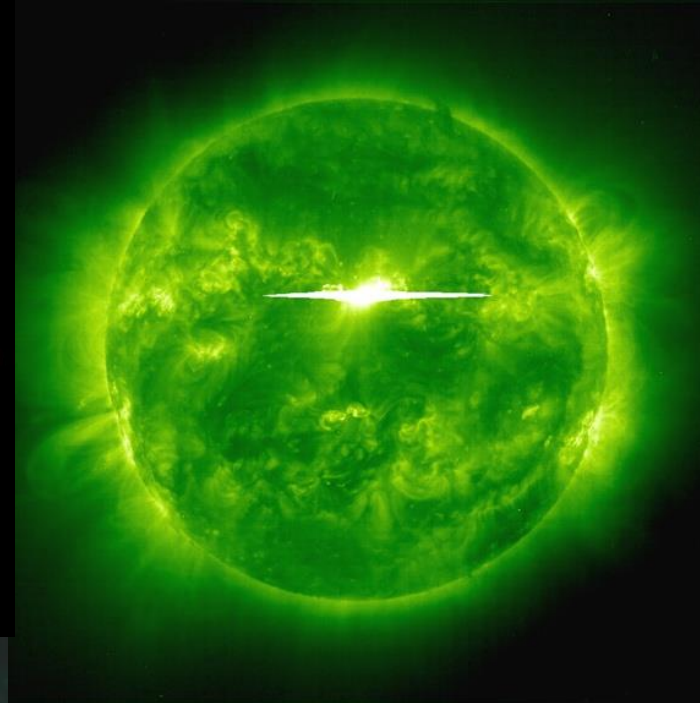
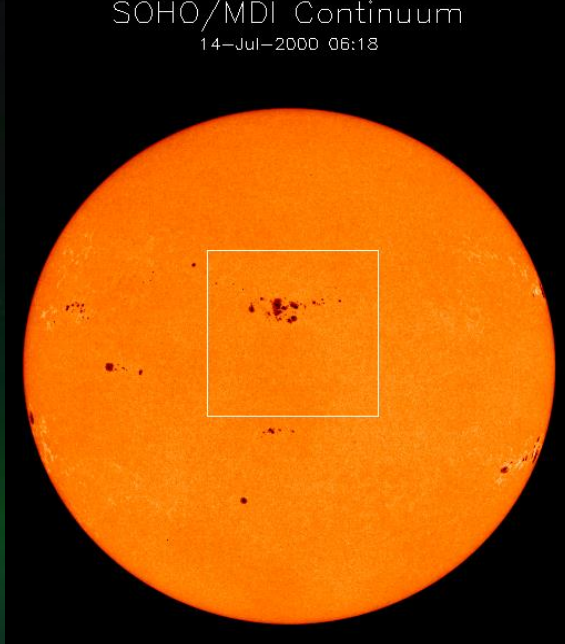
Magnetized Plasma
18-96 hours to Earth



MIT
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OBSERVATORY

Solar Flares/Solar Radio Bursts (Radio Blackouts – R Scale)

SOHO/MDI Continuum
14-Jul-2000 06:18



J.J. Kunches, NOAA

Solar Flare –
A violent explosion in
the Sun's atmosphere;
energy equivalent of a
hundred million
hydrogen bombs.

- Arrival: 8 min from Sun to Earth
- Duration: minutes to 3 hrs
- Daylight-side impact

Impacts



GPS Network

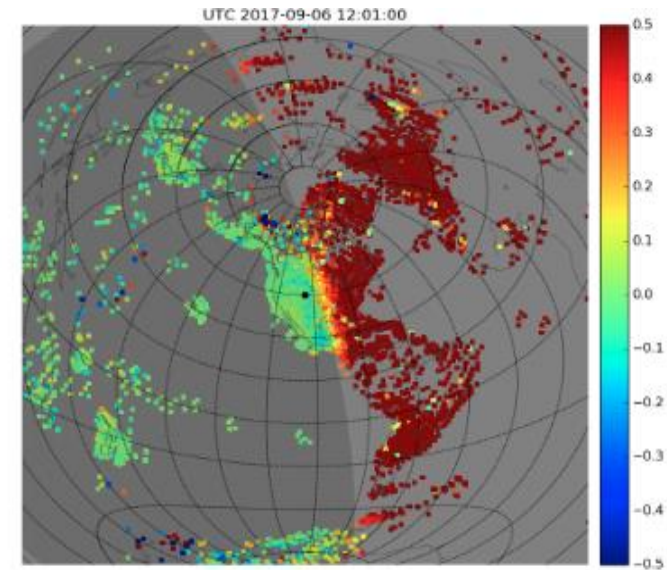
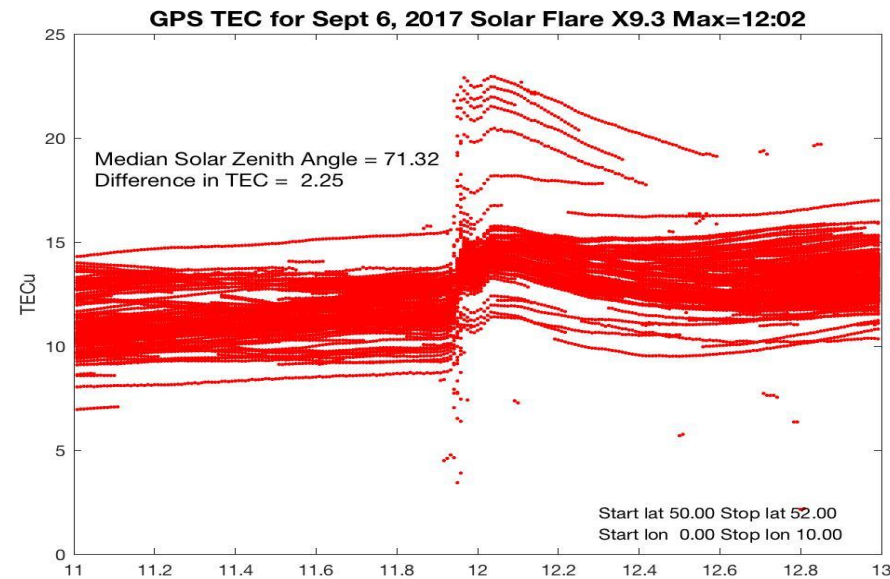
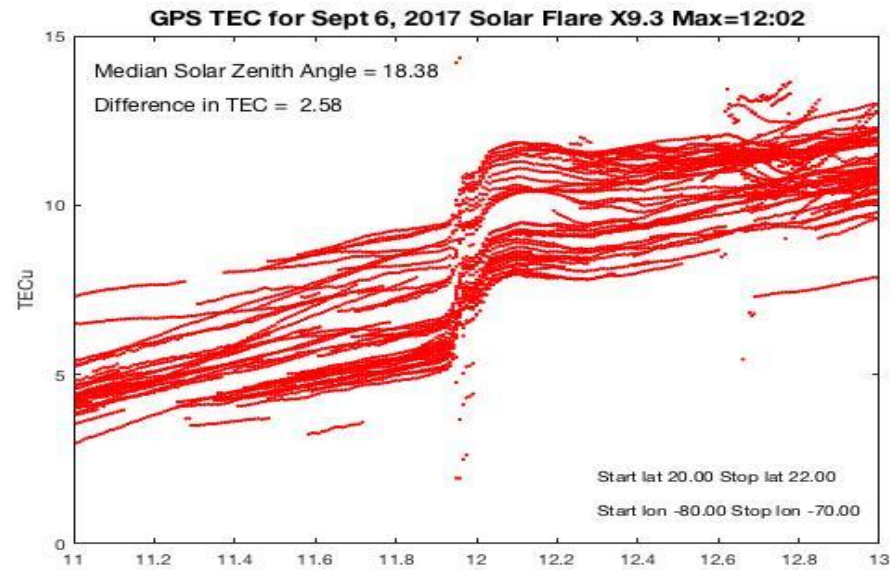
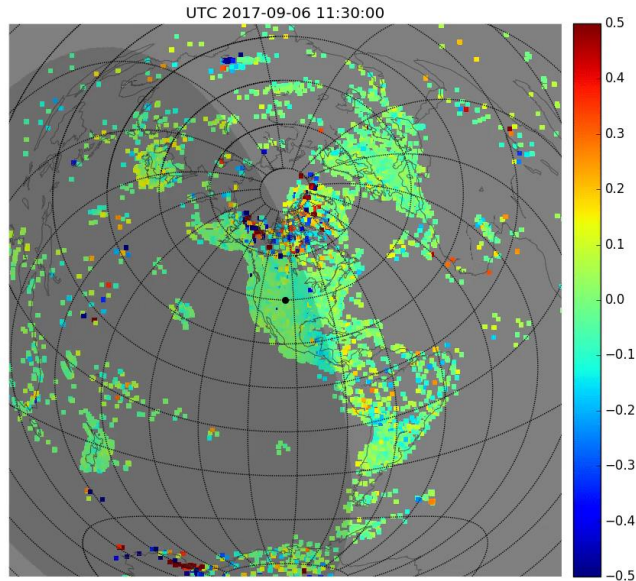


Communications
Ground and Space-based

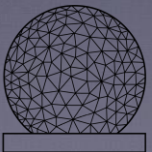
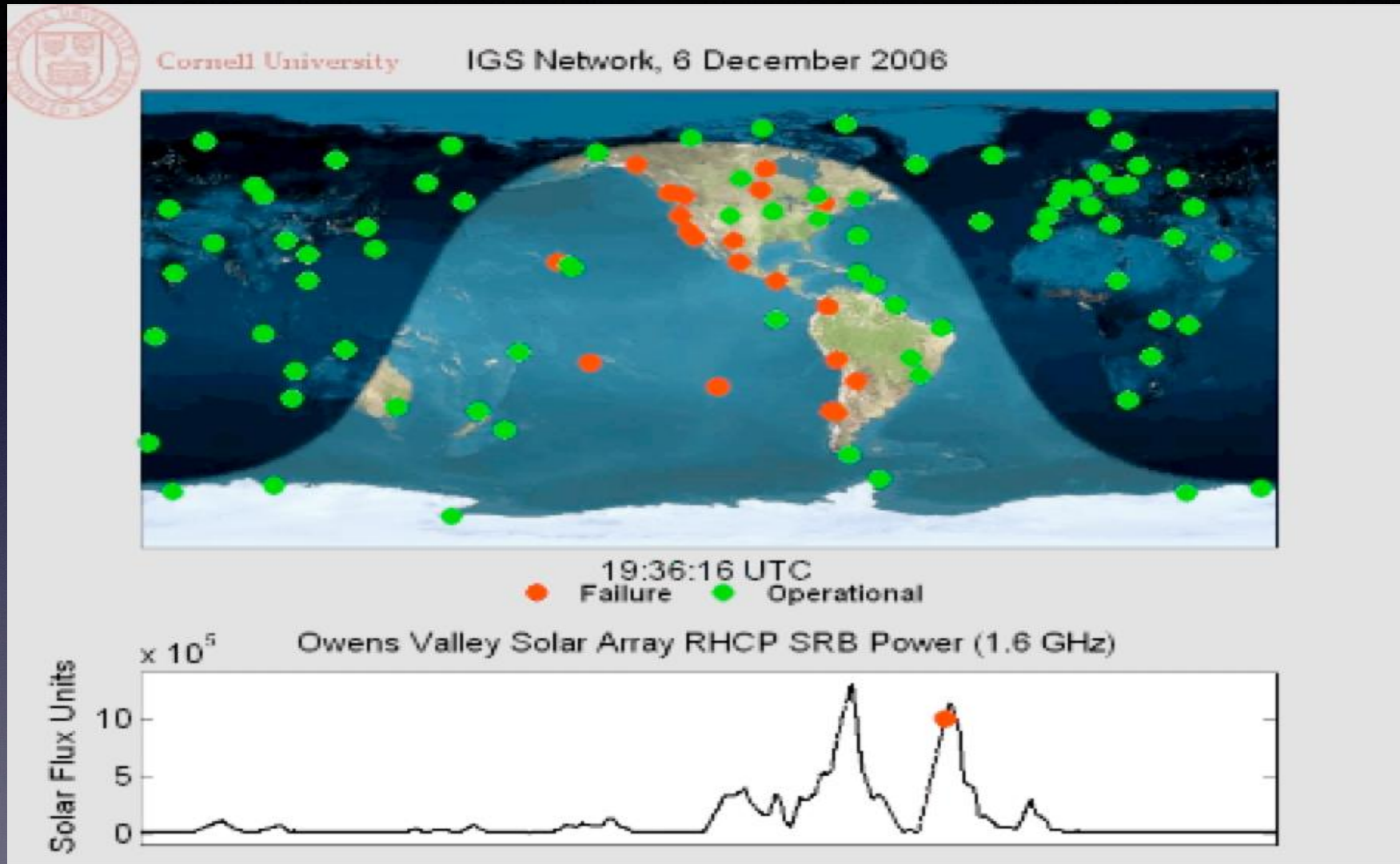


Radar

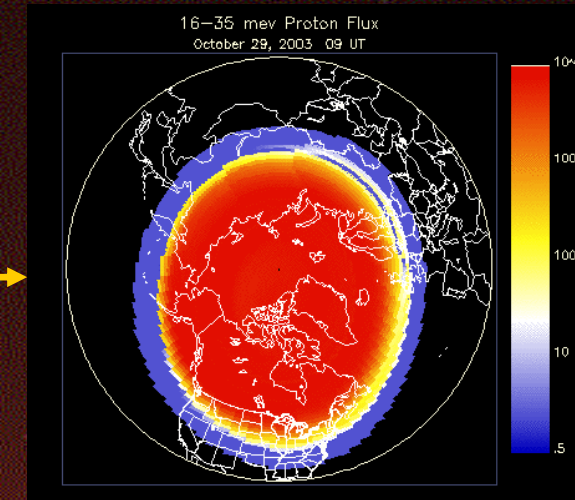
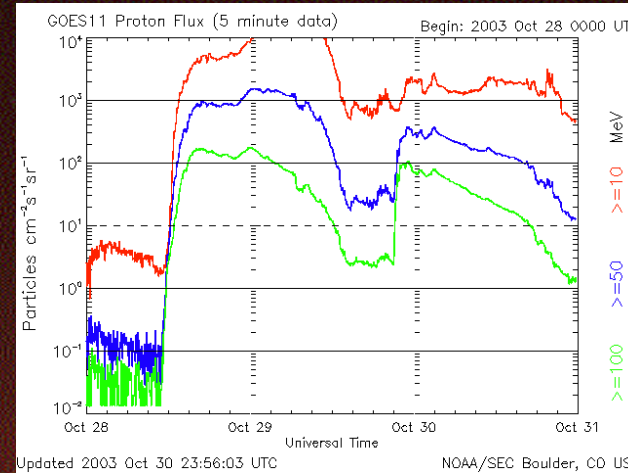
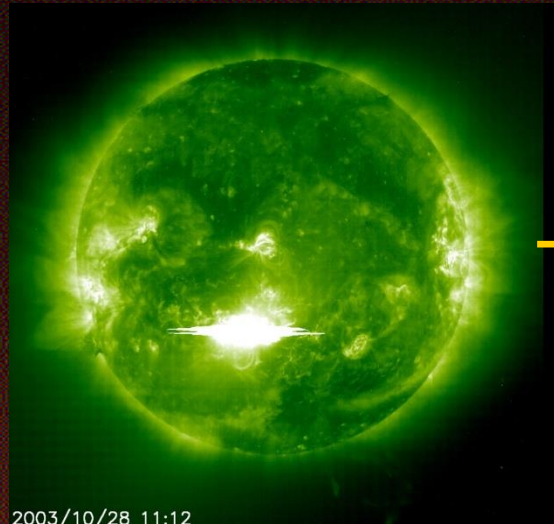
Sept 6, 2017 X9.3 Solar Flare



Increase in background noise level due to Solar Radio Bursts



Solar Radiation Storms (S Scale)



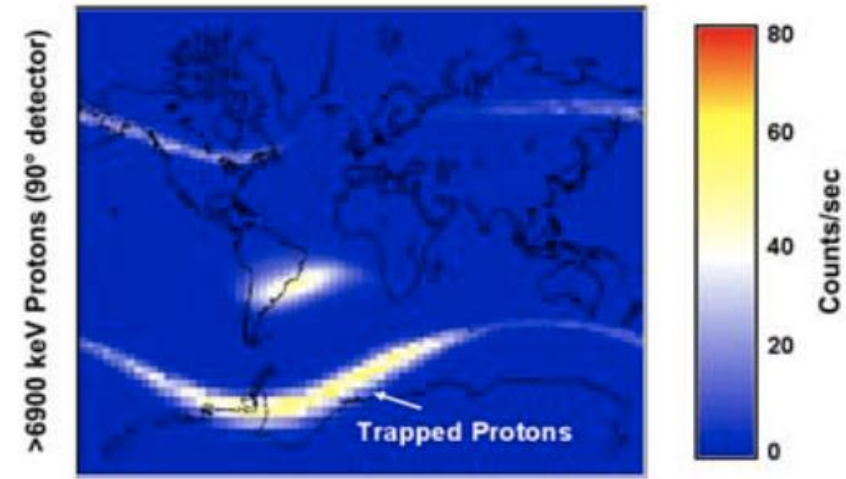
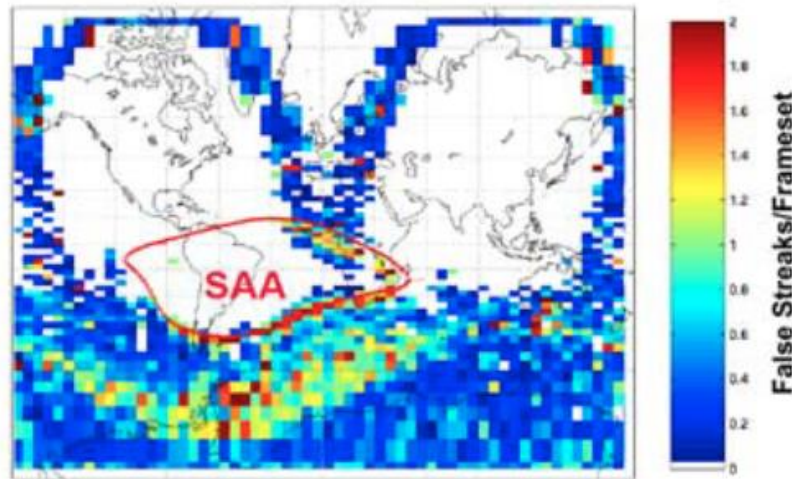
- Arrival: 30 minutes to several hours
- Duration: hours to days

Impacts...

- Satellite Operations (range from loss of data to loss of satellite)
- Aviation (communications and exposure concerns)
- High latitude HF comm outage
- Manned Spaceflight

Optical Tracking from Space

- Subtract background stars and look for streaks; calculate angular position



Space environment effects on GEO satellites

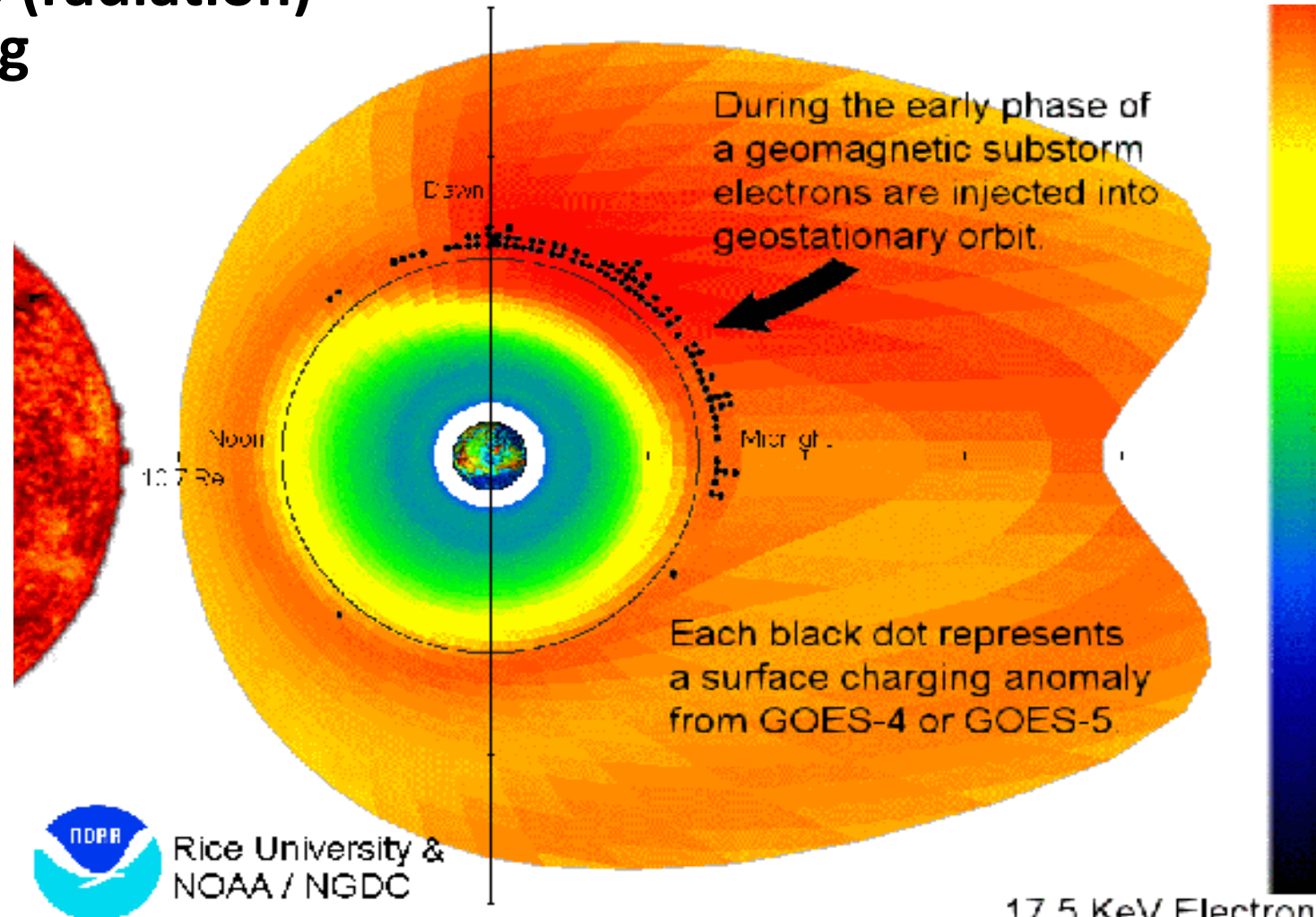
Electron and ion particle effects (radiation)

Internal and Surface Charging

Magnetic field variability

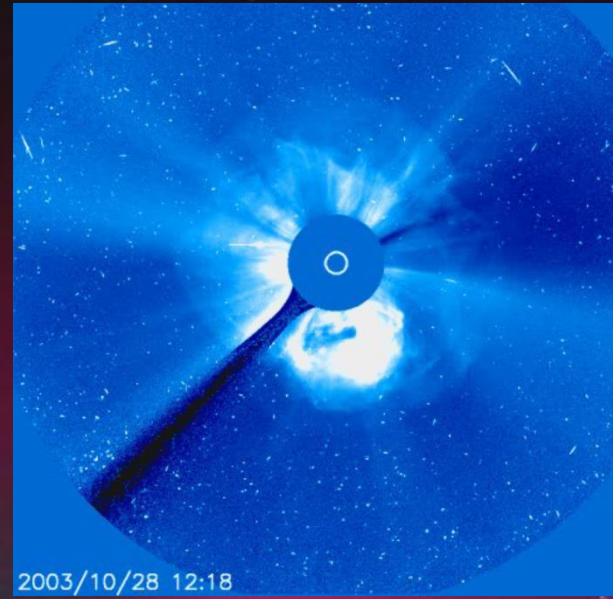
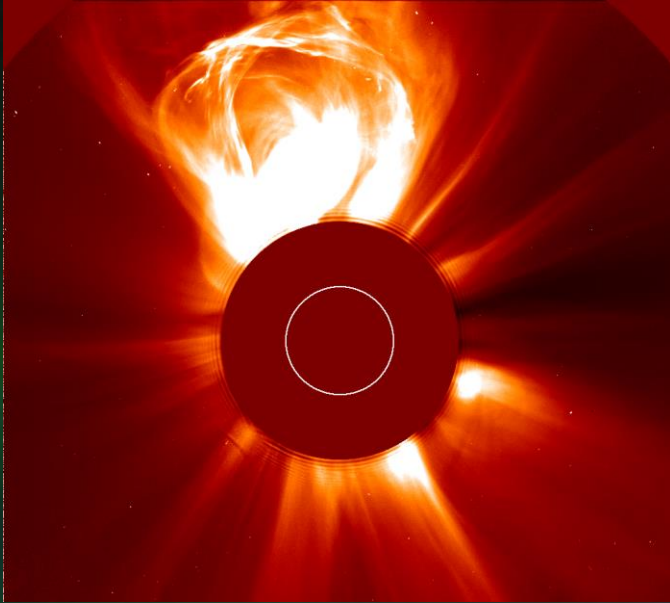
Affects attitude control

*Certain places known to be regions
of enhanced satellite anomaly
occurrence*



Geomagnetic Storms (G Scale)

Coronal Mass Ejections (CMEs)
create geomagnetic storms



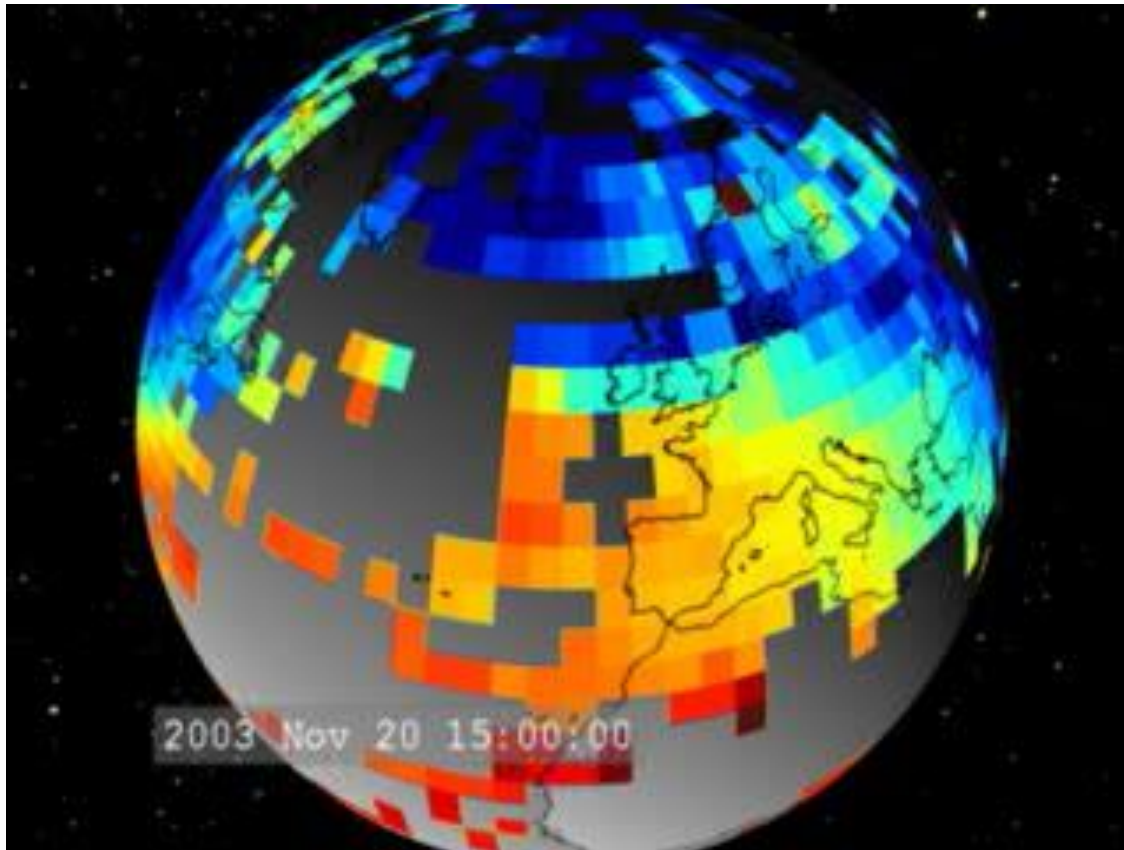
- Arrival: ~20 – 90 hours
- Duration: hours to a day
- Creates Ionospheric storms

Impacts...

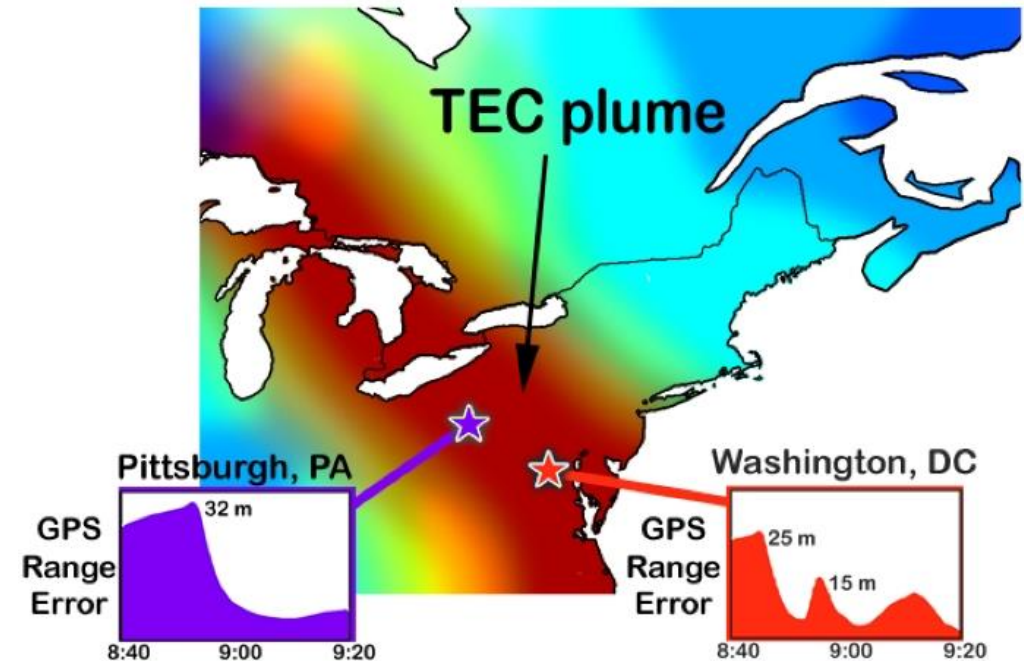
- Satellite Operations
- Aircraft operations
- Power grid operations
- GNSS operations
- Pipelines



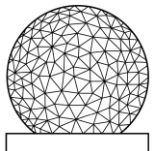
Large Gradients in the TEC lead to GNSS range errors



November 20, 2003



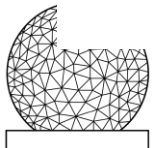
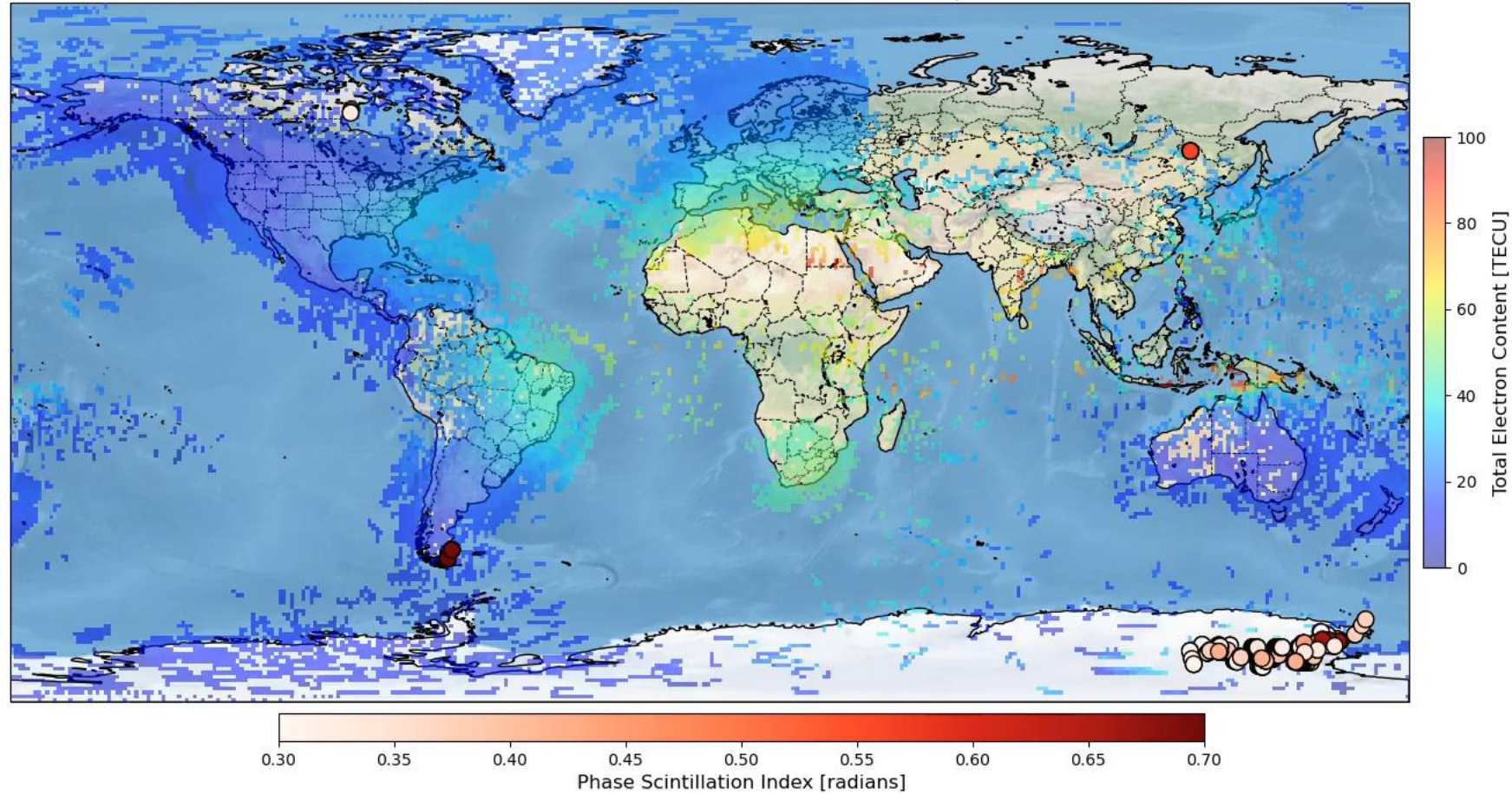
Courtesy of A. Mannucci



MIT
HAYSTACK
OBSERVATORY

SigmaPhi/TEC May 10, 2024

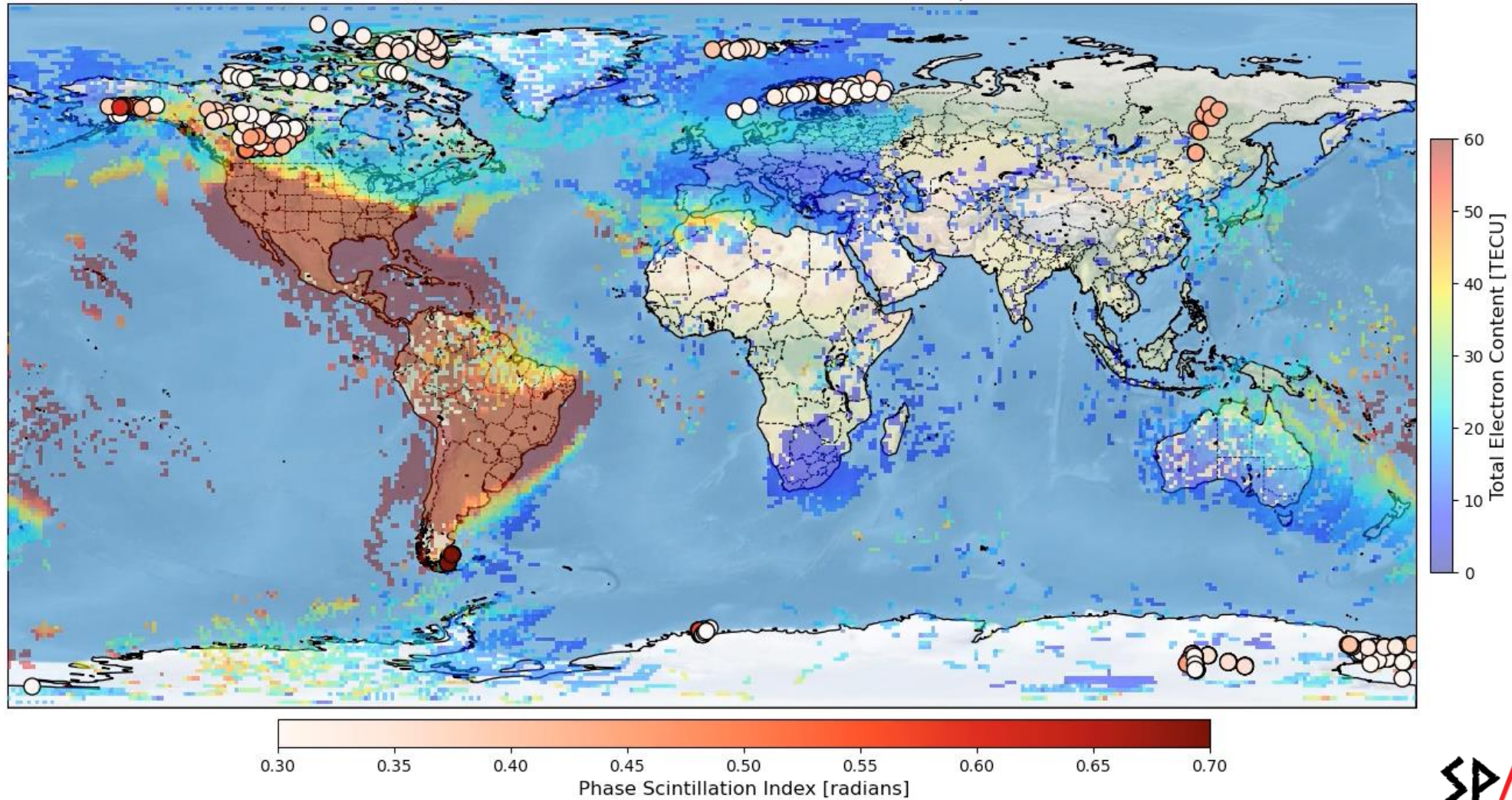
Scin-TEC Orthographic Global Plots (Time: 12:00 - 12:05 UT | Date: 05/10/2024)



**HAYSTACK
OBSERVATORY**

SigmaPhi/TEC May 10, 2024

Scin-TEC Orthographic Global Plots (Time: 21:20 - 21:25 UT | Date: 05/10/2024)



Outline

Overview – space weather

Radio wave propagation (GNSS)

Ionospheric Irregularities

 **Spacecraft Anomaly/ Satellite Drag**

Atmospheric drag

- Satellites in Low Earth Orbit (LEO) (<2,000 km altitude) are significantly impacted by atmospheric drag
- Atmospheric drag is largest source of error in orbit prediction:

$$F_i = \frac{1}{2} \rho A C_i v^2$$

F_i = Force, N

i = d(drag), l(lift), and s (side slip)

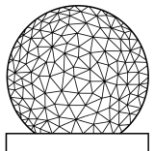
A = area, m^2

C_i = coefficient

v = spacecraft speed with respect to the atmosphere, m/s

ρ = atmospheric mass density, kg/m^3

The largest error is generally due to inaccurate estimates of the thermospheric density ρ



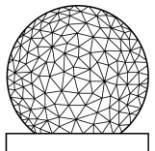
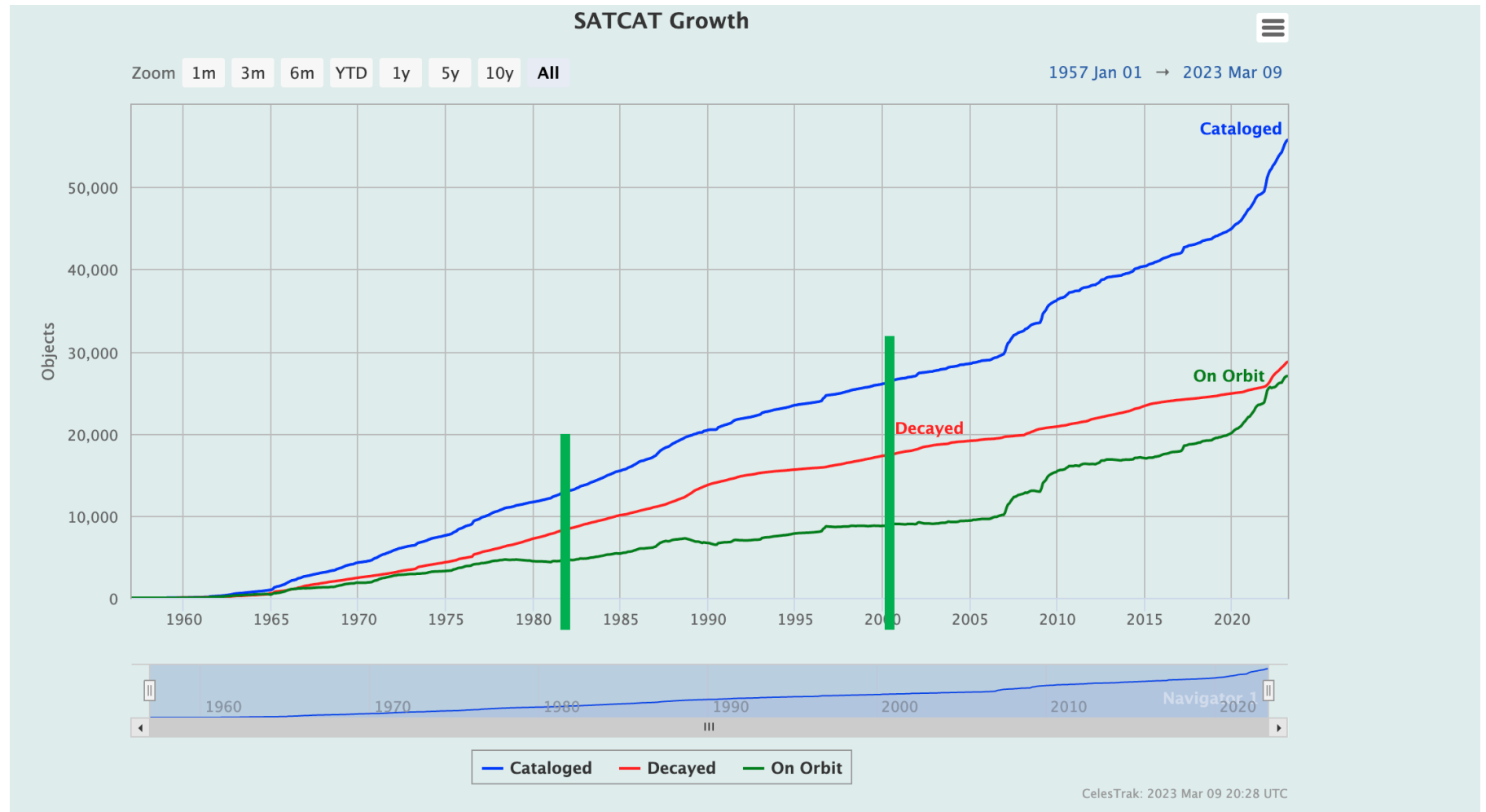
MIT
HAYSTACK
OBSERVATORY

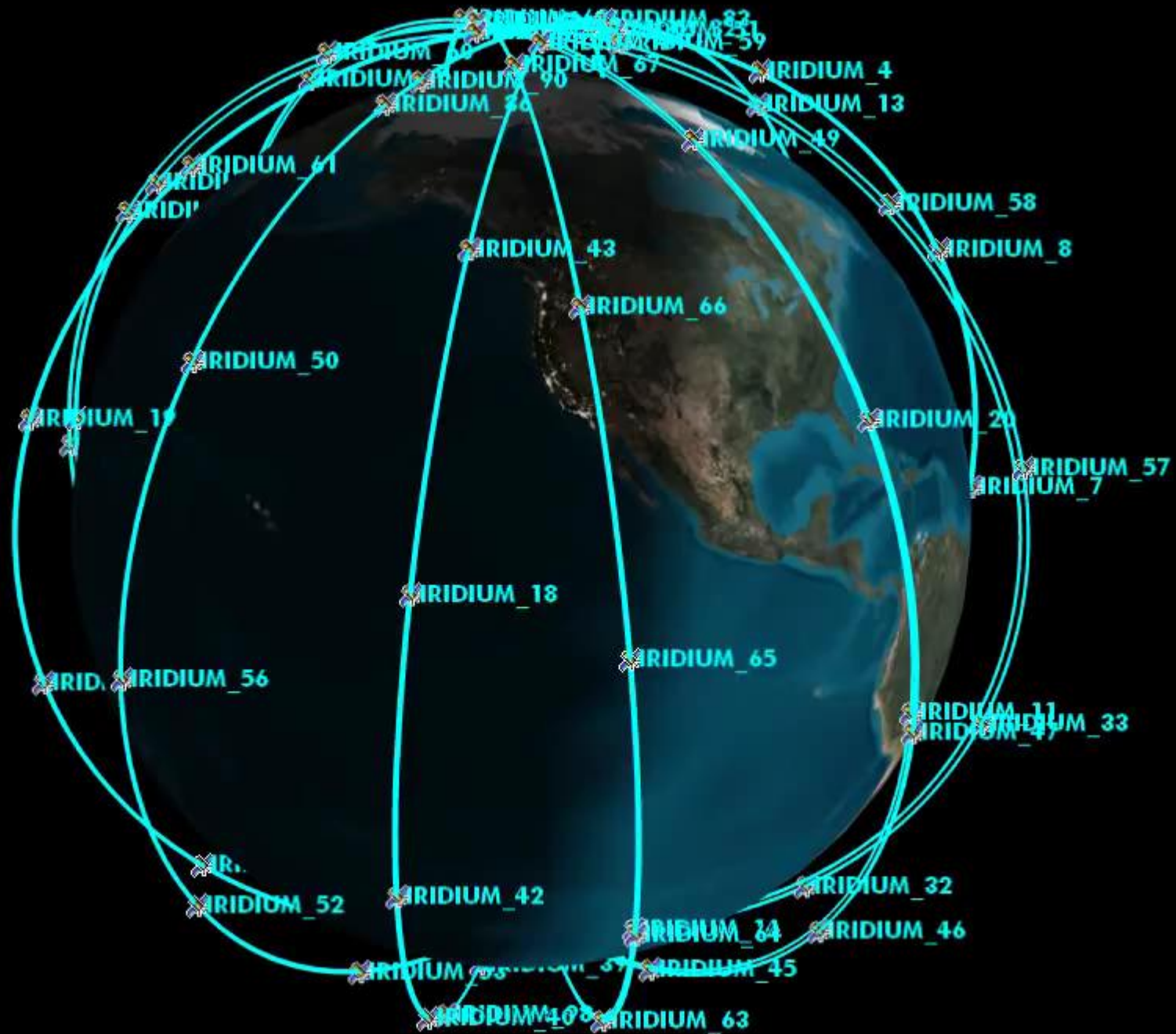
Courtesy of Julia Briden, MIT

Video Courtesy: NASA



Historical SATCAT Growth, 1957 to Present





Summary

- Space Weather Impacts:
 - Radiation hazards satellites (not necessarily advertised)
 - HF comm outages (noticed May 10-12, 2024)
 - Satellite tracking issues (not necessarily advertised)
 - GNSS degradation – some noted
- Correct estimation of Atmospheric Drag measurements more significant as number of satellites in LEO grows

