

Space Weather prediction

USING VARIOUS TECHNIQUES, INCLUDING MACHINE LEARNING

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Outline



Forecasting has always been a powerful tool



High efficiency vs Loss of explicability



ML is cheap + ML evolves extremely fast



Quick, nice results vs data limits & Deep Learning



We need to do better + we need to understand



The next steps



Forecasting has always been a powerful tool

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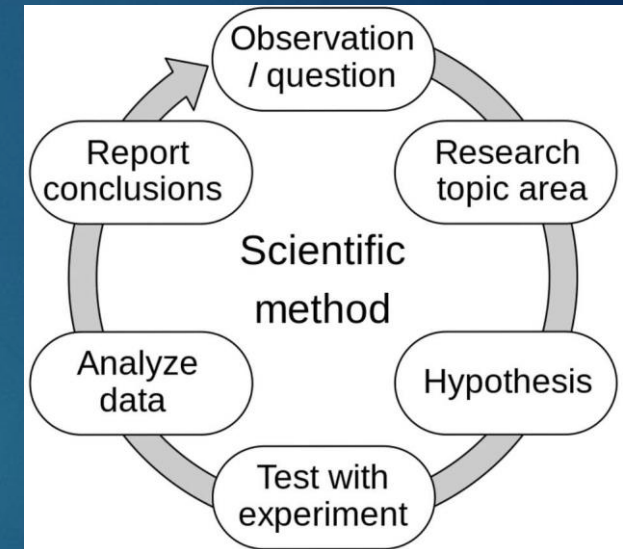
We have honed this ability first in the eons of the biological evolution, then encoding it in our culture, 'recently' by applying the scientific method to our everyday life problems.

The mathematical tools to understand why some events are more difficult to forecast than others.

The development in the last decades in sciences of complex systems and statistics have allowed us to better define the limits of predictability and - often- to extend those limits.

BUT! The robust forecast of flare eruptions still escapes us.

And also the apparently simpler problem of the propagation of a coronal mass ejection in the interplanetary medium has not been solved to the limit that we would like, while we fight with the uncertainties associated with the boundary conditions.

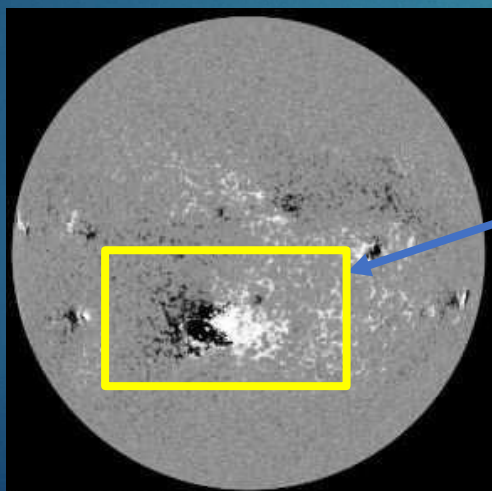


Scientific approach vs Machine Learning

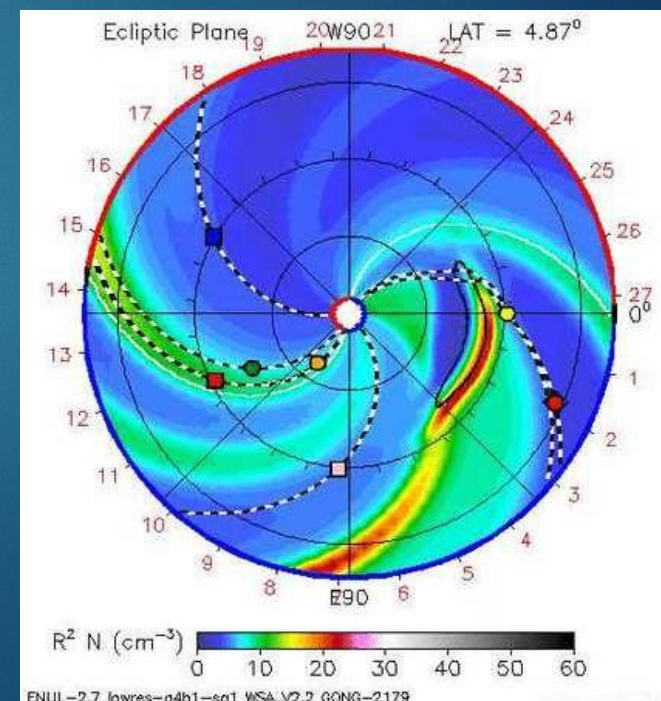
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- ▶ It is now feasible to compute the trajectory of plasma and magnetic field structures under the MHD equations in domains as large as the Heliosphere.
- ▶ Or try to predict -just by analyzing full disk images or the magnetograms- whether a given solar Active Region will release part of its stored energy as high energy photons and particles, or shoot out a coronal mass ejection.



Solar flares originate from magnetically active regions (ARs)

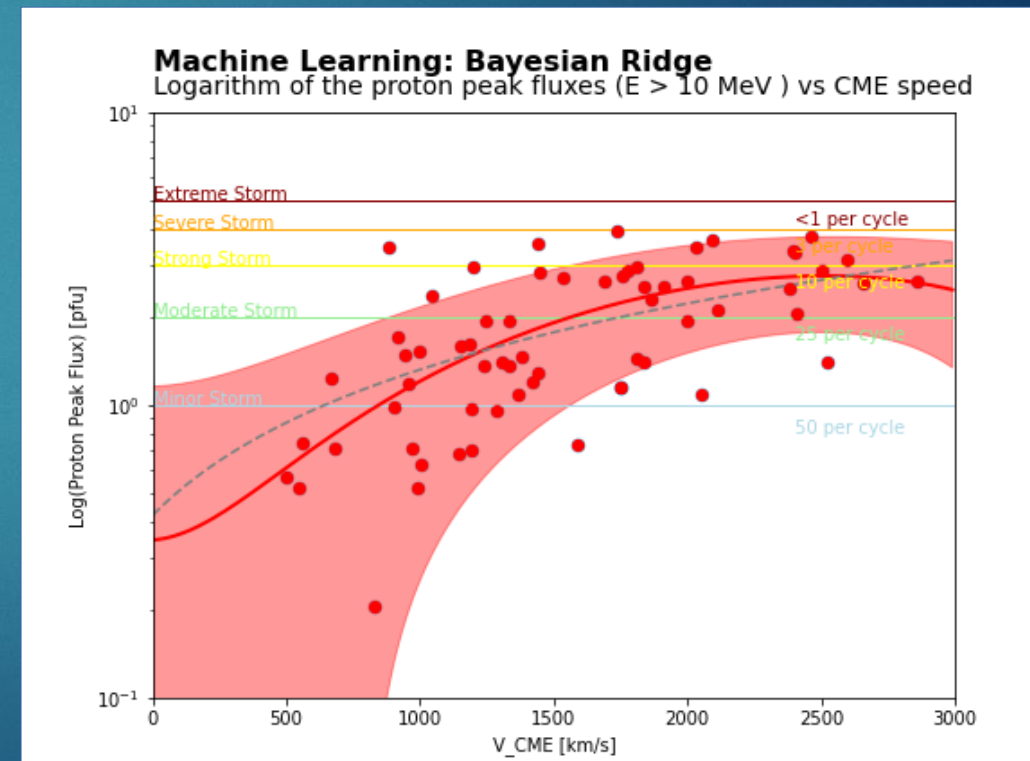
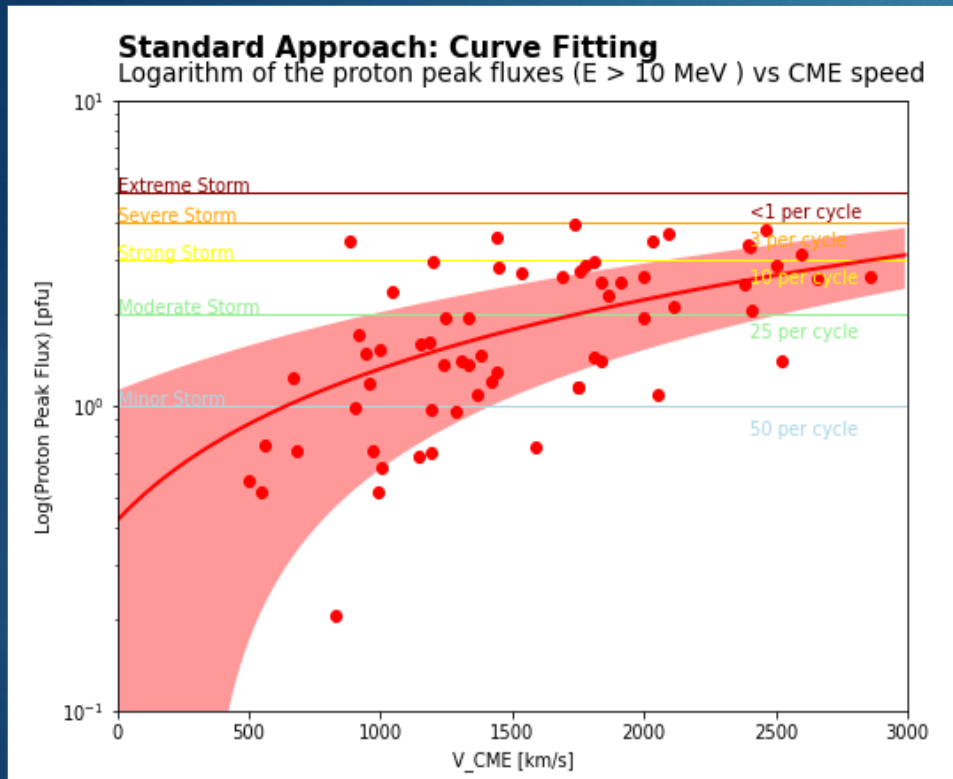


Scientific approach vs Machine Learning

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- Empirical, data-driven models and data fitting now enrolled into ML...



Papaioannou A, Sandberg I, Anastasiadis A, Kouloumvakos A, Georgoulis MK, Tziotziou K, Tsiropoula G, Jiggins P, Hilgers A. Solar flares, coronal mass ejections and solar energetic particle event characteristics. *Journal of Space Weather and Space Climate*. 2016;6:A42.

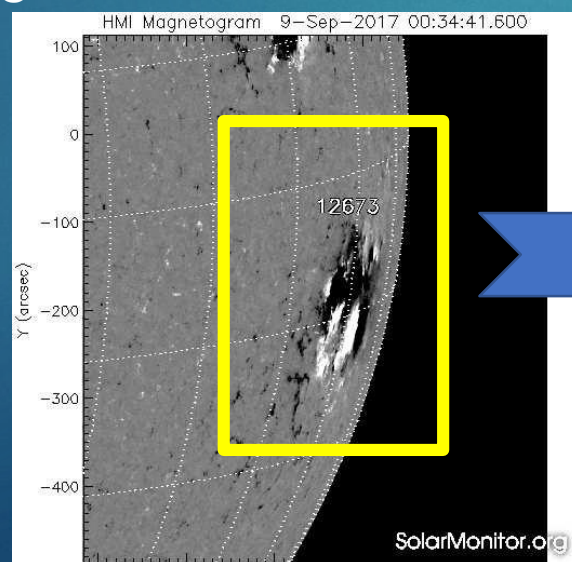
Stumpo, M., Benella, S., Laurenza, M., Alberti, T., Consolini, G. and Marcucci, M.F., 2021. Open issues in statistical forecasting of solar proton events: A machine learning perspective. *Space Weather*, 19(10), p.e2021SW002794.

Scientific approach vs Machine Learning

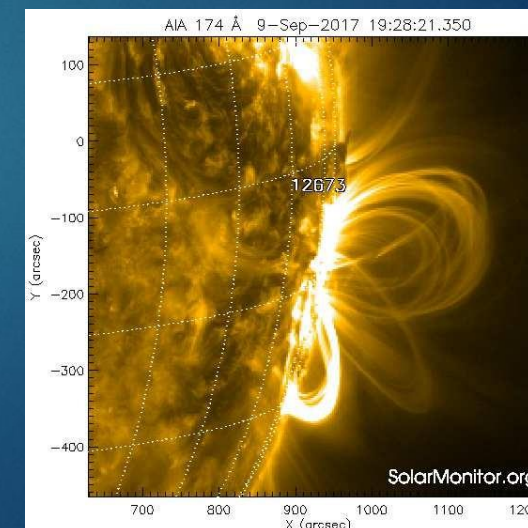
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- ▶ Exact space weather predictions are prevented either by intrinsic limitations or by the lack of knowledge about the present state.
- ▶ To counter these limitations, we have come up with clever and clever numerical techniques to solve the differential equations that typically describe our problems, and ensemble methods [3,4] to cope with measure errors and unknown variables. We strive to extract from remote measures all the relevant information [5], by applying our understanding of the physics of the problem, to feed this information into our forecasting algorithms.
- ▶ And we are getting better and better at this too.

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Will a solar flare of or above class **K** be generated in the next **Dt** hours?

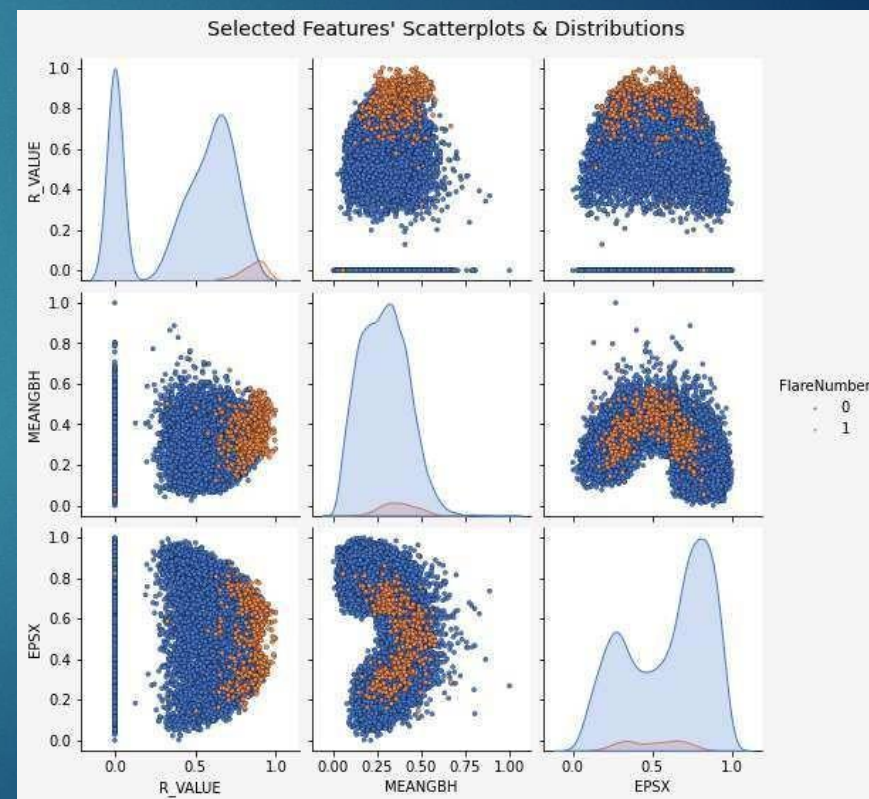


Scientific approach vs Machine Learning

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- ▶ In the absence of a definitive physical theory explaining the mechanisms of an AR, the best hope for forecasting solar eruptions lies in finding an empirical relationship between some well chosen features of ARs and the solar flares and CMEs.
- ▶ As a consequence, we have turned to the dark side and applied the methods of this hybrid of numerical methods, complex system science and statistics which is usually referred to as Machine Learning (ML).

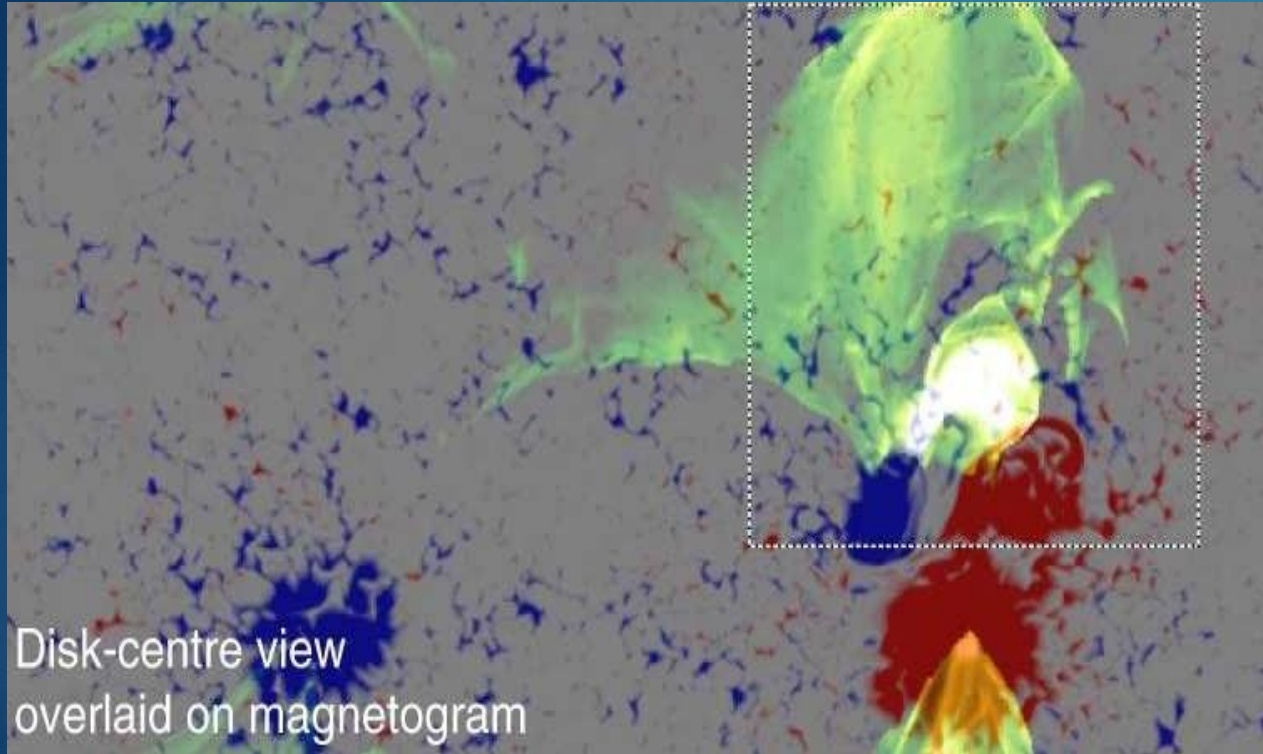
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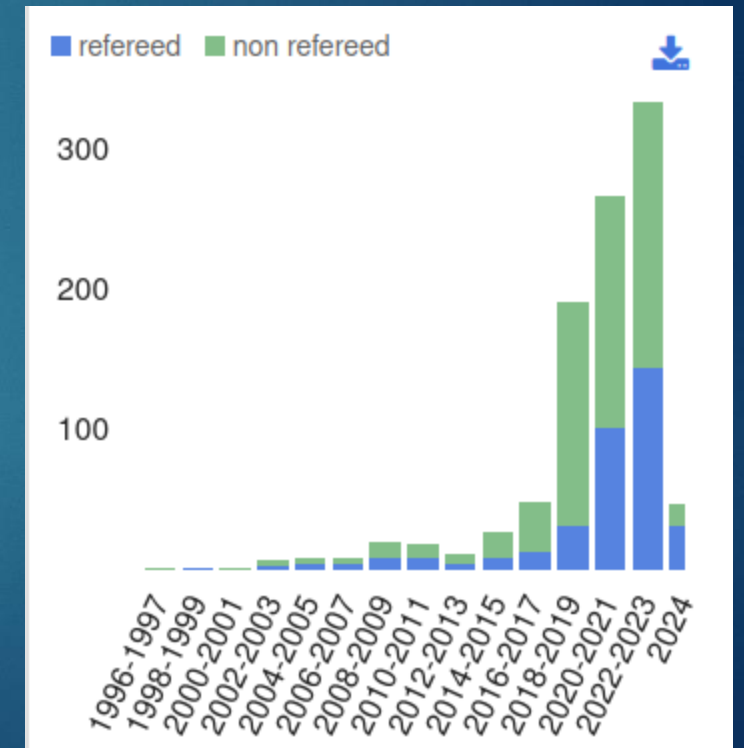
- Bobra, M. G., & Couvidat, S. (2015). Solar flare prediction using SDO/HMI vector magnetic field data with a machine-learning algorithm. *The Astrophysical Journal*, 798(2), 135.
- Florios, K., Kontogiannis, I., Park, S. H., Guerra, J. A., Benvenuto, F., Bloomfield, D. S., & Georgoulis, M. K. (2018). Forecasting solar flares using magnetogram-based predictors and machine learning. *Solar Physics*, 293(2), 28. <https://doi.org/10.1007/s11207-018-1250-4>

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Cheung, M. C. M., Rempel, M., Chintzoglou, G., Chen, F., Testa, P., Martínez-Sykora, J., ... & McIntosh, S. W. (2019). A comprehensive three-dimensional radiative magnetohydrodynamic simulation of a solar flare. *Nature Astronomy*, 3(2), 160-166. <https://doi.org/10.1038/s41550-018-0629-3>



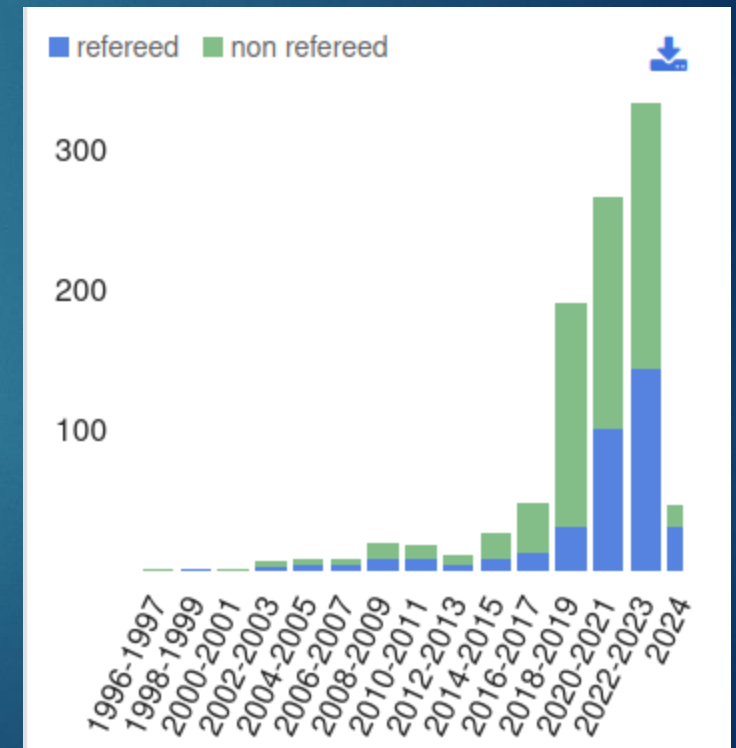
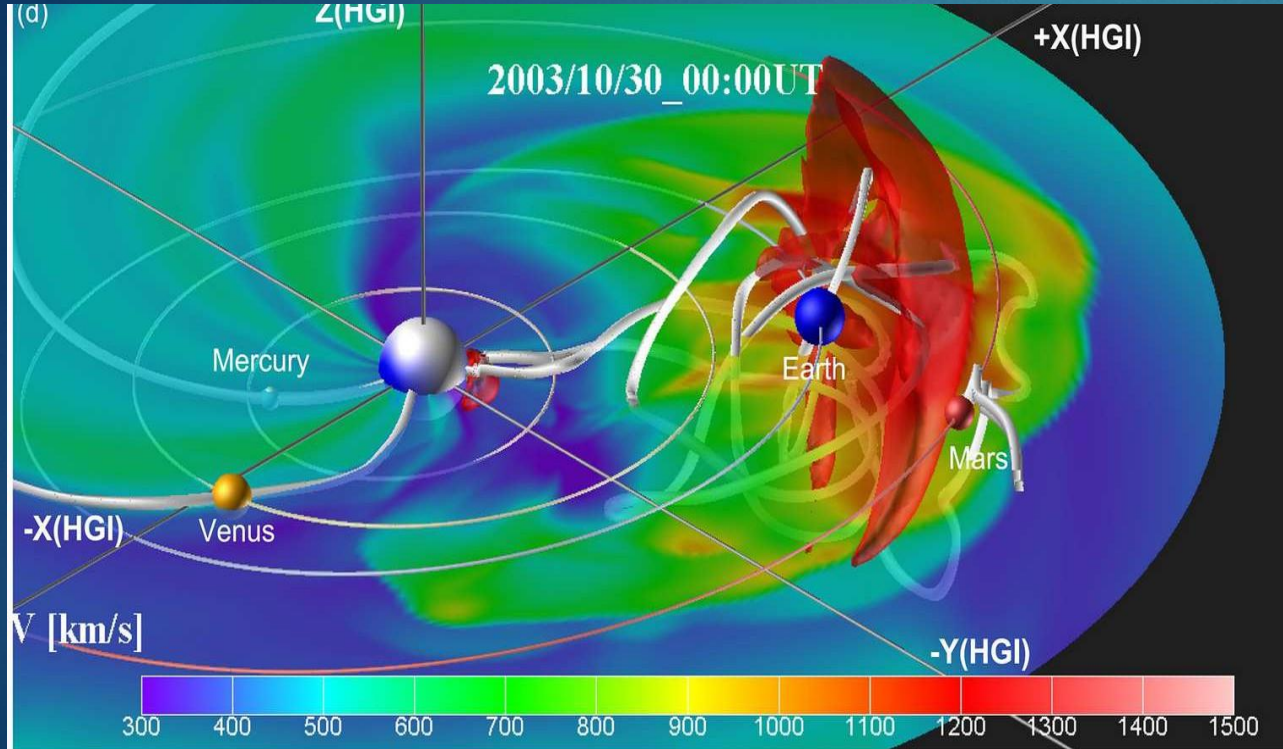
- Camporeale, E. (2019). The challenge of machine learning in space weather: Nowcasting and forecasting. *Space weather*, 17(8), 1166-1207. <https://doi.org/10.1029/2018SW002061>

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- Camporeale, E. (2019). The challenge of machine learning in space weather: Nowcasting and forecasting. Space weather, 17(8), 1166-1207. <https://doi.org/10.1029/2018SW002061>

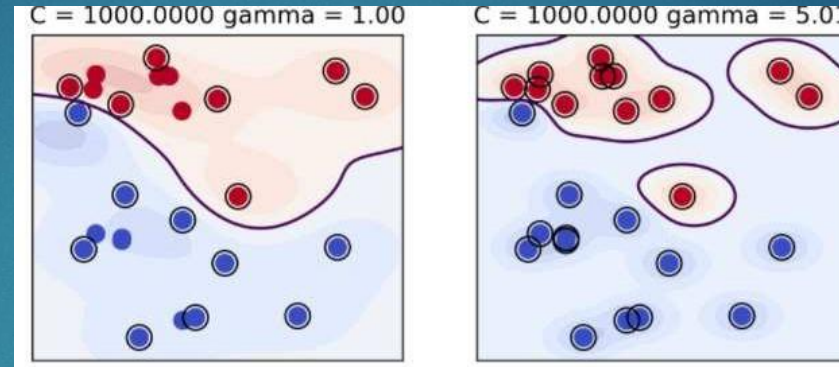
- Shiota, D., & Kataoka, R. (2016). Magnetohydrodynamic simulation of interplanetary propagation of multiple coronal mass ejections with internal magnetic flux rope (SUSANOO-CME). Space Weather, 14(2), 56-75. <https://doi.org/10.1002/2015SW001308>
- Bobra, M. G., & Ilonidis, S. (2016). Predicting coronal mass ejections using machine learning methods. The Astrophysical Journal, 821(2), 127. DOI 10.3847/0004-637X/821/2/127
- Chierichini, S., Liu, J., Korsós, M. B., Del Moro, D., & Erdélyi, R. (2024). CME Arrival Modeling with Machine Learning. The Astrophysical Journal, 963(2), 121. DOI 10.3847/1538-4357/ad1cee

ML: Quick, Nice results vs Correct Dataset use

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- ▶ Neural Networks
- ▶ SVM
- ▶ Random Forest
- ▶ Gaussian models
- ▶ Convolutional NN
- ▶



Flare Forecast

True Skill Score evolution:

➤ Barnes 2008:	0.24
➤ Florios 2018:	0.77
➤ Sun 2022:	0.90
➤ Deshmukh 2022:	0.90

BEWARE: sometimes there are errors in:

The hyperparameter tuning process

How the samples are used

How the cross-validation is used

- Barnes, G., & Leka, K. D. (2008). Evaluating the performance of solar flare forecasting methods. The Astrophysical Journal, 688(2), L107. [DOI 10.1086/595550](https://doi.org/10.1086/595550)
- Florios, K., Kontogiannis, I., Park, S. H., Guerra, J. A., Benvenuto, F., Bloomfield, D. S., & Georgoulis, M. K. (2018). Forecasting solar flares using magnetogram-based predictors and machine learning. Solar Physics, 293(2), 28. <https://doi.org/10.1007/s11207-018-1250-4>
- Sun, Z., Bobra, M. G., Wang, X., Wang, Y., Sun, H., Gombosi, T., ... & Hero, A. (2022). Predicting solar flares using CNN and LSTM on two solar cycles of active region data. The Astrophysical Journal, 931(2), 163. [DOI 10.3847/1538-4357/ac64a6](https://doi.org/10.3847/1538-4357/ac64a6)
- Deshmukh, V., Flyer, N., Van der Sande, K., & Berger, T. (2022). Decreasing false-alarm rates in CNN-based solar flare prediction using SDO/HMI data. The Astrophysical Journal Supplement Series, 260(1), 9. [DOI 10.3847/1538-4365/ac5b0c](https://doi.org/10.3847/1538-4365/ac5b0c)

Flares: Next steps?



More info from B → by using topological descriptors



Add the Time dimension → use B movies



Higher layers info → use B + chromosphere + corona



Adapt the Loss Functions → redefine error types



Explicability of Deep Learning → analysis of attention frames



Change of paradigm (e.g. Physics-Informed ML, ...)



Use of topological descriptors

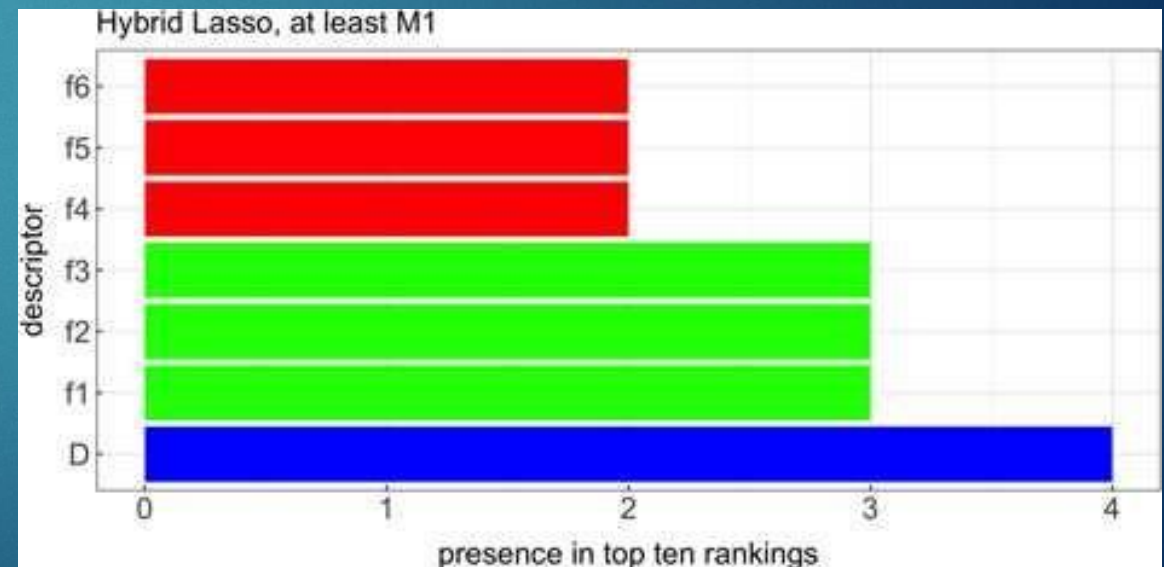
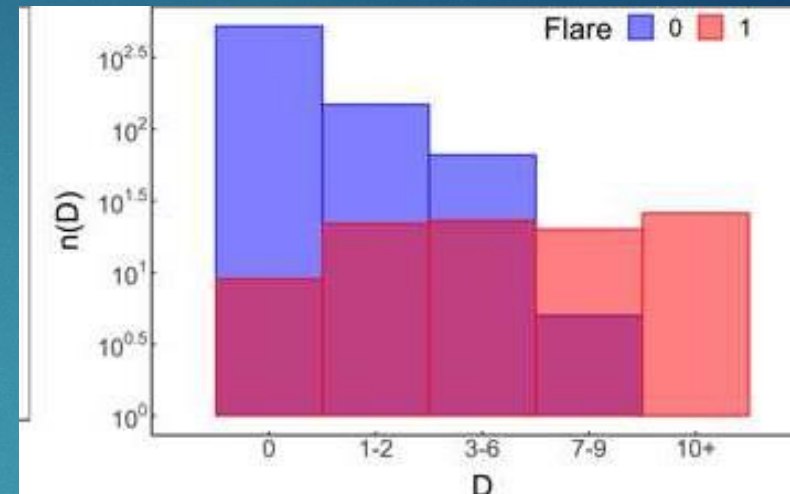
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YET another parameter from magnetic polarity inversion lines (PILs)

Topologically complex ARs are strongly correlated to flare emissions.

Therefore a topological descriptor that counts the number of separated **PILs** fragments in the ARs.

Feature ranking analysis tells us it is extremely relevant to reach high skill scores → it brings relevant information on the AR flaring potential.



Analysis of time sequences

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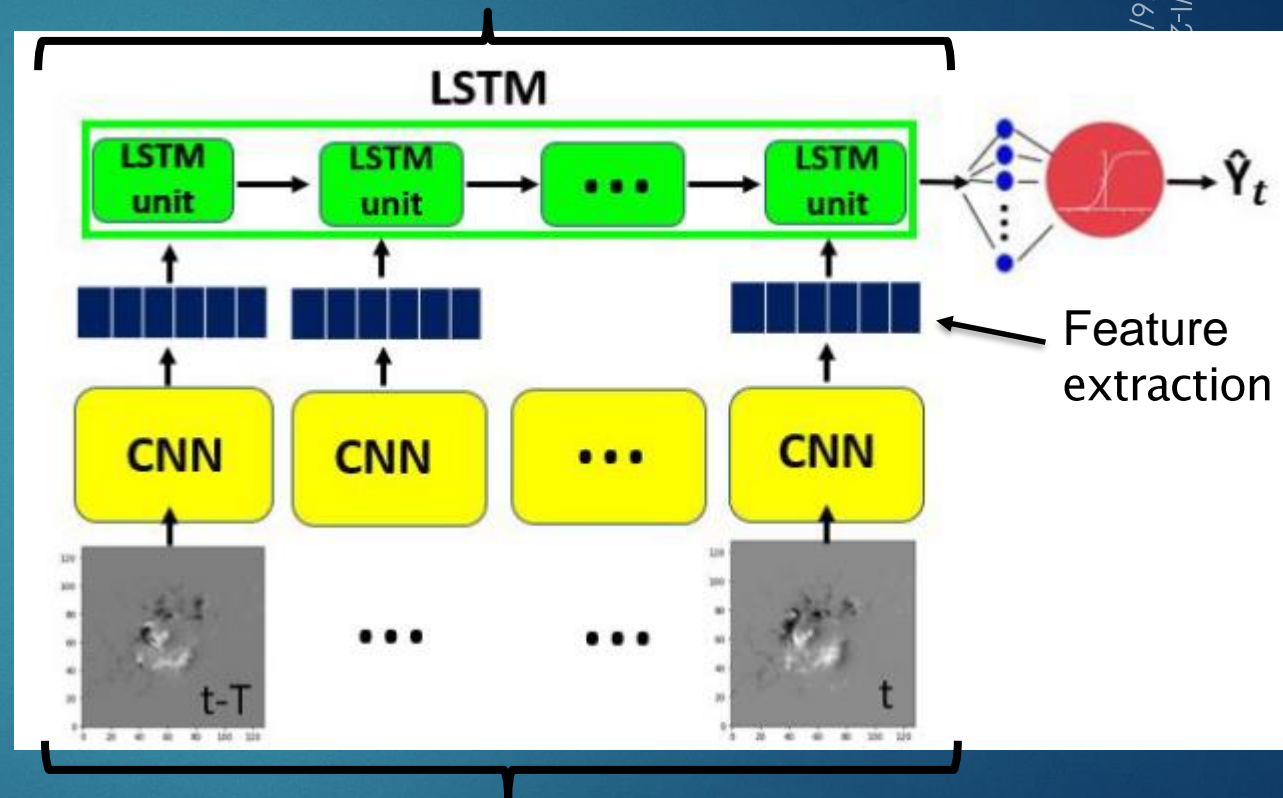
Long-term Recurrent Neural Network
LRCN = CNN + LSTM

BUT they get **similar TSS** values as more standard methods

→ HMI data do not contain enough information?

→ flares' stochasticity hampers the possibility of binary predictions, in favor of a probabilistic prediction?

Considering the “memory” of the process



Input: videos of HMI magnetograms

Guastavino, S., Marchetti, F., Benvenuto, F., Campi, C., & Piana, M. (2022). Implementation paradigm for supervised flare forecasting studies: A deep learning application with video data. *Astronomy & Astrophysics*, 662, A105.

<https://doi.org/10.1051/0004-6361/202243617>

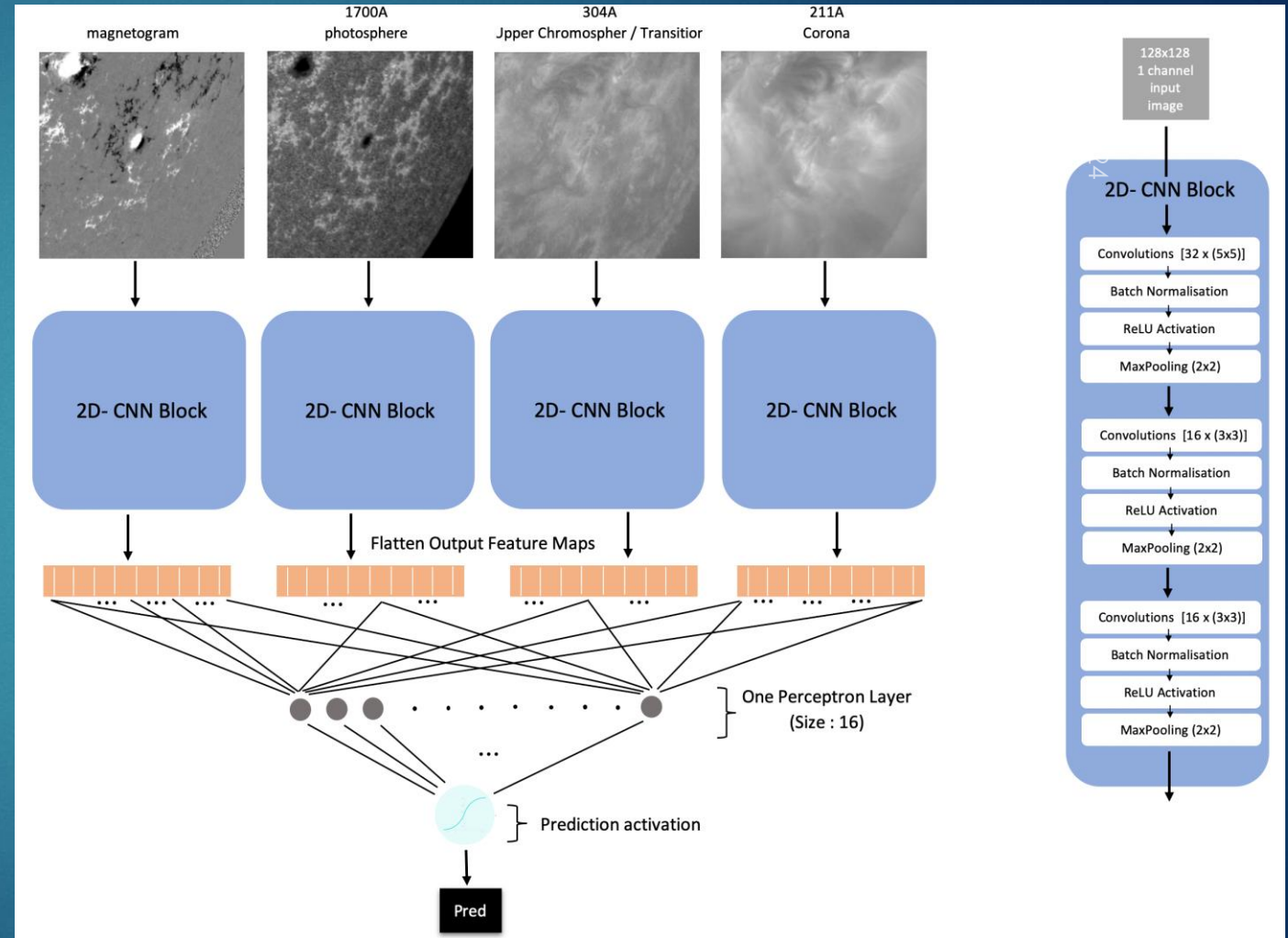
Campi, C., Benvenuto, F., Massone, A. M., Bloomfield, D. S., Georgoulis, M. K., & Piana, M. (2019). Feature ranking of active region source properties in solar flare forecasting and the uncompromised stochasticity of flare occurrence. *The Astrophysical Journal*, 883(2), 150. DOI 10.3847/1538-4357/ab3c26

Use of more layers

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Each layer used contributes to improved performances

Models combining SDO/AIA EUV images as inputs show improved performances compared to employing SDO/HMI photospheric magnetograms alone



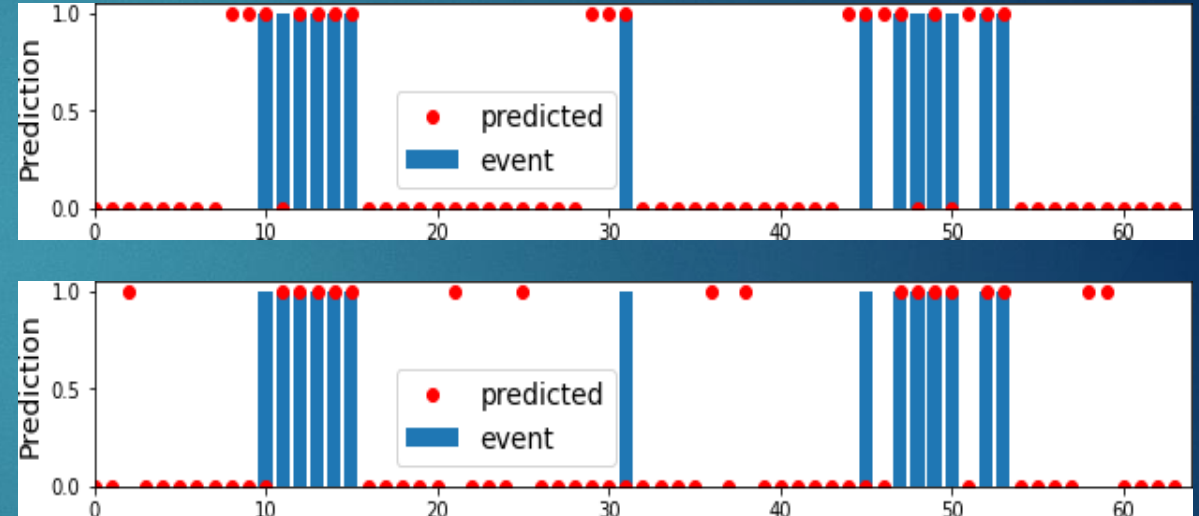
Redefine error types/Performance Scores

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Classical skill scores do not take into account the temporal distribution of the prediction

A possible solution:

define new skill scores which allow ranking the prediction errors on the basis of their distribution along time



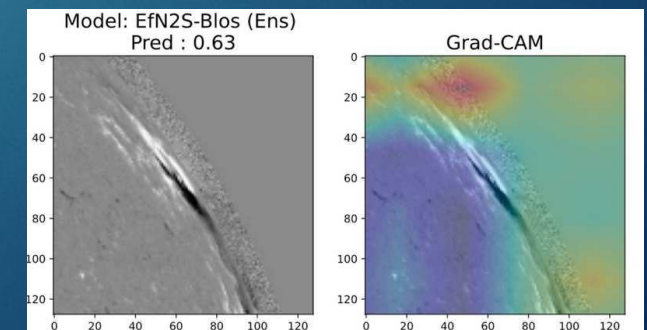
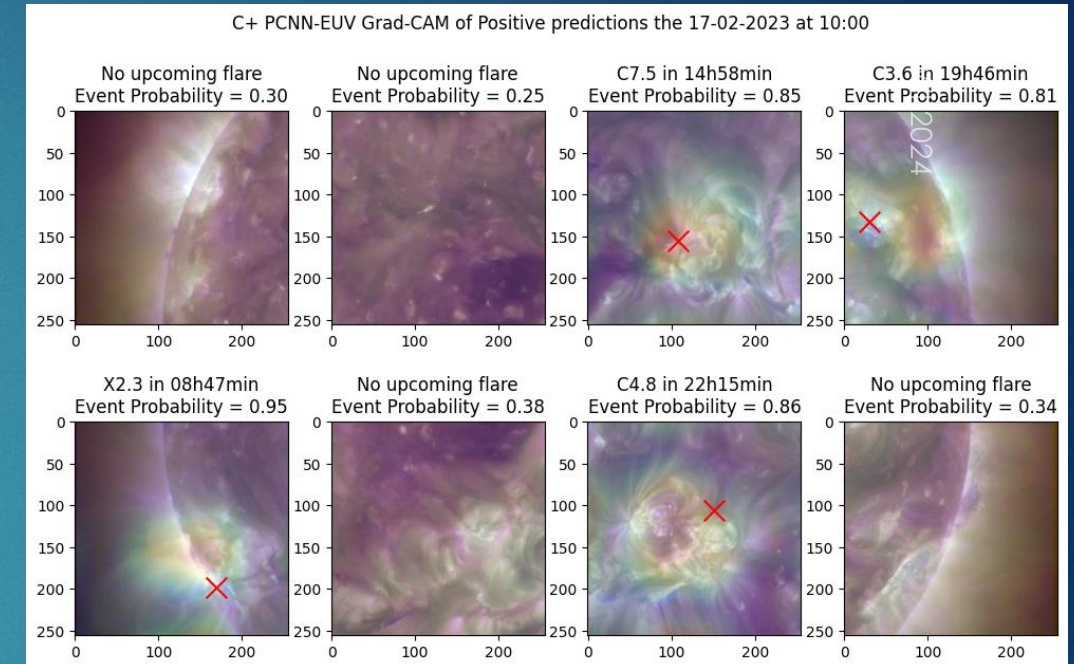
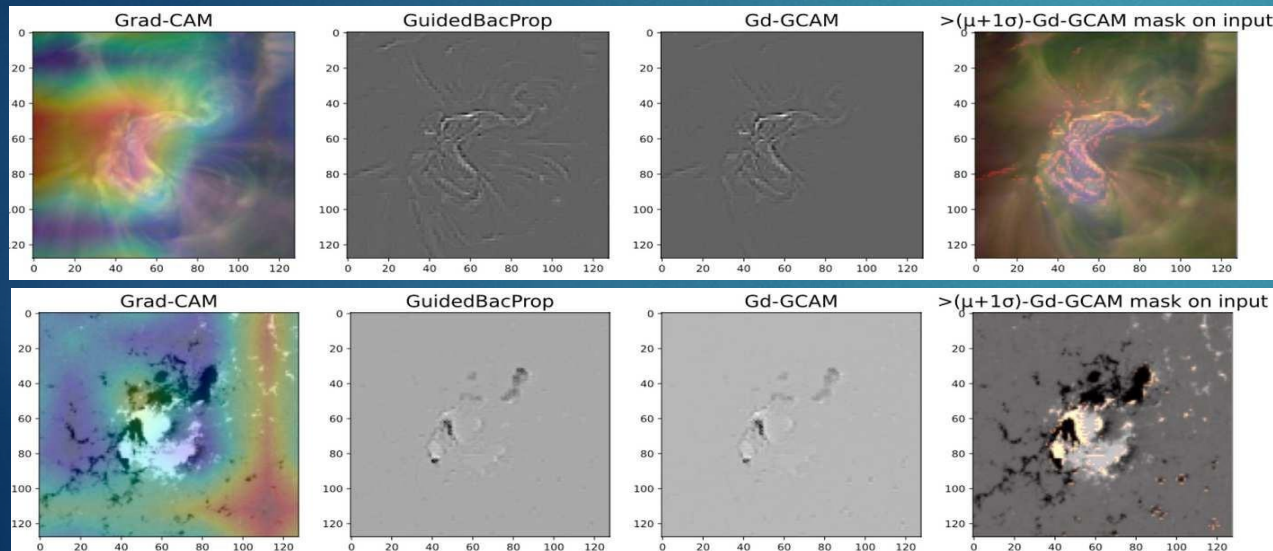
Two forecasts: same performance, but totally different predictive value.

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Interpretability of CNN

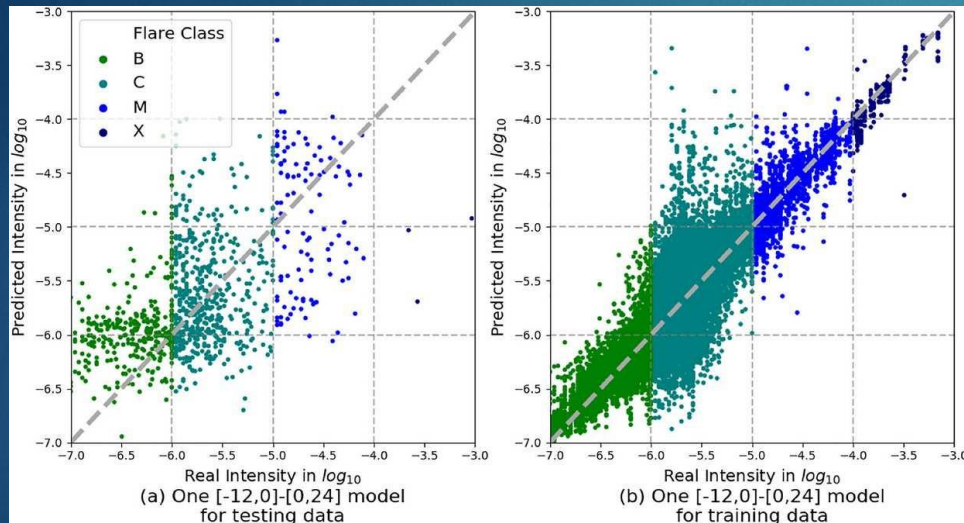
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Identify the most class-discriminative regions
+
the pixels contributing the most to the last convolutional layer



Change of paradigm

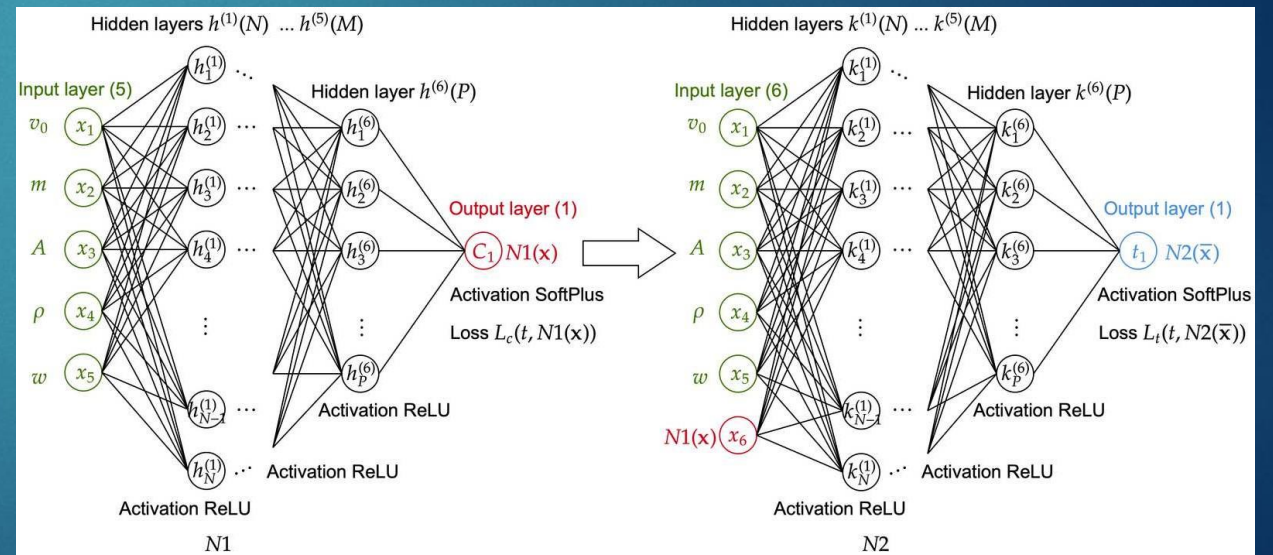
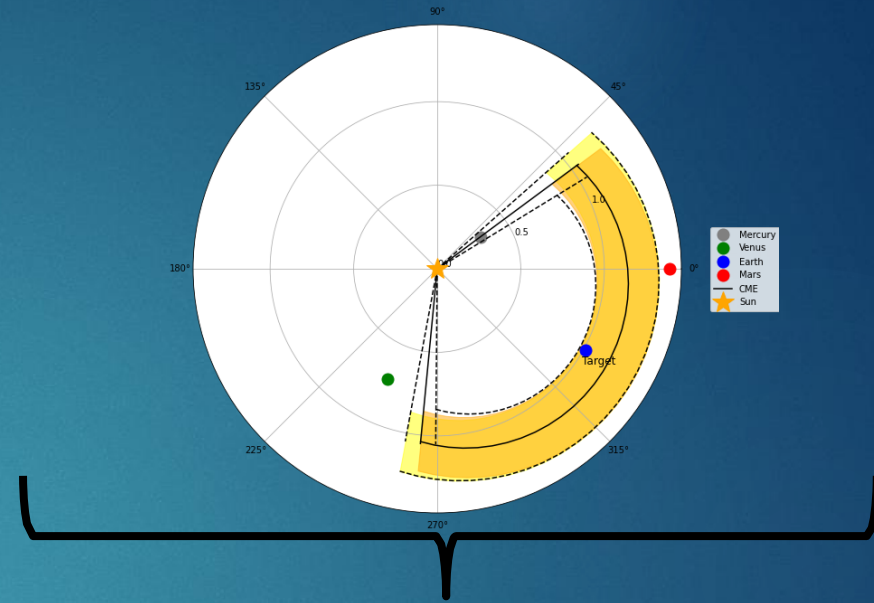
Regression instead of Classification
+
Put the Physics in the ML model



Jiao, Z., Sun, H., Wang, X., Manchester, W., Gombosi, T., Hero, A., & Chen, Y. (2020). Solar flare intensity prediction with machine learning models. *Space weather*, 18(7), e2020SW002440.

Guastavino, S., Candiani, V., Bemporad, A., Marchetti, F., Benvenuto, F., Massone, A. M., ... & Michele, P. (2023). Physics-driven machine learning for the prediction of coronal mass ejections' travel times. *The Astrophysical Journal*, 954(2), 151.

Chierichini, S., Francisco, G., Mugatwala, R., Foldes, R., Camporeale, E., De Gasperis, G., ... & Erdelyi, R. (2024). A Bayesian approach to the drag-based modelling of ICMEs. *Journal of Space Weather and Space Climate*, 14, 1.



Beyond the Black Box: What is next?

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ML is cheap!



ML is growing fast!



ML is fixing its (major) drawbacks



The data(-sets) are more and more important



ML is here to stay ?



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Thank You!

Extra Slides



CONCEPTS OF CONVOLUTIONAL NN

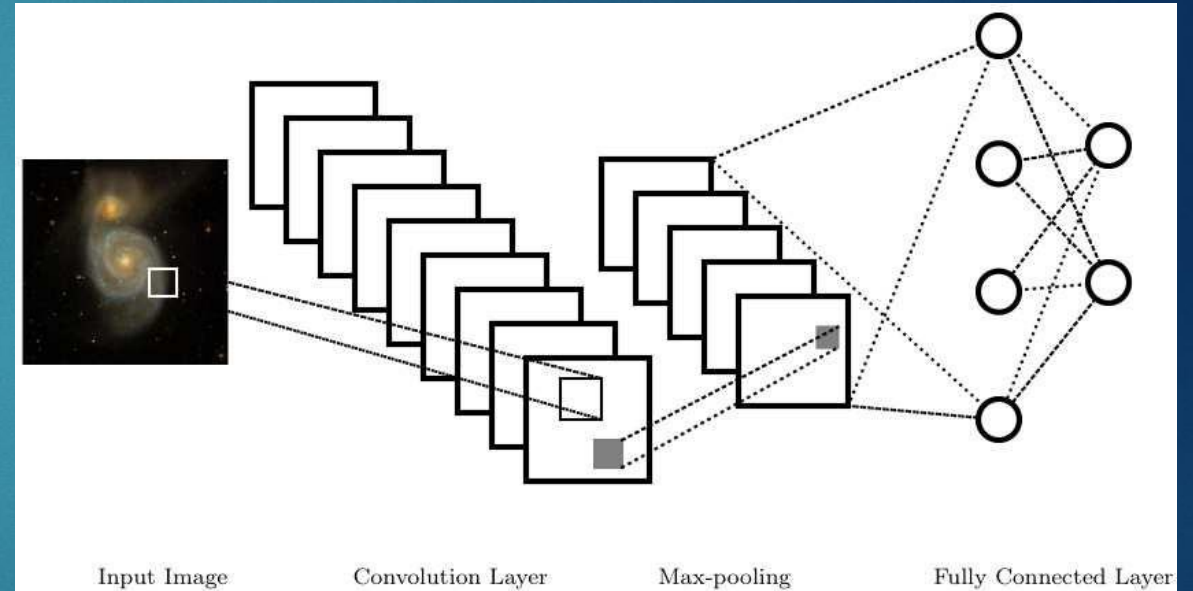
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NN designed to work with images and/or regularly sampled data

Components:

- Convolutional layer
- Non-linear activation functions
- Pooling/downsampling
- Fully connected layer for classification

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CONCEPTS OF LONG-TERM SHORT-TERM MEMORY

h_{t-1} - hidden state at previous timestep t-1 (short-term memory)

c_{t-1} - cell state at previous timestep t-1 (long-term memory)

x_t - input vector at current timestep t

h_t - hidden state at current timestep t

c_t - cell state at current timestep t

\otimes - vector pointwise multiplication $+$ - vector pointwise addition

\tanh - tanh activation function

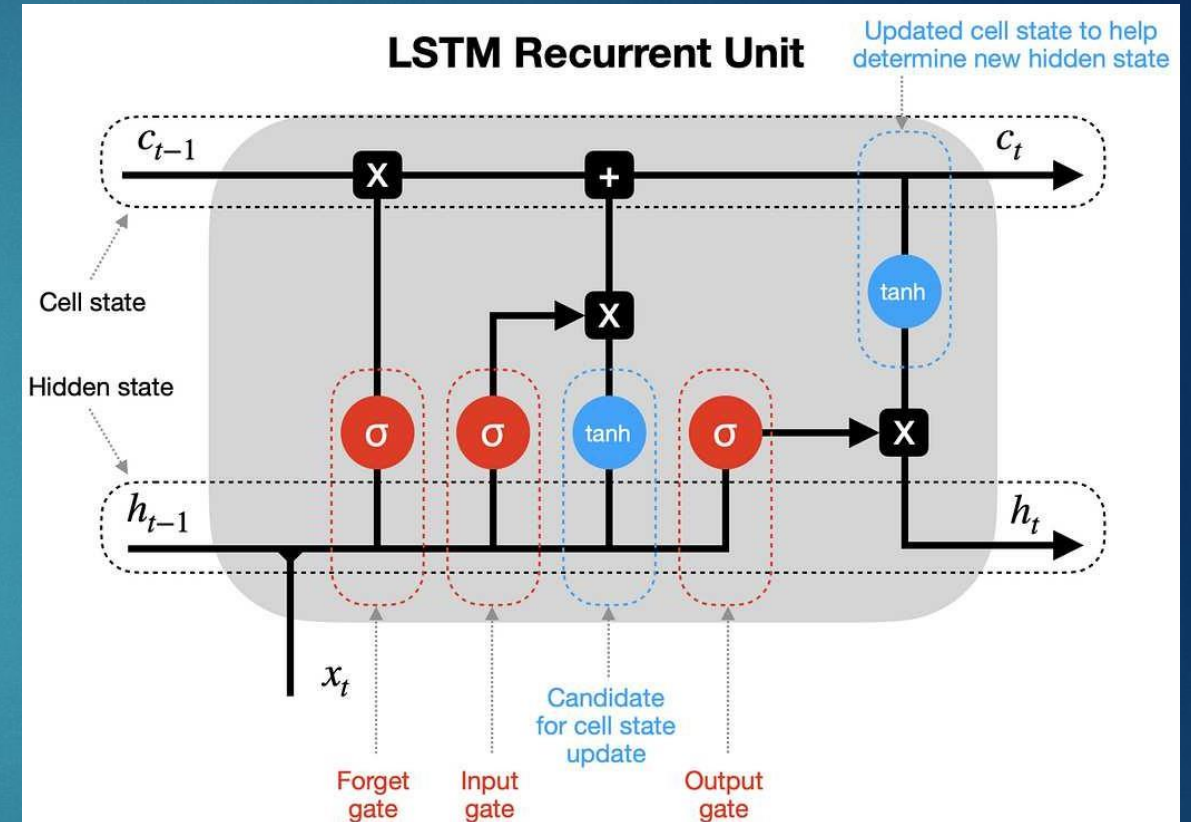
σ - sigmoid activation function

\parallel - concatenation of vectors

--- - states

--- - gates

--- - updates



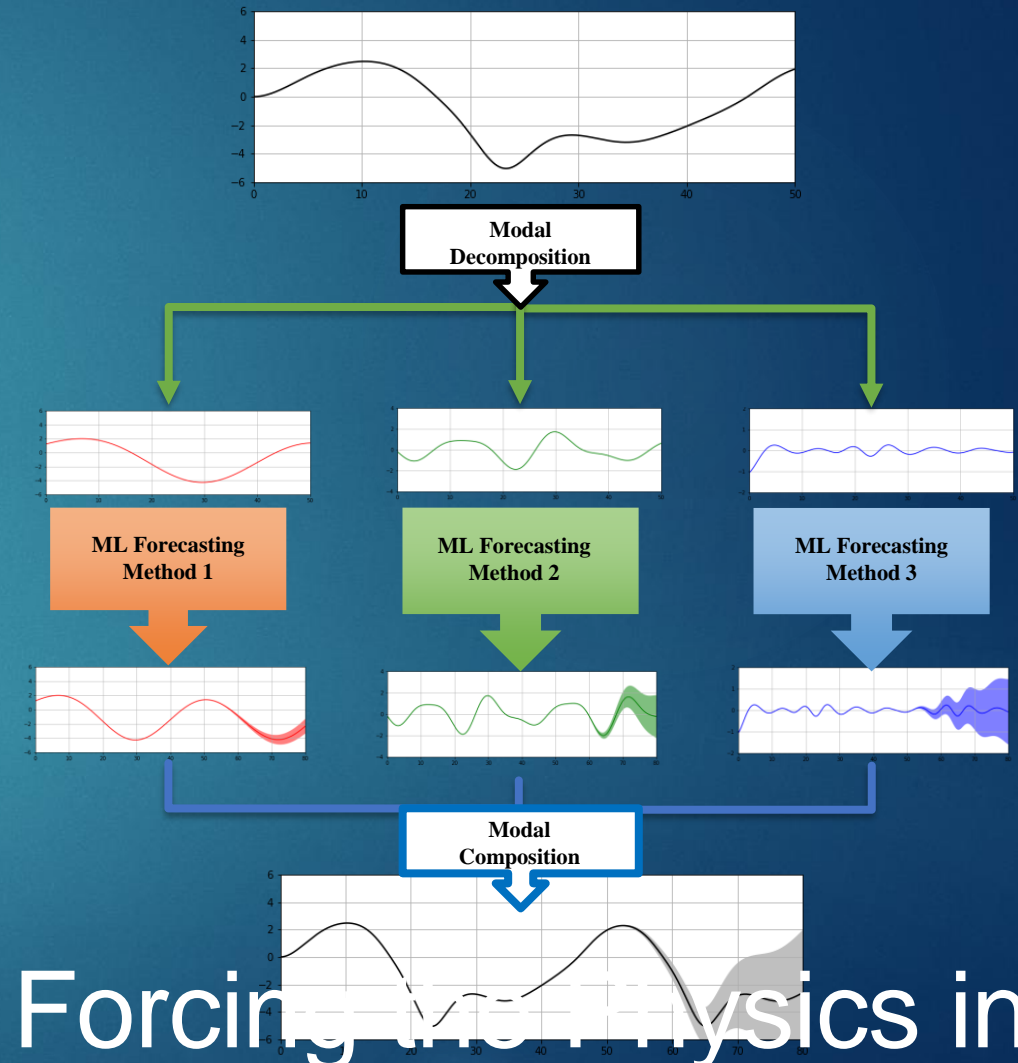
Forcing the Physics into ML approach

A complex signal, with many different timescales embedded.

→ We can try to mix different proxies to reproduce/forecast the signal.

OR:

→ We separate the signal's relevant scales into different modes, then reproduce/forecast the different modes mixing the proxies' modes (at or near the same timescales). Then, we reproduce/forecast the signal by combining the reproduced/forecast modes.



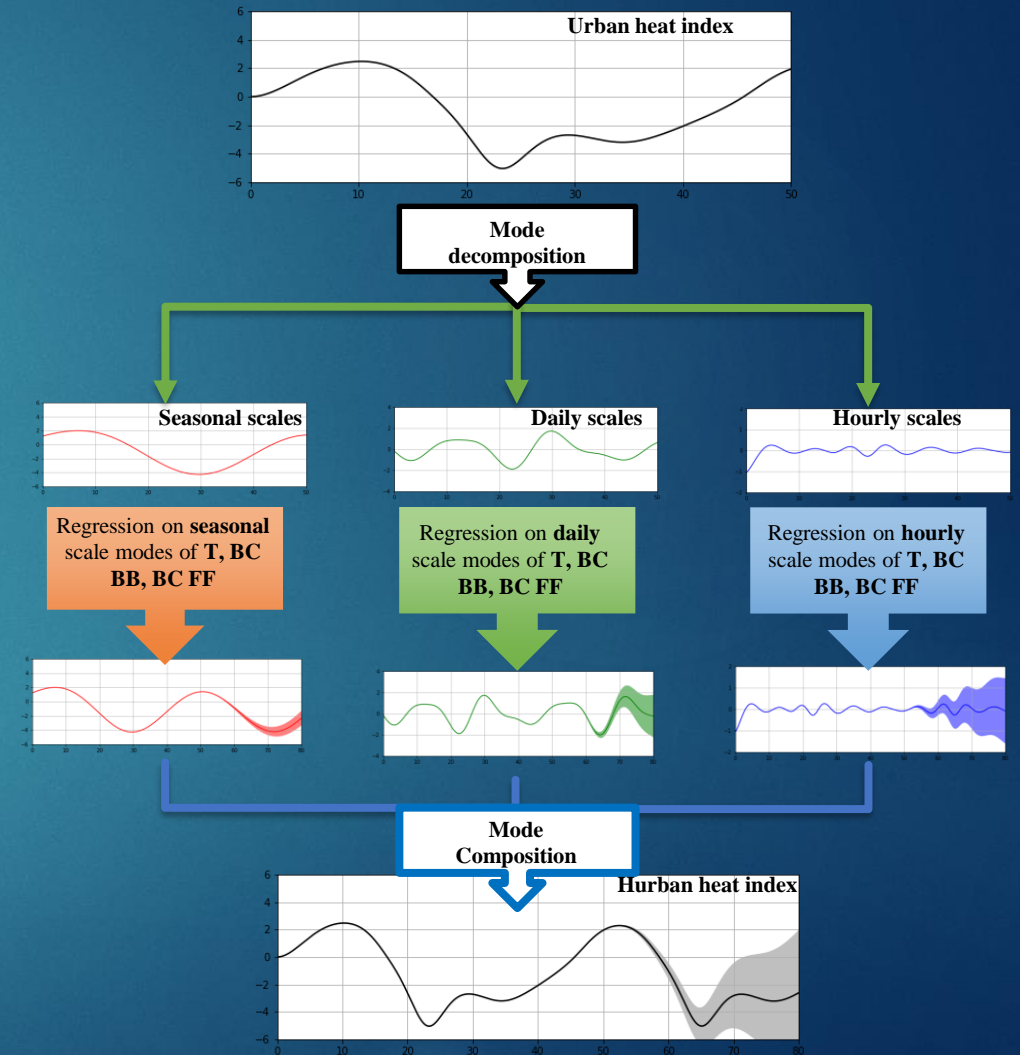
- Materassi, M.; Alberti, T.; Migoya-Oru , Y.; Radicella, S.M.; Consolini, G. Chaos and Predictability in Ionospheric Time Series. Entropy 2023, 25, 368. <https://doi.org/10.3390/e25020368>
- Reda, R., Stumpo, M., Giovannelli, L. et al. Disentangling the solar activity-solar wind predictive causality at Space Climate scales. Rend. Fis. Acc. Lincei 35, 49–61 (2024). <https://doi.org/10.1007/s12210-023-01213-w>
- Stumpo M, Consolini G, Alberti T et al (2020) Measuring Information Coupling between the Solar Wind and the Magnetosphere-Ionosphere System. Entropy 22(3):276. <https://doi.org/10.3390/e22030276>

Forcing the physics into
approach of CN

Applied to Black Carbon

An exercise: we know what we should get, we know how to interpret the results

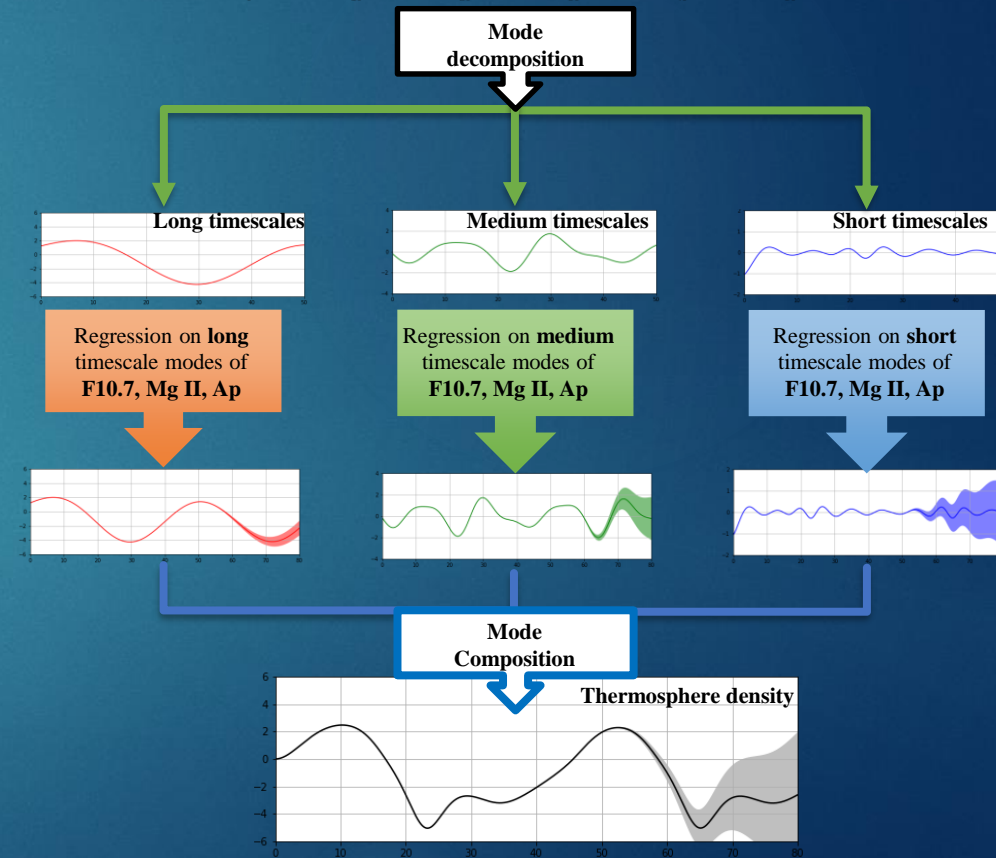
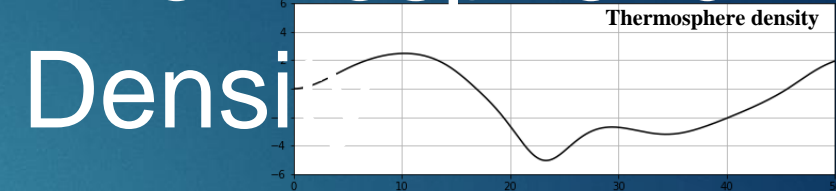
We can verify which quantities and which modes are more relevant to reproduce the different timescales of the original signal by permutation analysis.



Applied to Thermospheric Density

A test: we expect the outcome and we interpret the results

We can learn which quantities and which modes are more relevant to reproduce the different timescales of the original signal by permutation analysis.



- Cafiero, D. (2024 in preparation) From solar and geomagnetic indices to thermosphere density Master Thesis U. Rome "Tor Vergata"
- Bigazzi, A., Cauli, C., & Berrilli, F. (2020, June). Lower-thermosphere response to solar activity: an empirical-mode-decomposition analysis of GOCE 2009–2012 data. In *Annales Geophysicae* (Vol. 38, No. 3, pp. 789–800). Copernicus GmbH. <https://doi.org/10.5194/angeo-38-789-2020>
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