THE INTERNATIONAL SPACE WEATHER INITIATIVE (ISWI)

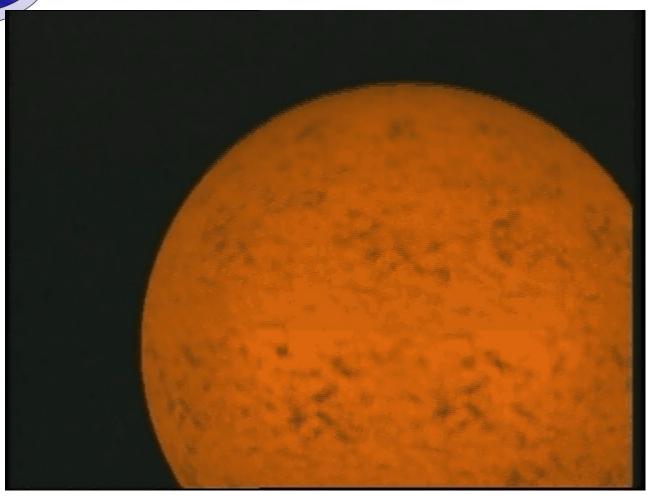
A FOLLOW-ON TO THE INTERNATIONAL HELIOPHYSICAL YEAR (IHY)

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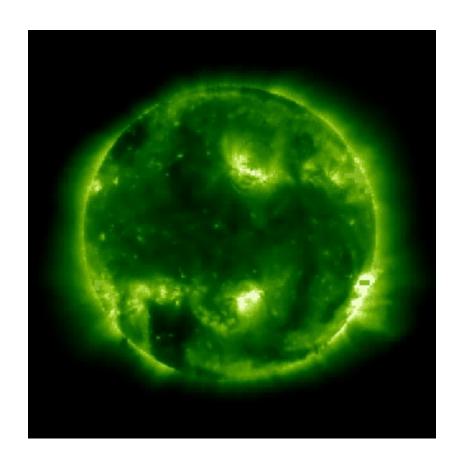
February 2011

Solar Magnetic Dynamo



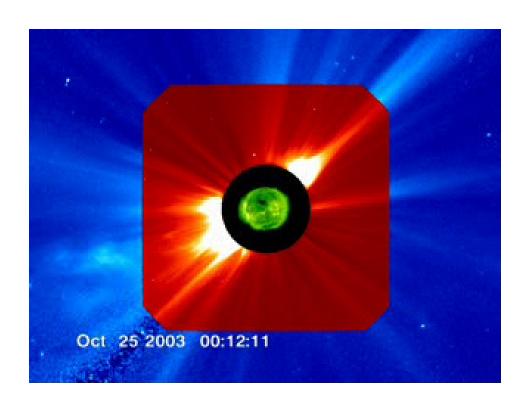


Coronal Activity



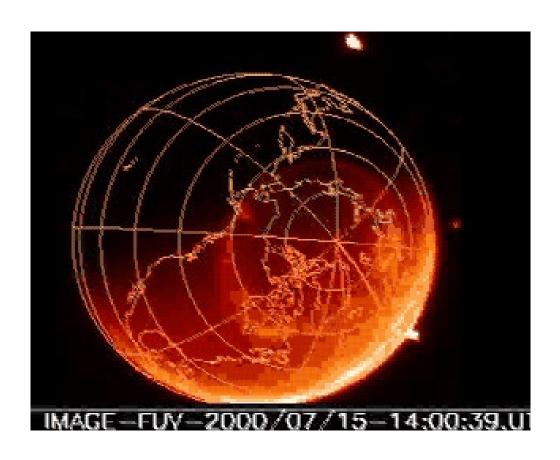


Space Weather





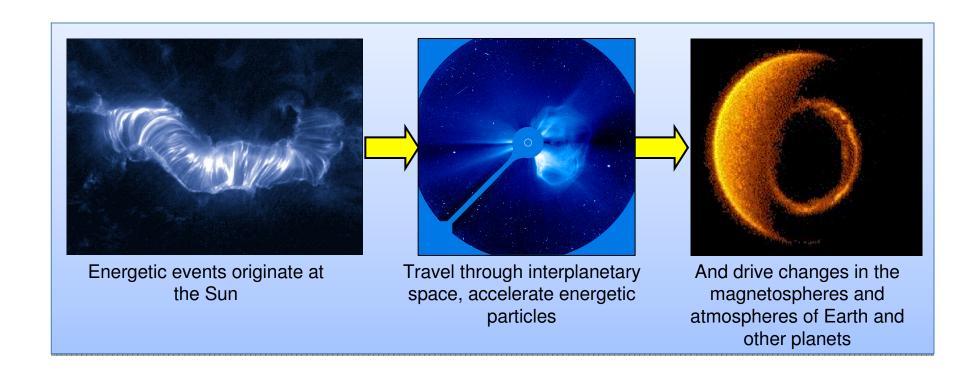
Earth Impact





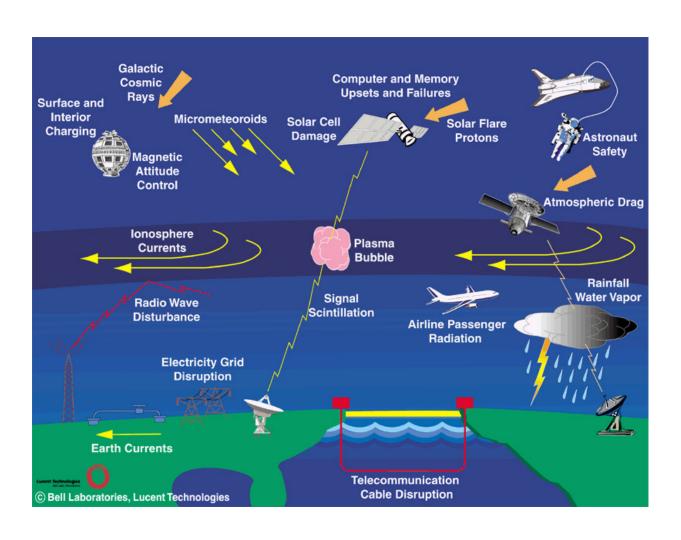
Key Point

Solar events influence planetary and interplanetary environments





ISWI



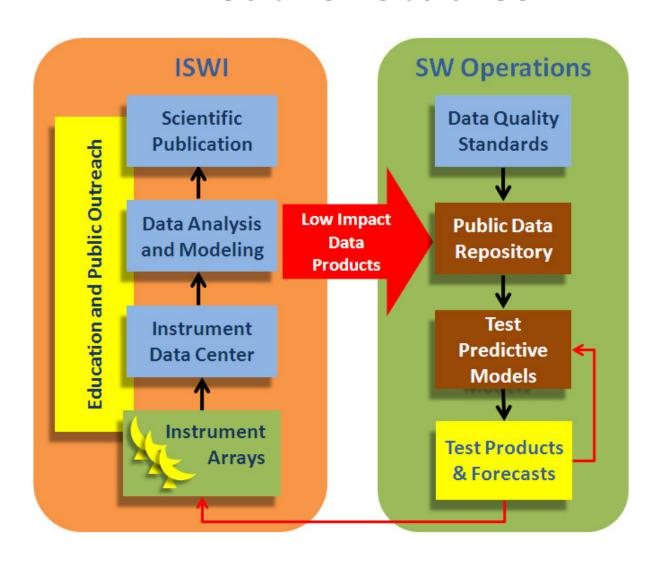


ISWI Objectives

- Develop the scientific insight necessary to understand the science, and to reconstruct and forecast near-Earth space weather
 - Instrumentation
 - Expand and continue deployment of new and existing instrument arrays
 - Data analysis
 - Expand data analysis effort for instrument arrays and existing data bases
 - Coordinate data products to provide input for physical modeling
 - Input instrument array data into physical models of heliospheric processes
 - Develop data products that reconstruct past conditions in order to facilitate assessment of problems attributed to space weather effects
 - Coordinate data products to allow predictive relationships to be developed
 - Develop data products to allow predictive relationships that enable the forecasting of Space Weather to be established
 - Develop data products that can easily be assimilated into real-time or near real-time predictive models
- Education, Training, and Public Outreach
 - University and Graduate Schools
 - Encourage and support space science courses and curricula in Universities that provide instrument support
 - Public Outreach
 - Develop public outreach materials unique to the ISWI, and coordinate the distribution

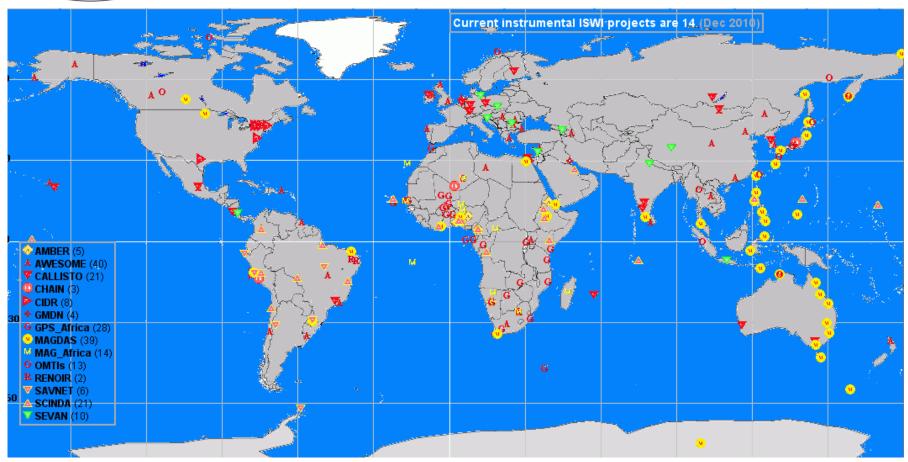


ISWI Contributes to Space Weather Studies





IHY Participation



- 14 Distributed instrument teams observatory program
- ~1000 participating locations



Principles of the ISWI Instrument Program

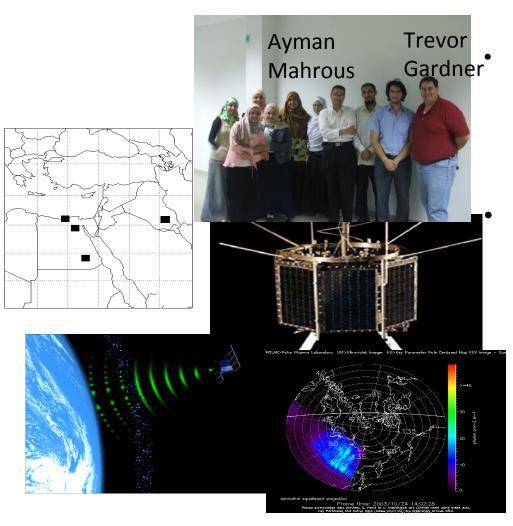
- The lead scientist or principle investigator funded by his/her country provides instrumentation (or fabrication plans) and data distribution service
- The host country provides the workforce, facilities, and operational support typically at a local university or research institute.
- Host scientists become part of science team
- All data and data analysis activity is shared
- All scientists participate in publications and scientific meetings where possible



Scientific Benefits: Why do this?

- By observing in new geographical regions, a more global picture of Earth's response to solar wind inputs can be obtained
- 24/7 solar observing in radio and H-alpha
- Arrays provide 3D information that can be used in tomographic reconstructions
- Long term these networks will provide real-time data valuable for forecasting and nowcasting
- Modeling improvements will allow better exploitation of existing data sets

Coherent Ionospheric Doppler Receivers (CIDRs)



Radio Altimetry and Ephemeris Satellites

- 150/400 MHz Radio Beacon
- Ionospheric TEC Correction Data
- Egypt-Trevor Gardner (UTAr)

Advantages over GPS

- More accurate, no need for plasmaspheric corrections by using LEO satellites.
- Can measure the spatial structure of the ionosphere.
- A powerful tool for topographic image of the ionosphere

Current Instrument Arrays

ID	INSTRUMENT	Lead Scientist	Country	Objective
1	Scintillation Network Decision Aid (SCINDA)	K. Groves keith.groves@hanscom.af.mil (Hanscom AFRL)	USA	Study equatorial ionospheric disturbances to aid in the specification and prediction of communications degradation due to ionospheric scintillation in the Earth's equatorial region
2	Ionospheric Tomography Network of Egypt (ITNE) Coherent Ionospheric Doppler Receiver (CIDR)	A. Mahrous amahrous@helwan.edu.eg (Helwan University) T. Garner garner@arlut.utexas.edu (University of Texas)	USA	To tomographically reconstruct the ionosphere and to provide input to data assimilation models
3	Atmospheric Weather Education System for Observation and Modeling of Effects (AWESOME) and Sudden Ionospheric Disturbance monitor (SID)	U. Inan inan@stanford.edu M. Cohen mcohen@stanford.edu D. Scherrer deborah@solar2.stanford.edu (Stanford University)	USA	Lightning, sprites, elves, relation to terrestrial gamma ray flashes, whistler induced electron precipitation, conjugate studies. Education and public outreach.
4	Remote Equatorial Nighttime Observatory for Ionospheric Regions (RENOIR)	J. Makela <u>imakela@illinois.edu</u> (University of Illinois)	USA	Study the equatorial/low-latitude ionosphere/thermosphere system, its response to storms, and the irregularities that can be present on a daily basis
5	African GPS Receivers for Equatorial Electrodynamics Studies (AGREES)	E. Yizengaw (Boston College) ekassie@igpp.ucla.edu M. Moldwin (University Mich)	USA	Understand unique structures in equatorial ionosphere, low/mid latitude plasma production, effect of ionospheric and plasmaspheric irregularities on communications
6	African Meridian B-field Education and Research (AMBER)	M. Moldwin (University Mich) mmoldwin@igpp.ucla.edu E. Yizengaw (Boston College)	USA	Understand low latitude electrodynamics, ULF pulsations, effect of Pc5 ULF on MeV electron population in inner radiation belts

Current Instrument Arrays (continued)

ID	INSTRUMENT	LeadScientist	Country	Objective
7	Compound Astronomical Low-cost Low-frequency Instrument for Spectroscopy and Transportable Observatory (CALLISTO)	A.Benz benz@astro.phys.ethz.ch C. Monstein monstein@astro.phys.ethz.ch (ETH)	Switzerland	Study of radio flares caused by solar activity in view of space weather and climate change
8	South Atlantic Very Low frequency Network (SAVNET)	JP. Raulin raulin@craam.mackenzie.br (University Presbiteriana)	Brazil	Study of the SAMA region at low ionospheric altitudes and its structure and dynamics during geomagnetic perturbations
9	Magnetic Data Acquisition System (MAGDAS)	K. Yumoto yumoto@serc.kyushu-u.ac.jp (Kyushu University)	Japan	Study of dynamics of geospace plasma changes during magnetic storms and auroral substorms, the electromagnetic response of iono-magnetosphere to various solar wind changes, and the penetration and propagation mechanisms of DP2-ULF range disturbances
10	African Dual Frequency GPS Network	C. Amory-Mazaudier <u>christine.amory@lpp.polytechnique.fr</u> (CETP/CNRS)	France	To increase the number of real-time dual- frequency GPS stations worldwide for the study of ionospheric variability, response of the ionospheric total electron content (TEC) during geomagnetic storms over the African sector

Current Instrument Arrays (continued)

ID	INSTRUMENT	LeadScientist	Country	Objective
11	Space Environmental Viewing and Analysis Network (SEVAN)	A.Chillingarian <u>chili@aragats.am</u> (Aragats University)	Armenia	A network of particle detectors that aims to improve fundamental research of the particle acceleration in the vicinity of the Sun and the space environment, as well as to provide forewarnings of dangerous consequences of space storms
12	Global Muon Detector Network (GMDN)	K. Munakata kmuna00@shinshu-u.ac.jp (Shinsu University)	Japan	To identify the precursory decrease of cosmic ray intensity that takes place more than one day prior to the Eartharrival of shock driven by an interplanetary coronal mass ejection
13	Flare Monitoring Telescopes (FMT) under the Continuous H-alpha Imaging Network (CHAIN)	S. UeNo <u>ueno@kwasan.kyoto-u.ac.jp</u> K. Shibata (Kyoto University)	Japan	Time variation and 3D velocity field of solar activity, flares, filament eruptions and shock waves (Moreton waves) by using multi-wavelength H-alpha images of the full-disk Sun.
14	Optical Mesosphere Thermosphere Imagers (OMTIs)	K. Shikawa shiokawa@stelab.nagoya-u.ac.jp (Nagoya University)	Japan	Dynamics of the upper atmosphere through nocturnal airglow emissions http://stdb2.stelab.nagoya-u.ac.jp/omti/index.html



UN-NASA Workshop Series on IHY

- First Workshop (2005)
 - UN, ESA, NASA, and UAE sponsored
- Second Workshop (2006)
 - UN, NASA and India sponsored
- Third Workshop (2007)
 - UN, JAXA, NASA and Japan
- Fourth Workshop (2008)
 - UN, ESA, NASA, JAXA, and Bulgaria
- Fifth Workshop (2009)
 - UN, ESA, NASA, and South Korea

These Workshops have been highly successful at establishing new collaborations between instrument providers and hosts









1st UN ISWI Meeting in Egypt

- November 6-10, 2010
- Hosted by Prof. Ayman Mahrous at Helwan University (about 30 km south of Cairo)
- About 200 registered participants
- Future Workshops planned for
 - •Nigeria 2011
 - •Equador 2012 workshops.





ISWI Space Science Schools



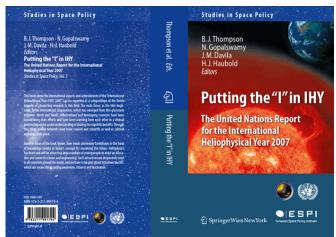
- 6 Schools organized during the IHY
- First ISWI Space Science School in Bahir Dar University November 2010 in Ethiopia
- This year:
 - 1. Slovakia Summer
 - 2. Kinshasa, RDC Sept
 - 3. Rabat, Morocco Dec 5-16
 - 4. Lagos, Nigeria August
 - 5. Abuja, Nigeria (GNSS) Oct
 - 6. Mumbai, India (VLF)- Oct

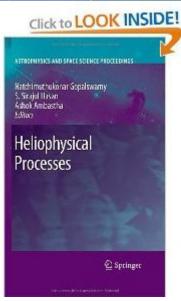


ISWI Space Science School Financial Sponsors



ISWI Space Science Publications





- ISWI schools follow the tradition begun during the IHY
- Two books
 - Putting the I in the IHY:
 describes results of the IHY
 - Heliophysical Processes:
 lectures presented at the school in India



What is Next?

- Continue to identify appropriate sites for instrument new deployments
- Identify additional instruments for deployment
 - We are interested in adding additional experiments to the current list of 14
- Begin to use this new data in modeling and predictions
 - Just starting up, help us do this right
- Additional information
 - http://iswi-secretariat.org
 - Twitter: ISWINews