A Geophysical approach to assess Space Weather impacts on Earth

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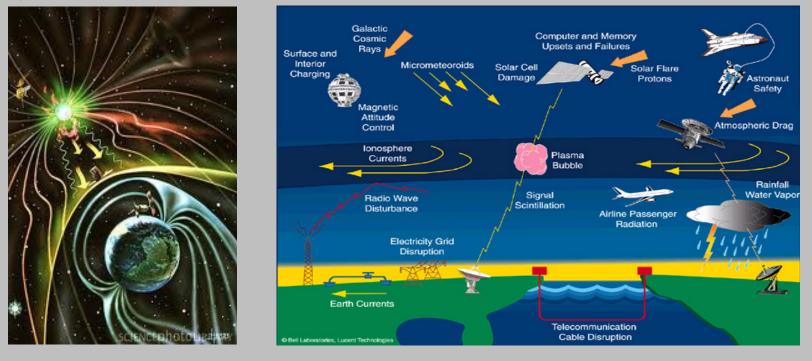


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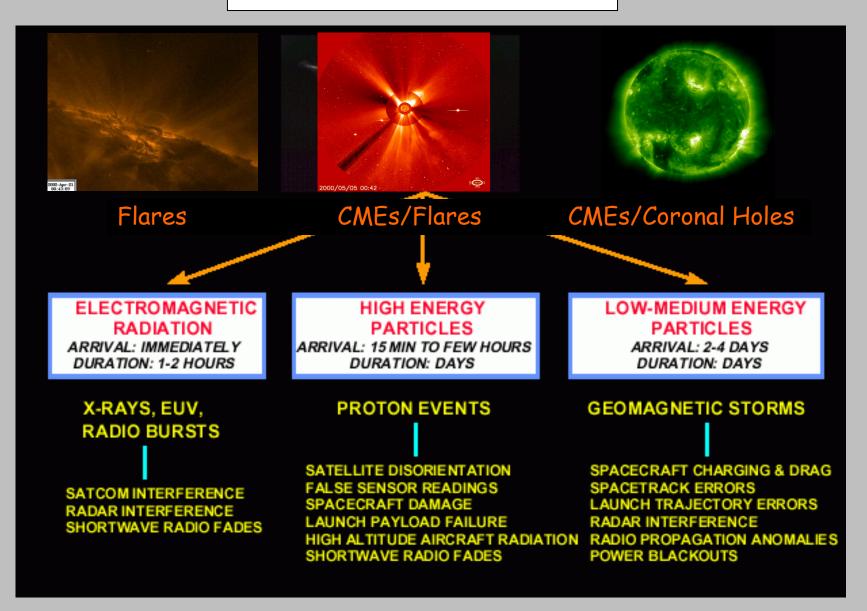
Space Weather

The term "space weather" refers to conditions on the Sun and in the solar wind, magnetosphere, ionosphere, and thermosphere that can influence the performance and reliability of space-borne and ground-based technological systems and that can affect human life or health.

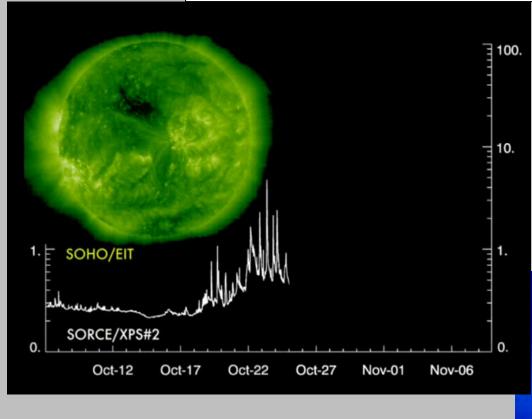


The underlying physics is not sufficiently well understood, and thus we do not have **appropriate warning tools** at hands. Further fundamental research is needed in order to finally understand the processes that tie our home planet to its parent star, the Sun.

Space Weather



Space Weather: causes

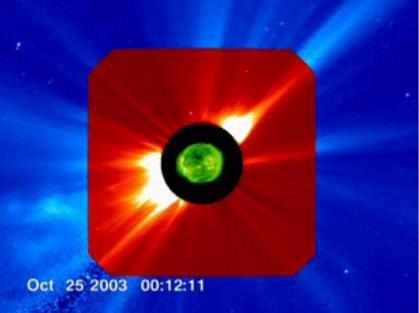


← 07/10/2003 to

to 09/11/2003 →

CMEs are fundamental for predictions on Space Weather conditions

Coronal Mass Ejection (CME) $1 \text{ AU} = 150 \ 10^6 \text{ km} \sim 110 \text{ solar } \emptyset$ Arrival time at $1 \text{ AU} \sim 1.5 - 4 \text{ days}$



Space Weather: growing demand on SW products

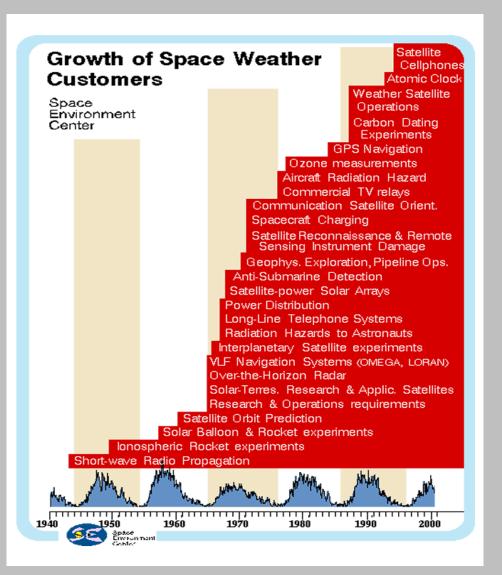
Our society is much more dependent on technology than ever before.

The most rapidly growing sector of the communication market is satellite based:

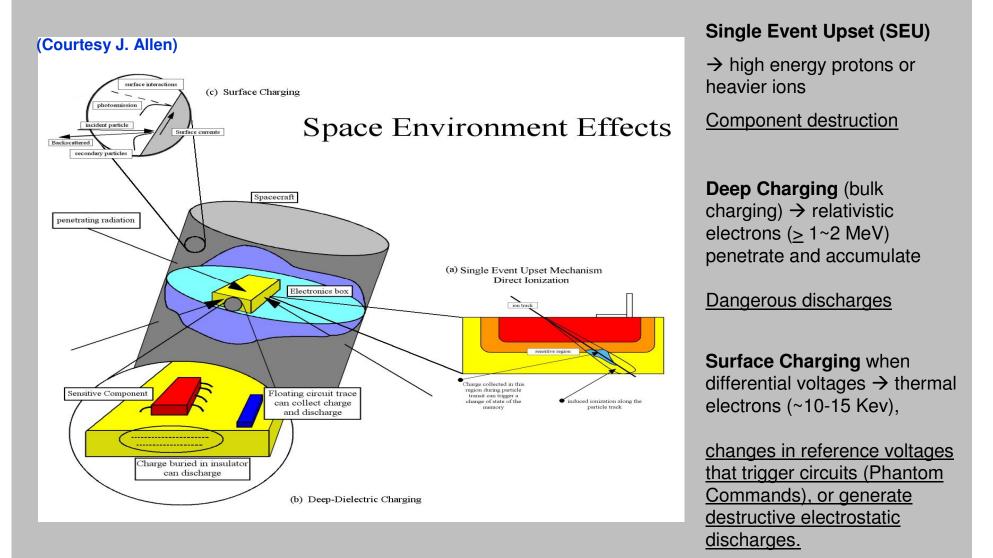
broadcast TV/Radio long-distance telephone service, cell phones, pagers internet, finance transactions

Change in technology: more sensitive payloads high performance components lightweight and low cost

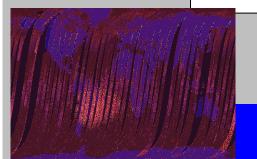
Humans in space: more and longer manned missions



Space Weather: effects

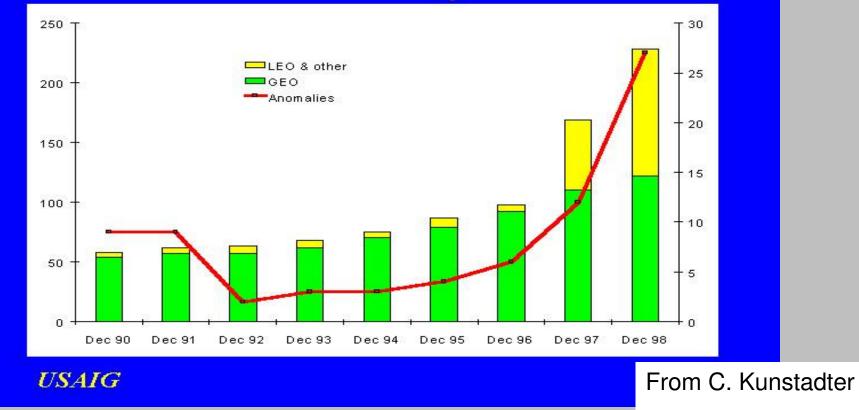


Space Weather: effects



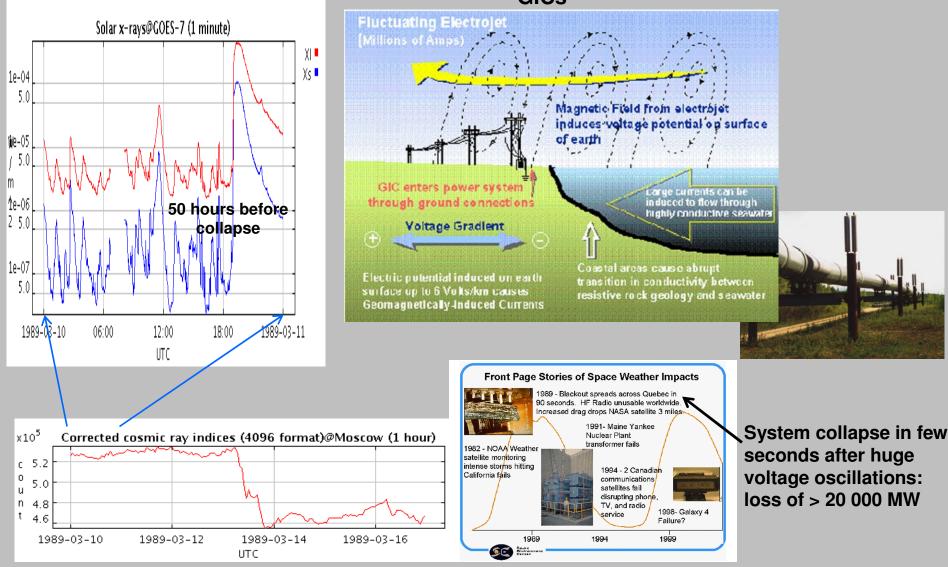
The number of reported anomalies increases with the number of satellites. But situation can be worst since many anomalies may not be reported

Satellite Anomaly Trends



Space Weather: effects

GICs



ISWI (Former IHY)

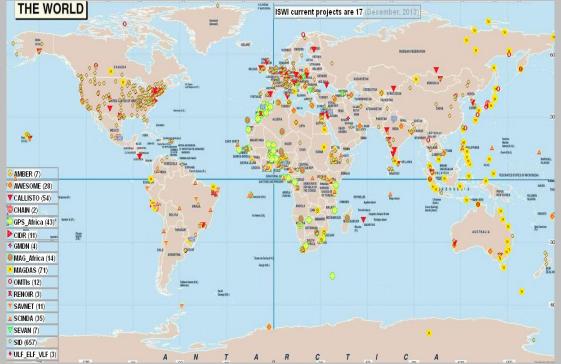
ISWI is a program of international cooperation to advance the space weather science by a combination of instrument deployment, analysis and interpretation of space weather data from the deployed instruments in conjunction with space data, and communicate the results to the public and students.

15 Instrumental Arrays

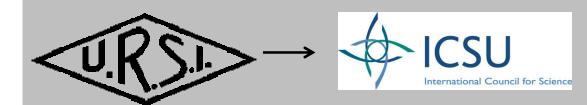
- heliospheric, solar, ionospheric, atmospheric physics
- deployed in 122 countries

11 Training schools since 2010 (ISWI)

Societal impacts workshop



During ~ 10 years IHY + ISWI \rightarrow outstanding results \rightarrow effort needs to continue





Union Radio Scientifique Internationale Past President: François Lefeuvre

Scientific Committee on Solar-Terrestrial Physics President: Nat. Gopalswamy



Coordinated at Centro de Radioastronomia e Astrofísica Mackenzie Escola de Engenharia Universidade Presbiteriana Mackenzie, São Paulo, SP, Brazil

The main objective was to incentivize the development on a long-term basis of a **regional centre** for the use of Geophysics as a new approach for **Natural Disaster prevention**, and to assess the **impacts of Space Weather conditions** on the Earth's environment and technological systems. Since radio communications and observations are important to assess geophysical hazards, the former objective implies a natural collaborative activity between Latin-American scientists involved in radio and geophysical sciences.

ICSU/URSI-SCOSTEP: Main achievements

• Foster new collaborations with scientists and institutions within (10 countries) and outside South America (7 countries) Brazil: Peru: Argentina: Mexico; Ecuador; Colombia ;

Brazil; Peru; Argentina; Mexico; Ecuador; Colombia ; Chile ; Paraguay ; Uruguay ;Venezuela

- Definition of thematic working groups WG reports
- Instrumental facilities upgrade/extension
- Development of skills and human resources





SAVNET **CRAAM/EE**





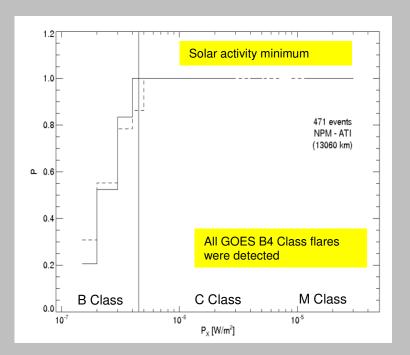
11 VLF tracking receiver stations deployed in Brazil, Peru, Argentina e México. 6 years of operation since 2007

- Long-term and transient solar activity (Ly- α ; solar flares)
- D Region absorption models
- mesospheric disturbances (T, NO, O_3)
- Physics of the lower ionospheric (C/D) regions
- Atmos. Physics (TGFs)
- Subionospheric radio propagation modeling
- Search for seismic-EM effects
- Detection of Remote astrophysical objects



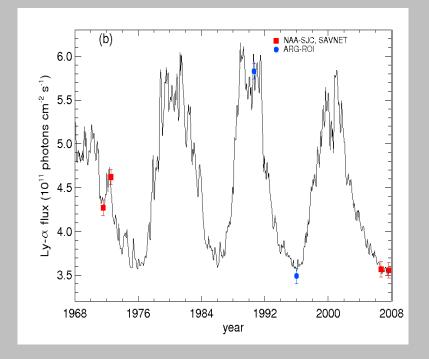
Improving Space Weather Forecasting in the Next Decade, 51st Session of the Scientific and Technical Subcommittee of the Committee on the Peaceful Uses of Outer Space (COPUOS), 10-11 February 2014, Vienna, Austria

Solar Flare Detection



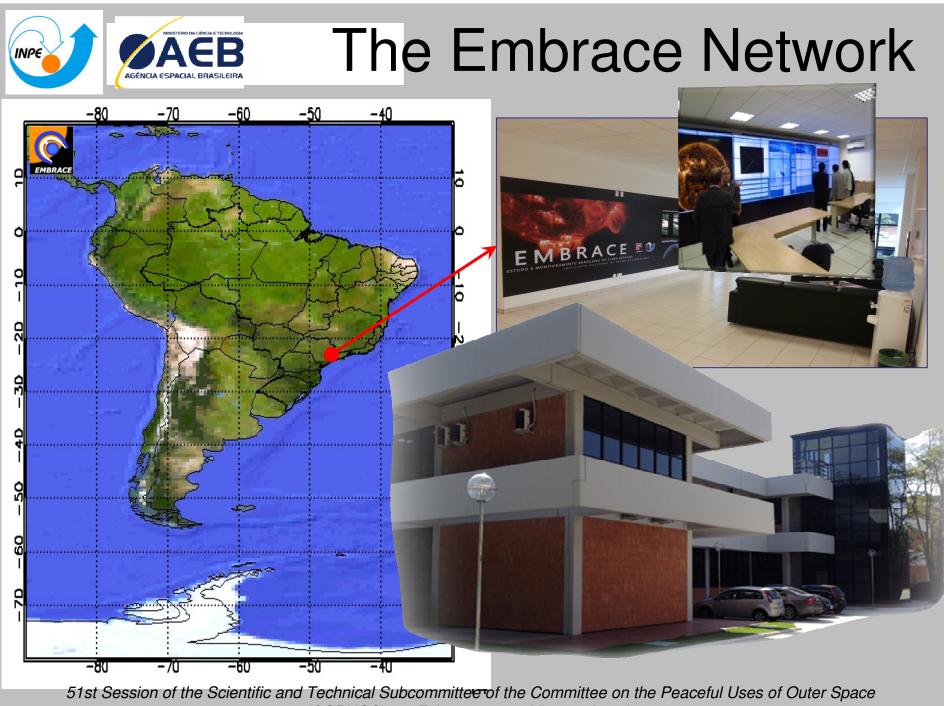
For this period of solar activity, the lower detected solar flare B 2.7 \rightarrow 2.7 10⁻⁷ W/m² \rightarrow all \geq B 4 Class events are detected with 100 % probability. Smallest SXR flare detected so far \rightarrow B1.8

Ionospheric C-Region



Indirect monitoring of the Lyman- α solar radiation

The Lyman- α line is the strongest solar radiation, important for the energetic balance in the Earth atmosphere. It cannot be observed from the ground.



(COPUOS), 12 February 2014, Vienna, Austria

The Embrace Network



Workshop de Programa de Clima Espacial do INPE com Usuários 11 de outubro de 2013 - Auditório Fernando de Mendonça - INPE



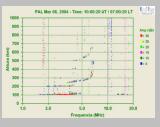


Users interested in diagnostics of Space Weather effects on different technological systems: GNSS, communication systems, geolocalization, energy distribution network, academic professionals

As a result many products are provided on real-time:

Not only but also

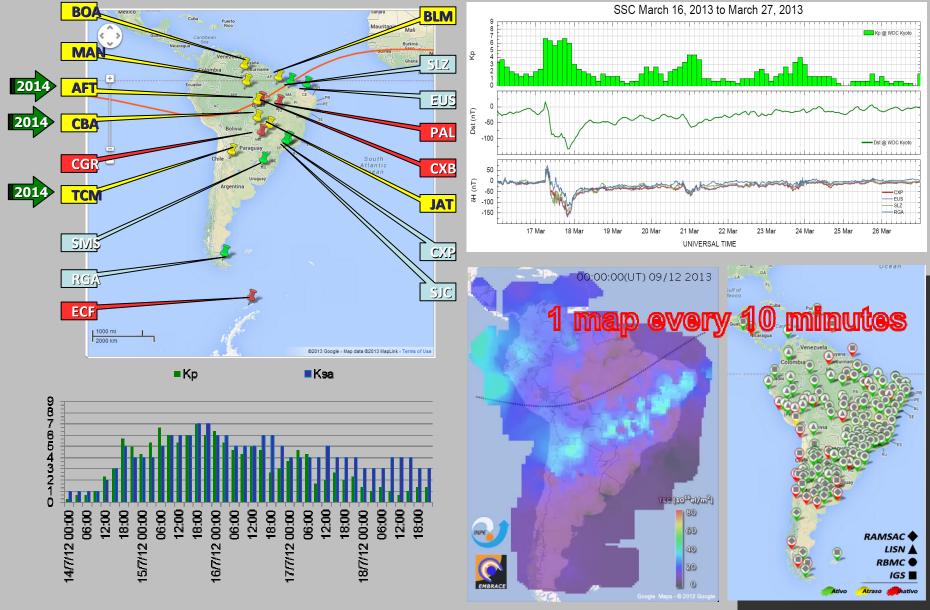
- Daily Space Weather activity bulletin
- TEC (Predicted values SUPIM)
- Ionosonde indices (f0F2, h'F, hmF2, MUF)



- Ionospheric plasma bubble activity index
- Magnetic indices

The Embrace Network

INPE



MCT / INPE / CEA / DAE - IONO - Clezio Marcos De Nardin - clezio.denardin@.inpe.br

CONCLUSIONS

Nowadays, almost everyone can feel the effects of the Space Weather dynamics \rightarrow constantly growing dependency of our society on technology (ground and space)

A better understanding of the Space Weather dynamics needs efforts on the study of fundamental processes at the Sun (solar flares, CMEs, solar cycle variations etc ...).

The lower ionosphere plasma is a medium very sensitive to external forcing: radiation, energetic particle fluxes, magnetic clouds. We can use it as a large sensor of external disturbances to monitor Space Weather effects. This is what actually does the VLF technique.

Multidisciplinary science:

Heliospheric, solar, magnetospheric, ionospheric, atmospheric physics \rightarrow

- large international collaborations
- several Unions need to be involved and to participate (URSI, SCOSTEP, COSPAR, IAGA, ...)
- put together and combine the data from the existing monitoring instrumental networks, extend the existing arrays and install new instrumental facilities (ISWI)
- modeling effort is needed for prediction of the Geo-Space Dynamics (ISWI)
- development of skills and human resources (ISWI)

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